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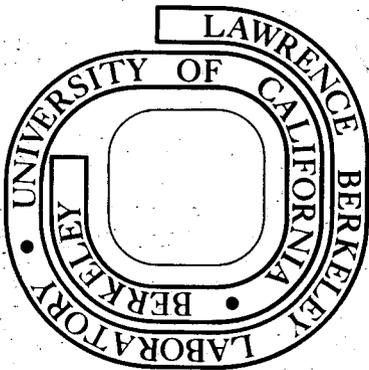
MICROSEISMS IN GEOTHERMAL EXPLORATION:  
STUDIES IN GRASS VALLEY, NEVADA

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(Ph. D. thesis)

**MASTER**

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DEDICATION

*This dissertation is dedicated to my  
wife Julie Yu-Chu for her patient  
encouragement and understanding.*

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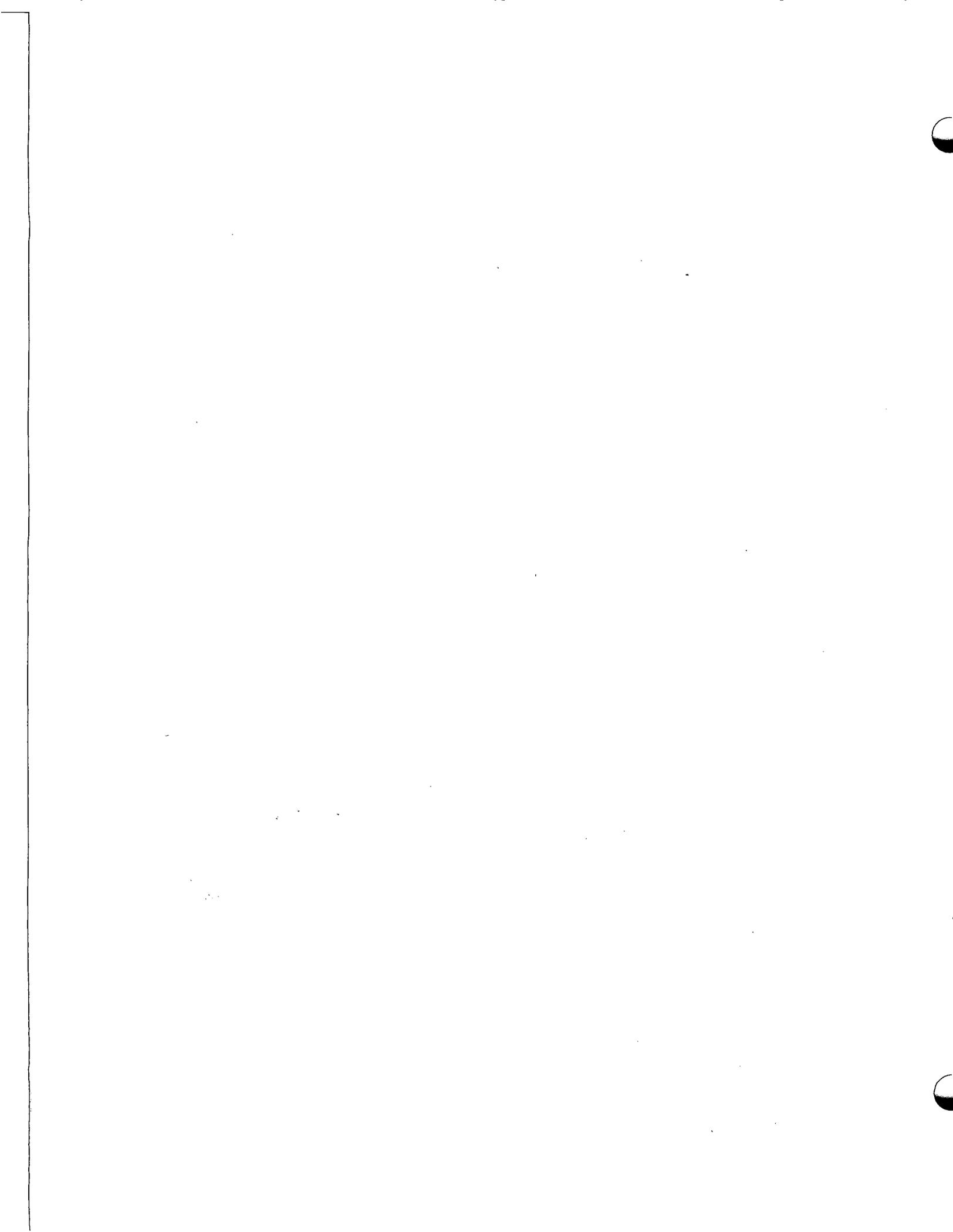


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MICROSEISMS IN GEOTHERMAL EXPLORATION:  
STUDIES IN GRASS VALLEY, NEVADA

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ABSTRACT

Frequency-wavenumber ( $f$ - $k$ ) spectra of seismic noise in the bands  $1 \leq f \leq 10$  Hz in frequency and  $|k| \leq 35.7$  cycles/km in wavenumber, measured at several places in Grass Valley, Nevada, exhibit numerous features which can be correlated with variations in surface geology and sources associated with hot spring activity. Exploration techniques for geothermal reservoirs, based upon the spatial distribution of the amplitude and frequency characteristics of short-period seismic noise, are applied and evaluated in a field program at a potential geothermal area in Grass Valley, Nevada. A detailed investigation of the spatial and temporal characteristics of the noise field was made to guide subsequent data acquisition and processing. Contour maps of normalized noise-level derived from carefully sampled data are dominated by the hot spring noise source and the generally high noise levels outlining the regions of thick alluvium. Major faults are evident when they produce a shallow lateral contrast in rock properties. Conventional seismic noise mapping techniques cannot differentiate noise anomalies due to buried seismic sources from those due to shallow geological effects. The noise radiating from a deep reservoir ought to be evident as body waves of high phase velocity with time-invariant source azimuth. A small two-dimensional array was placed at 16 locations in the region

to map propagation parameters. The f-k spectra reveal local shallow sources, but no evidence for a significant body wave component in the noise field was found. With proper data sampling, array processing provides a powerful method for mapping the horizontal component of the vector phase velocity of the noise field. This information, as well as the accurate velocity structure, will enable us to carry out seismic ray tracing and eventually to locate the source region of radiating microseisms. In Grass Valley, and probably in most areas, the 2-10 Hz microseismic field is predominantly fundamental mode Rayleigh waves controlled by the very shallow structure.

#### ACKNOWLEDGMENTS

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## I. INTRODUCTION

Two methods have been proposed for attempting to utilize microseisms for delineating geothermal reservoirs. The first is based on the speculation that hydrothermal processes deep in the reservoir radiate seismic wave energy in the frequency band 1 to 100 Hz. If this phenomenon exists, the exploration method becomes a relatively straightforward "listening" survey, using stations on a 0.5 to 2 km grid. Contours of noise power on the surface should delineate noise sources. This is the "standard" noise survey used widely in geothermal exploration. A second approach interprets the noise field as propagating elastic waves of appropriate type, e.g., fundamental mode Rayleigh waves, and inverts their propagation characteristics to obtain the distribution of medium properties, i.e. velocity and attenuation, both laterally and vertically. The propagation parameters of ambient microseisms so measured will also locate distinctive radiation sources. With sufficient knowledge of the wave nature of the microseisms and a reasonably accurate velocity-depth model, a fixed, non-aliased array can be used in a beam-steering mode to define the source region of radiating noise. Both approaches, as used in typical surveys, suffer greatly when data are contaminated by non-seismic noise, by interfering seismic wave trains, or by improper temporal and spatial data sampling. These pervasive problems have combined to render noise analysis at best a qualitative geophysical method and have substantially limited the acceptance of the seismic noise survey as an integral element in geothermal exploration.

This study attempts to avoid such problems by sophisticated analysis of microseismic data in an evaluation of the feasibility of ground noise

studies in geothermal site delineation. We report a series of investigations undertaken near Leach Hot Springs in Grass Valley, within the region of generally high heat flow in northern Nevada. We first quantify the spatial and temporal variations of ground noise in the region and find that the seismic noise spectrum is strongly affected by the near-surface geology at the recording site. In fact, with broadband seismic sensors in a mapping technique using amplitudes and frequencies, one can outline lateral variations in alluvial thickness. This "standard" mapping technique cannot differentiate noise enhancement due to shallow structure from noise enhancement due to buried seismic source. On the other hand, we find that the mapping of wave propagation parameters provides additional information about the noise field. However, the successful application of this technique requires some understanding of the wave nature of microseisms. We used multiple-sensor arrays to study the seismic coherency as a function of frequency and spatial separation. Based on this information, we designed an array to collect propagating microseismic data. The array data were processed by both the frequency domain beam-forming method (BFM) and the maximum-likelihood method (MLM). From the dispersion curves obtained in our array study, we verify that the seismic noise consists primarily of fundamental mode Rayleigh waves.

In Chapter 2, I consider the geothermal ground noise within the content of wide-band microseisms and discuss the origin of seismic noise in three different period ranges. In Chapter 3, I describe the area of field study and the systems used for field data acquisition. The conventional seismic noise power contour technique is evaluated

in Chapter 4. Finding weakness in the conventional mapping method, I present a sophisticated array mapping technique in Chapter 5, along with field procedures, data processing schemes, and the results of the Grass Valley study. In the concluding chapter, I summarize the results of the study and present recommendations for extensions of the work.

## II. EARTH NOISE

### 2.1 Introduction

The study of microseisms, or earth noise, involves interplay of meteorology, oceanography and seismology. The investigation of microseisms has attracted seismologists since the beginning of this century. A bibliography covering work up to 1955 (Gutenberg and Andrews, 1956) lists over 600 articles on the subject. In a recent review, Båth (1974) provided an additional 69 articles related to spectral studies of microseisms.

A typical seismic noise spectrum recorded at a quiet site (Figure 2.1) shows:

- 1) The amplitude increases with period for periods longer than 30 sec. Over this band seismic noise is due to atmospheric loading (Savino, et al., 1972).
- 2) Two distinctive maxima appear around 6-8 sec and 14-20 sec. These are related to ocean waves.
- 3) Low amplitude short-period noise ( $T < 0.8$  sec) associated with human activity and local natural disturbances, including weather.

A distinct worldwide minimum in seismic noise occurs around 30 to 40 sec, separating the very long-period microseisms of atmospheric origin from those long-period waves due to ocean swell. (Savino, et al., 1972; Whorf, 1972). Another prominent minimum around 0.8 sec separates short-period local noise from long-period ocean swell (Figures 2.1 and 2.2).

The noise spectrum,  $S(f)$  can be expressed as

$$S(f) = I(f) E(f) G(f) \quad (2.1)$$

where:

$I(f)$  = the source spectrum,

$E(f)$  = the transfer function of entire path, excluding near surface effects at the sensor, and

$G(f)$  = the transfer function of the shallow section at the recording site.

The noise spectrum reflects the nature of the source spectrum, and the acoustic properties of the medium both along the path and beneath the recording site. At a quiet bedrock location, away from strong local surface or subsurface sources, the noise spectrum generally exhibits a frequency dependence of  $f^{-1}$  to  $f^{-3}$ , with local reversals at the peaks (Figures 2.1 and 2.2). In Figure 2.1 the power spectral density obtained by Fix (1972) corresponds to the minimum value recorded at a quiet site (Brune and Oliver, 1959) and to the typical background noise spectrum in 10 to 100 sec range at Albuquerque (Peterson et al., 1976). The results of Peterson et al. (1976) were recorded by the broad-band bore-hole seismometers of the Seismic Research Observatories.

In this chapter, I discuss the sources of seismic noise, emphasizing those observations of particular relevance for geothermal ground noise. In the last section, I discuss the effect of local geology on seismic noise.

## 2.2 Short-period seismic noise ( $T < 0.8$ sec)

Microseisms in the period range 0.8 sec to 0.01 sec (1.25 Hz to 100 Hz) are generated by human activity such as traffic or machinery (usually called "cultural" noise), by natural surface disturbances due

to wind, or by subsurface activity such as microearthquakes or geothermal processes. Over this frequency range the seismic energy attenuates rapidly with increasing distance from the source. As indicated by Douze (1967), short-period surface seismic noise consists mostly of fundamental mode Rayleigh waves. He also found that at depths of 2 to 3 km beneath the earth's surface, where the fundamental Rayleigh mode has decreased to negligible amplitude, the remainder of the noise consists of random body waves.

Cultural noise. Noise sources related to human activity are responsible for diurnal variations of background noise level as seen in the Waiotapu region of New Zealand (Whiteford, 1970), and at East Mesa, California (Iyer, 1974), both geothermal regions. At East Mesa the freeway, with frequent truck traffic, produces wide-band (up to 10 Hz) noise as well as large amplitude variations which can be detected at distances up to 8 km. Measurements near roads in the Waiotapu region show that traffic can produce ground motion up to  $700 \times 10^{-9}$  m/sec about 13 dB above a quiet site, in 0.5 - 100 Hz, at a distance of 0.5 km (Whiteford, 1970).

Rivers, canals, and waterfalls. Iyer and Hitchcock (1976) reported that the river flowing in Long Valley, California, generates noise at frequencies above 6 Hz and that the noise is attenuated by 12 dB 1 km from the river. In East Mesa, irrigation canals seem to be continuous wide-band sources of noise attenuating rapidly with distance, reaching a fairly steady level beyond 3 km. At the power drops (small waterfalls) along the canals, however, a distinctive noise component is generated in a narrow frequency band around 2.5 Hz (Iyer, 1974). Iyer and Hitchcock

(1974) found that in Yellowstone National Park the waterfalls generate a narrow band seismic noise at 2 Hz which is clearly different in appearance from the noise generated by the hydrothermal process.

Wind. Although both Whiteford (1970) and Iyer (1974) reported poor correlation between the amplitude of short-period seismic noise with wind speed, Frantti (1963) reported that high wind influenced short-period noise. His results are presented in Figure 2.3 for three components of motion. The wind effect influences a broad frequency range, particularly on the horizontal component oriented parallel to the wind direction. Peterson et al. (1976) have shown that the effect of wind-generated noise decreases as the depth of sensor location increases.

Geysers, hot springs, and fumaroles. Reports relating short-period microseisms to geothermal processes are numerous. We will discuss this topic in more detail in Section 2.5.

### 2.3 Long-period seismic noise (2 sec < T < 20 sec)

Early studies of microseisms concentrated on the observation and the origin of 6 - 8 sec microseisms which have been reported as a world-wide phenomenon. The amplitude of microseisms varies from site to site and it varies temporally at a given site. The term "microseismic storm" is used to refer to the occasional periods of unusually large amplitude microseismic activity. Since the dominant microseism period at 6 - 8 sec corresponds to half the period of a pronounced peak in the ocean wave spectrum, and since large amplitude microseisms are always observed near the coast, these microseisms are inferred to be of

oceanic origin. Longuet-Higgins (1950) has shown theoretically that the 6 - 8 sec microseisms are generated by the standing wave phenomenon in ocean wave motion. The standing waves are caused by the interference of groups of waves of the same wavelength, but not necessarily of equal amplitude, travelling in opposite directions. The mean pressure on the bottom beneath a train of standing waves is not constant, as in a progressive wave, but fluctuates with an amplitude independent of the depth and proportional to the square of the wave height. The oscillation of standing waves produces the type of energy required for the generation of ground movement. This type of energy is unattenuated by water depth, and in phase at all points of the bottom, suitable for producing coherent seismic waves. Further, the frequency of this pressure variation is twice the fundamental frequency of the waves. The favorable environments in which the standing wave motion may be strong are:

- 1) the wake of a moving storm,
- 2) the center of a storm, and
- 3) reflection from a coast.

The generation of 6-8 sec microseisms by ocean storms has been reported (Iyer, 1958; Toksöz and Lacoss, 1968; Haubrich and McCamy, 1969; and Vinnik, 1971), suggesting the first or second possible source of generation. Other observations related to ocean waves near the coastline indicate the third possible source (Haubrich et al., 1963).

The double-peaked spectral feature characterized by two distinctive maxima at 6-8 sec and 14-20 sec respectively has attracted attention since the development of high gain long period seismographs. Microseisms with the same dominant period as ocean surface waves are referred to

as primary frequency microseisms (PF), as distinct from double frequency microseisms (DF) or secondary microseisms, which have half this period. The microseisms of the double frequency contain more energy than those of the primary frequency. The primary and secondary microseisms are generated from the same meteorological disturbances at the east and west coasts of the United States (Oliver and Page, 1963; Haubrich et al., 1963). The primary microseisms are probably generated as a result of the incidence of gravity waves in water upon a coastline; this is in general agreement with the classical "surf" theory. The generative strip is presumably confined to shallow water, 100 miles up or down the coast (Haubrich et al., 1963). The secondary microseisms are apparently due to standing water wave oscillation, in agreement with Longuet-Higgins' theory, in shallow and/or deep water.

#### 2.4 Very long period seismic noise (T > 30 sec)

The power spectral density of very long period seismic noise in 30 to 130 sec range rises smoothly with period, as shown in Figure 2.1 for the vertical component. Haubrich and Mackenzie (1965) and Capon (1969) have observed that noise in this period range consists of non-propagating energy, as indicated by low coherency between two seismometers separated a few km, plus some fundamental Rayleigh mode energy. The non-propagating seismic noise is apparently generated by atmospheric pressure variations. Sorrells et al. (1971) have noted that during windy intervals there is strong correlation between local atmospheric pressure changes and the noise recorded by a vertical seismograph located on the surface. In contrast, over the same range of periods,

there is no correlation between seismic noise recorded in a mine at a depth of 183 m and local atmospheric pressure changes, except during the passage of acoustic gravity waves. The model for generating earth noise in this period range is related to atmospheric pressure disturbances that are relatively coherent over a fraction of a wavelength and propagate with jet stream velocities (30-100 m/sec) which are much slower than seismic waves (Savino, et al., 1972). An idealization of this model corresponds to a static, atmospheric pressure loading consisting of random pressure disturbances acting over equal areas on the earth surface. This idealized model explains the level and shape of the spectrum of very long-period earth noise and the attenuation of microseismic ground motion with depth.

#### 2.5 Geothermal ground noise

New Zealand. Clacy (1968) first suggested that seismic noise increased near geothermal reservoirs. His first results northeast of Lake Taupo, New Zealand, were based on contours of total noise amplitude in the frequency band 1 to 20 Hz. In subsequent surveys at Wairakei, Waiotapu, and Broadlands geothermal areas, he found that the local noise amplitude anomalies were characterized by a dominant frequency of 2 Hz, whereas, away from the area of the anomaly, frequencies higher than 3 Hz predominated. On the other hand, Whiteford (1970) found in repeat surveys of the same areas that neither the shape of the frequency spectrum nor its dominant frequency conformed to any regional pattern. Whiteford measured the absolute ground motion in the Waiotapu geothermal area and found that, within a distance of 1 to 2 km of the high heat

flow area, the average minimum ground particle velocity was greater than  $150 \times 10^{-9}$  m/sec, while further away, the amplitude of the ground movement decreased by a factor of about 3 and, in addition, exhibited pronounced diurnal variations.

Imperial Valley, California. In the United States, a similar survey was first carried out southeast of the Salton Sea by Goforth et al. (1972) who suggested an empirical relationship for geothermal reservoirs between high temperature gradient and high seismic noise level. Their results showed a significant increase in the noise power in the frequency band of 1 to 3 Hz at sites above the reservoir. They estimated the power spectrum at each site with ten 200-second data segments taken over eight hours of night-time recording. The contour map of the total power in the frequency band of 1 to 3 Hz was similar to the temperature gradient contour map. Douze and Sorrells (1972) conducted a similar survey over the nearby East Mesa area, where they found that the total seismic power in the 3 to 5 Hz band exhibited spatial variations similar, in general, to the gravity and heat flow fields. East Mesa was later surveyed by Iyer (1974), with significantly different results. Iyer measured seismic noise by averaging 20 of the lowest values of the RMS amplitude in several narrow frequency bands, using data blocks of 81.92 sec selected from four hours of digital data. He did not find an anomaly in seismic noise associated with geothermal activity but only the noise from canals and freeway traffic.

Yellowstone National Park. The seismic pulsation associated with several geysers in Yellowstone National Park is believed to be indicative of

the heating of water in the underground reservoir and the eruption triggered by the superheated system. Nicholls and Rinehart (1967) have studied the seismic signature of several geysers in the park and inferred that their predominant pulse frequencies are quite similar, in the range of 20-60 Hz, presumably due to steam action. The very low frequency seismic pulses recorded at Old Faithful, Castle, Bead, Plume and Jewel geysers are believed to be associated with some type of water movement. The maximum amplitude of seismic pulses recorded in the park is  $5.08 \times 10^{-5}$  m/sec. At Old Faithful Geyser the maximum amplitude is  $2.54 \times 10^{-5}$  m/sec at 30-50 Hz.

Iyer and Hitchcock (1974) have also found good correlation between geothermal activity and high seismic noise levels in 1 to 26 Hz in the park. The ground noise level in non-geothermal areas of the park is approximately  $13$  to  $15 \times 10^{-9}$  m/sec at 1 to 26 Hz. In the Lower and Upper Geyser Basins where there are numerous geysers and hot springs, the average noise level is in general higher than  $50 \times 10^{-9}$  m/sec, and reaches a value of  $672 \times 10^{-9}$  m/sec near the Old Faithful Geyser. In the Norris Basin, another highly active geyser basin in the park, the noise level varies from 50 to  $500 \times 10^{-9}$  m/sec. Part of the observed noise in the Lower, Upper, and Norris Geyser Basins is no doubt generated by the hydrothermal activity at the surface. The measurements near Old Faithful indicate that high-frequency noise, in the 8 to 16 Hz band, is generated during the geyser eruption; the noise level of lower frequencies is not affected by the eruption cycles. Noise levels around Mammoth Hot Springs are two to five times higher than in the surrounding area. There is no geyser or fumarole

here, and the geothermal water is relatively cooler than at Norris and the other geyser basins. Hence, it is very unlikely that the seismic noise observed here is generated near the surface. The noise anomaly observed in the area between Lower Falls and Mud Volcano could be caused by ground amplification effects in the soft sedimentary deposits.

Other Areas. Correlations have also been reported between geothermal activity and high seismic ground noise in the Vulcano Islands, Italy (Luongo and Rapolla, 1973), the Coso geothermal area, China Lake, California (Combs and Rotstein, 1975), and Long Valley, California (Iyer and Hitchcock, 1976). High frequency noise,  $f > 8$  Hz, in the vicinity of geysers, fumaroles, and hot springs is associated with hydrothermal activity near the surface and during the geyser eruption. Low frequency noise,  $f < 8$  Hz, is not affected by geyser eruption cycles and is probably generated at depth. Other than those active sources the noise power anomaly may also result from lateral variation in near-surface velocity, particularly where low velocity alluvium is involved.

## 2.6 Local geological structure effects

It has been noted that the seismic waves observed from earthquakes and explosions are strongly affected by the near-surface geology at the recording site. The amplitude of seismic waves is generally smaller at a bedrock site than at an alluvium site. Low-velocity surface materials tend to amplify incident body waves (Borcherdt, 1970; Murphy et al., 1972) and surface waves (Lysmer and Drake, 1972) in selective frequency bands. The amplification effects of body waves result from resonance phenomena caused by large reflection coefficients

related to velocity and thickness of the overburden. In the case of the surface waves, the shallow structure can provide a waveguide at particular frequencies corresponding to the maxima and minima in the group velocity dispersion curve. Similar effects have been reported for seismic noise, where amplitude and dominant frequency are characterized by the geology of the recording site (Kanai and Tanaka, 1961; Borchardt, 1970; Iyer and Hitchcock, 1976). In general, the seismic noise at a quiet bedrock site exhibits a smooth spectrum and small amplitude, whereas on a surface of deep weathering or thick sedimentary overburden, the noise spectrum shows a large peak in a particular frequency band.

For a normally incident-plane SH wave at the lower boundary of an elastic layer over an elastic half-space, the theoretical resonance frequencies are

$$f = \frac{(2n-1)\beta}{4h} \quad \text{for } n = 1, 2, 3, \dots, N,$$

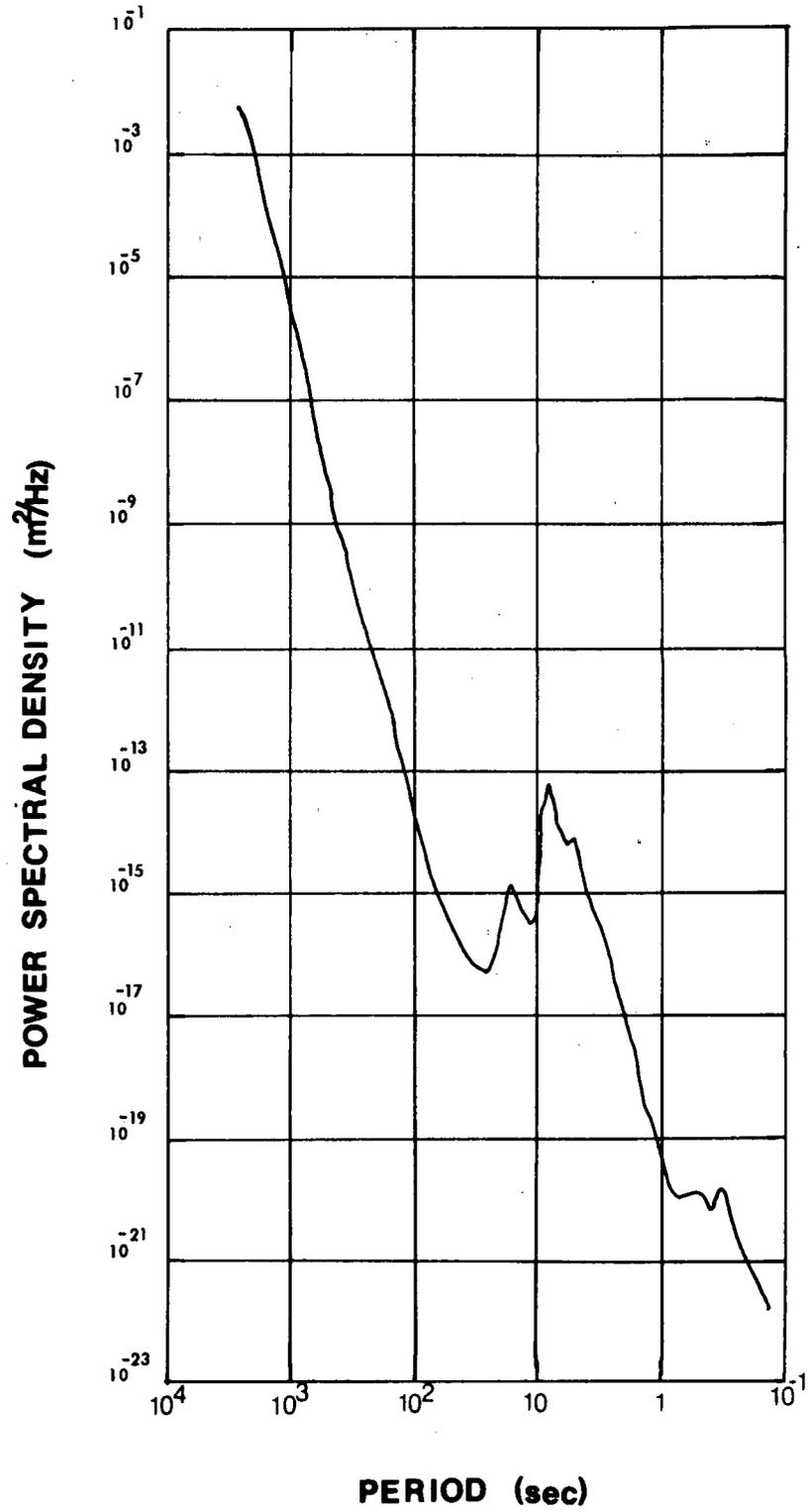
where  $\beta$  = shear-wave velocity of the layer and

$h$  = thickness of the layer.

This relation also holds for the case of P-wave, using the compressional wave velocity. Consequently, the spectrum of recorded motion will show distinctive peaks at the resonance frequencies. In practice, however, complex geological structure will complicate the nature of the spectrum (Kanai and Tanaka, 1961; Kanai, et al., 1966).

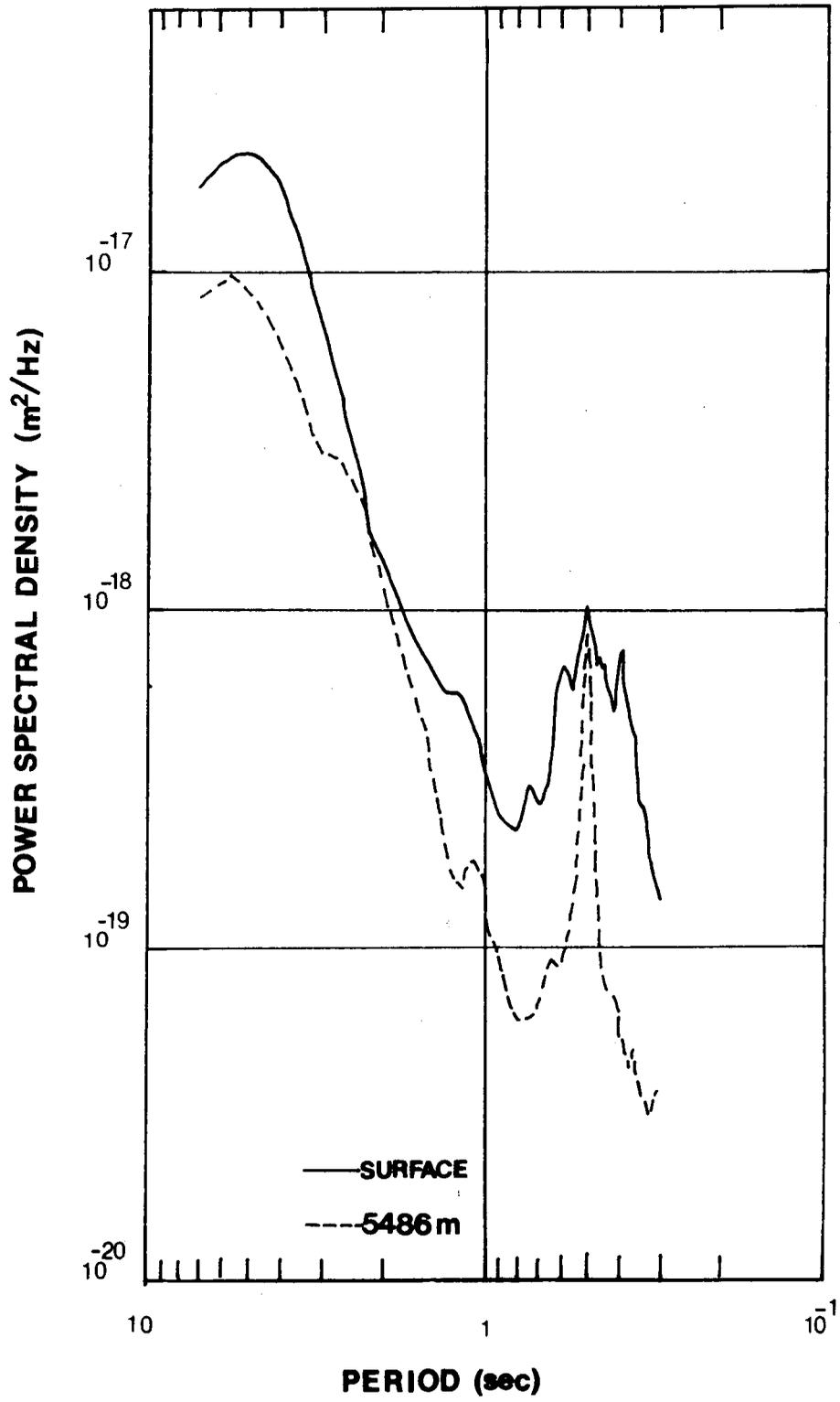
FIGURE CAPTIONS

- Figure 2.1 Power spectral density of the vertical component earth noise recorded at the Queen Creek Seismological station, Arizona (after Fix, 1972). This spectrum is typical of a quiet site. The distinctive humps at 8 and 15 sec arise from ocean sources.
- Figure 2.2 Power spectral density of the vertical component short-period earth noise recorded at the surface and at a depth of 5486 m in Grapevine, Texas (after Douze, 1967). Note the attenuation with depth.
- Figure 2.3 Earth noise spectra showing the effect of high wind (>20 mph) and low wind (<5 mph) in three components of short-period particle motion at Rural Valley, Pennsylvania (after Frantti, 1963).



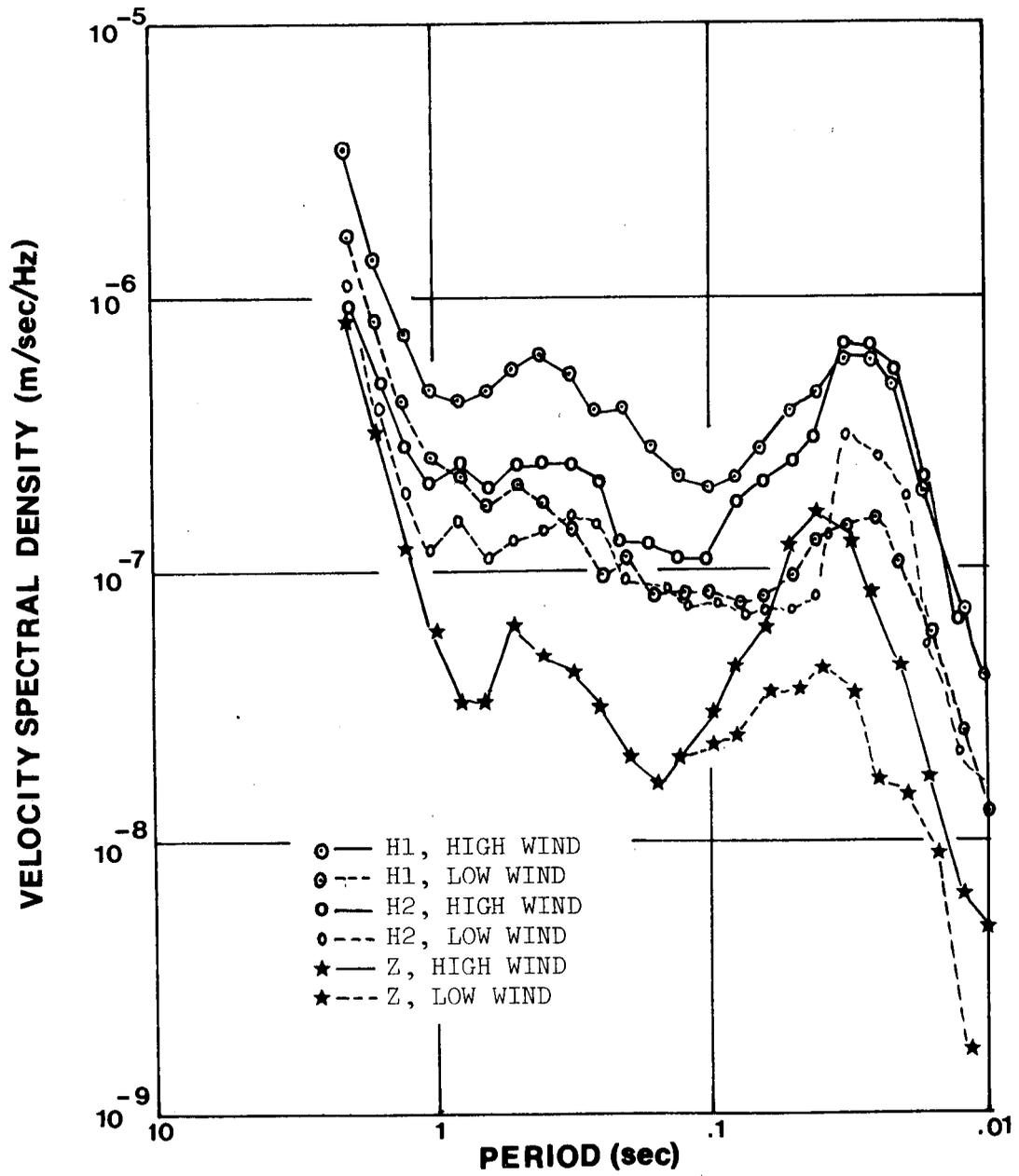
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Figure 2.1



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Figure 2.2



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Figure 2.3

### III. FIELD PROGRAMS

#### 3.1 Introduction

In this chapter I discuss the data collection phase of the seismic ground noise study. The location, geological setting and tectonic history of the potential geothermal resource region are presented. Data acquisition is discussed in the concluding sections.

#### 3.2 Area of Study

Leach Hot Springs, in Grass Valley, Nevada, is located 30 km south of Winnemucca (Figure 3.1). Grass Valley is a typical valley of the Basin and Range province with normal faulting, major earthquakes, and hot springs occurring along the valley margins. The valley is bounded by the Sonoma and Tobin Ranges to the east, and the basalt-capped East Range to the west. The valley narrows south of the hot springs as it approaches the Goldbanks Hills (Figure 3.2). These ranges are composed of Paleozoic sedimentary rocks, or Triassic siliceous clastic and carbonate rocks. Some granitic intrusions, probably of Triassic age, are found in the Goldbanks Hills; elsewhere the granites are probably of Cretaceous age. The valley is filled with Quaternary alluvial sediments. There are sparse occurrences of Tertiary rhyolite and basalt, and in the vicinity of Leach Hot Springs there are Quaternary sinter deposits. The distribution of major lithologic units in the region is illustrated on the geologic map (Figure 3.3).

As is characteristic of hot spring systems found in northern Nevada, Leach Hot Springs is located on a fault, strongly expressed by a 10 to 15 m high scarp trending NE. Normal faulting since mid-Tertiary

has offset rock units vertically several tens to several hundreds of meters. As shown on the fault and lineament map (Figure 3.2), the present day hot springs occur at the intersection of a major NE-trending fault and the more common NNW-SSE trending lineament on the eastern side of the valley.

Leach Hot Springs is within the high heat flow area of northern Nevada indicated in Figure 3.1. This high heat flow area is often called the "Battle Mountain high" (Sass *et al.*, 1971) and exhibits heat flow values in the range of 1.5 to 3.5 HFU (1 HFU = 10 cal/m<sup>2</sup> sec). The diffuse region of elevated heat flow over the Basin and Range province is generally thought to be an expression of high temperature in the lower crust and upper mantle, and it seems reasonable to interpret the localized Battle Mountain high as an effect of fairly recent intrusion of magma into the earth's crust. Quaternary volcanism within the province supports this hypothesis.

### 3.3 Other geophysical data in this area

Geophysical data were obtained primarily along 17 survey lines (Figure 3.4), although not all methods were employed on every line. Line E is typical. Bouguer gravity anomaly, P-wave delay data, and seismic reflection data, presented in Figure 3.5 for line E, indicate that the greatest thickness of sediments and major faulting occur near the eastern valley margin. The major lithologic units from the seismic reflection section are Quaternary alluvium (1.8 km/sec), Tertiary sedimentary and volcanic rocks (2.9 km/sec), Paleozoic rocks (4.0 km/sec), and deep basement (5.0 km/sec), respectively. The basement surface

rises gently to the west, but is apparently up-faulted at the eastern boundary faults as indicated by the Bouguer gravity map (Figure 3.6).

The low apparent resistivity zone beneath E2W-E4W (Beyer, et al., 1976), found in the dipole-dipole resistivity survey, has been identified with Tertiary sediments. Since the heat flow value in this zone is not high by Battle Mountain standards (2.24 HFU), the accumulation of conductive sediments, such as ancient playa deposits in the deepest portion of the valley, is probably responsible for the resistivity anomaly.

The only portion of the Grass Valley area that is seismically active is in Panther Canyon and in the valley immediately west of it. This seismic zone is one of complicated faulting and frequent micro-earthquakes. The area is dominated by a strong NE trending gravity feature which offsets topography and the Bouguer gravity anomaly (Figure 3.6). There is a strong electrical conductivity high and a high heat flow of 4.9 HFU. More details of the geophysical data obtained in the Grass Valley area are given by Beyer, et al., (1976).

#### 3.4 Seismic data acquisition system

A portable seismic network, with up to 12 stations linked by radio telemetry to a recording system mounted in a small two-wheeled trailer, was designed for simplicity, flexibility and ease of installation. It proved possible for two men to deploy the sensors and check out the telemetry in about one day. Ease of network emplacement made it possible to modify the array as data were collected and to design field experiments with multiple objectives.

A 4.5 Hz vertical-component geophone, a high-gain amplifier (60-120 dB), a voltage controlled oscillator, and a radio transmitter constitute the station site equipment. The block diagram of one typical station is shown in Figure 3.7. A 0.1 watt transmitter gives a range of about 20 km for average topography. In applications using all 12 geophones spaced over a small aperture array (50m), the radio links were eliminated and signals were transmitted by wire to the recording trailer. In the trailer are housed the radio receivers, FM discriminators, a 14-channel slow-speed FM tape recorder (0.12 ips, 0-40 Hz, 8 days; or 0.24 ips, 0-80 Hz, 4 days), the timing system, and batteries. A slow-speed smoked-paper recorder was used as a monitor. Figure 3.8 gives the details of the central recording system housed in the trailer. The system has about 40 dB dynamic range (peak-to-peak measurement), limited primarily by the tape recorder.

Data are played back at the Seismographic Station on the Berkeley campus and selectively digitized at a rate of 40 samples per second. The transfer function (ground particle velocity to volts from playback discriminator) of the recording and playback system before digitization is expressed as the product over the transfer function of all system elements

$$H(f) = H_g(f) \cdot |H_A \cdot H_{RC}(f) \cdot \frac{4}{VCOFS} \cdot H_p(f)| \text{ volts/m/sec.} \quad (3.1)$$

The phase term of  $H_g(f)$  is different in some channels. Other terms are recording/playback/digitization systems which are the same for all channels, so no phase term is involved.

$$H_g(f) = G \cdot \frac{R_D}{R_S + R_D} \cdot \frac{f^2}{[(f_S^2 - f^2)^2 + (2\zeta f_S f)^2]^{1/2}} \exp[j(\theta - \frac{\pi}{2})], \quad (3.2)$$

where  $G$  = generator constant of geophone = 77 volt/m/sec for the GEOSPACE GSC-8D and HS-1 geophones,

$R_D$  = damping resistance in ohms, typically 2000 to 3000 ohms,

$R_S$  = coil resistance of geophone = 3400 ohms or 1310 ohms,

$f_S$  = natural frequency of geophone, nominally 4.5 Hz,

$\zeta$  = damping factor,

$$\theta = \tan^{-1} \frac{2\zeta f f_S}{f^2 - f_S^2} \quad (3.3)$$

The exponential term in Equation (3.2) is the phase difference introduced by the geophone. In the small-aperture array data acquisition mode, we have used geophones of two different damping factors, i.e.  $\zeta = 0.3$  and  $\zeta = 0.6$ . The phase correction is, therefore, necessary to correct those data used for f-k analysis. The geophone phase is the only phase correction term required in the data processing, since we have used matched phase responses to every other system element. We define the positive phase term  $\exp[j\theta(f)]$  to be phase delay for  $\theta > 0$ . The characteristics of geophones used in this study are listed in Table 1.

The modulus of transfer functions of the two-pole RC low-pass filter and the high-pass filter in the SPRENGNETHR AS-110 amplifier with mid-band gain,  $H_A$ , is

$$|H_A \cdot H_{RC}(f)| = |H_A| \cdot \frac{1}{1 + (\frac{f}{f_H})^2} \cdot \frac{1}{1 + (\frac{f}{f_L})^2} \quad (3.4)$$

where  $f_H$  and  $f_L$  are cut-off frequencies of the filters. "VCOFS" in Equation (3.1) represents the maximum voltage of the amplified signal set to correspond to +250 Hz of full-scale FM modulation about the

TABLE 1

Damping Characteristics of Geophones

Geophone identification	Damping resistance $R_D$ (k $\Omega$ )	Coil resistance $R_S$ (k $\Omega$ )	Damping factor
3	3.3	1.31	0.6
4	1.5	1.31	0.6
5	3.0	1.31	0.6
6	3.0	1.31	0.6
7	4.3	1.31	0.6
9	5.0	1.31	0.6
11	2.0	1.31	0.6
12	3.0	1.31	0.6
13	3.5	1.31	0.6
14	3.0	1.31	0.6
16	3.0	1.31	0.6
17	3.5	1.31	0.6
18	3.0	1.31	0.6
B1	infinite	3.4	0.3
B2	infinite	3.4	0.3
B3	infinite	3.4	0.3
B4	infinite	3.4	0.3
B5	infinite	3.4	0.3
BY1	infinite	3.4	0.3
BY2	infinite	3.4	0.3

VCO (voltage controlled oscillator) frequency. The tapes playback discriminators contain 4-pole Butterworth filters with

$$|H_p(f)| = \frac{1}{[1 + (\frac{f}{f_p})^8]^{1/2}} \quad , \quad (3.5)$$

where  $f_p$  is the tape recorder bandwidth (e.g., 40 or 80 Hz, depending on the tape recorder used). Since full-scale voltage from the tape playback discriminator is +4V, and the digitizer full-scale is +2048 counts, the transfer function to the digitized signal is

$$H_D(f) = H(f) \cdot \frac{2048}{4} \cdot |H_{BW}(f)| \text{ counts/m/sec.} \quad (3.6)$$

where  $H_{BW}(f)$  is the anti-alias filter (3 four-pole Butterworth filters) at low-pass corner  $f_{BW}$ ;

$$|H_{BW}(f)| = \frac{1}{[1 + (\frac{f}{f_{BW}})^8]^{3/2}} \quad . \quad (3.7)$$

The modulus of transfer function of the system in counts/millimicron/second is shown in Figure 3.9, assuming  $R_D / (R_S + R_D)$  to be unity. This is the only correction required for VSD estimation, since the phase information is not essential.

FIGURE CAPTIONS

- Figure 3.1 Prominent thermal springs and the Battle Mountain high heat flow region in northwestern Nevada. Leach Hot Springs is located in the center of the high heat flow area.
- Figure 3.2 Mapped faults and pertinent geophysical traverses in the Leach Hot Springs area. Hachured lines indicate down-faulted sides of scarplets; ball symbols indicate downthrown side of other faults. Star shows location of Leach Hot Springs. Heavy solid lines are survey lines E, B, and G with tick marks every 1 km. AC, A2N, and GP are observation sites discussed in text.
- Figure 3.3 Lithologic map, Leach Hot Springs area, Qal: Alluvium, Qos: older sinter deposits, Qsg: sinter gravels, Qtg: Quaternary-Tertiary gravels and fanglomerates, Tb: Tertiary basalt, Tr: Tertiary rhyolite, Tt: tuff, Ts: Tertiary sedimentary rocks, Kqm: quartz monzonite, Kg: granitic rock, md: mafic dike, TRg: Triassic granitic rocks, TR: Undifferentiated Traissic sedimentary rocks, P: undifferentiated Paleozoic sedimentary rocks.
- Figure 3.4 Geophysical survey lines in Grass Valley, Nevada.
- Figure 3.5 Profiles on line E of Bouguer gravity anomaly, P-wave delay, migrated seismic reflection section, and the instantaneous microseismic field, showing east margin fault (trace at 1E) and the maximum sediment thickness around 2W. Averaged section compressional velocities shown: (A) 1.8 km/sec Quaternary alluvium,

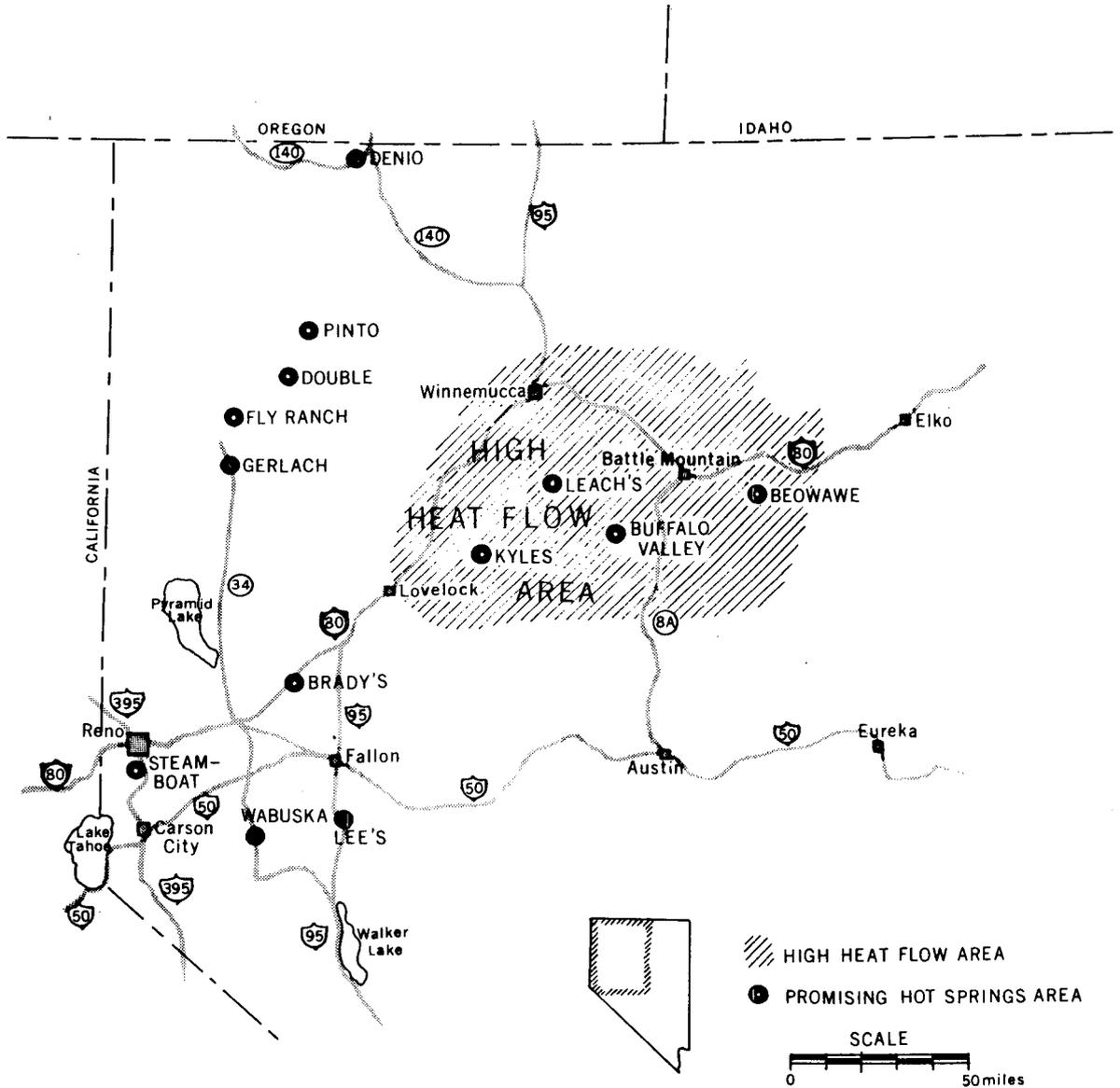
(B) 2.9 km/sec Tertiary sediments, and (C) 4.0 km/sec Paleozoic rocks. A detailed explanation of the microseismic field contour is given in Figure 4.6.

Figure 3.6 Bouguer gravity anomaly map of Grass Valley, Nevada. Station locations are shown as dots on the map. The gravity low axis along the eastern side of the valley corresponds to the greatest thickness of sediment.

Figure 3.7 Block diagram of a typical seismic station.

Figure 3.8 Block diagram of the central receiving and recording system. The splitter box at the center of the figure indicates the device distributing a multiplexed signal to two or more discriminators.

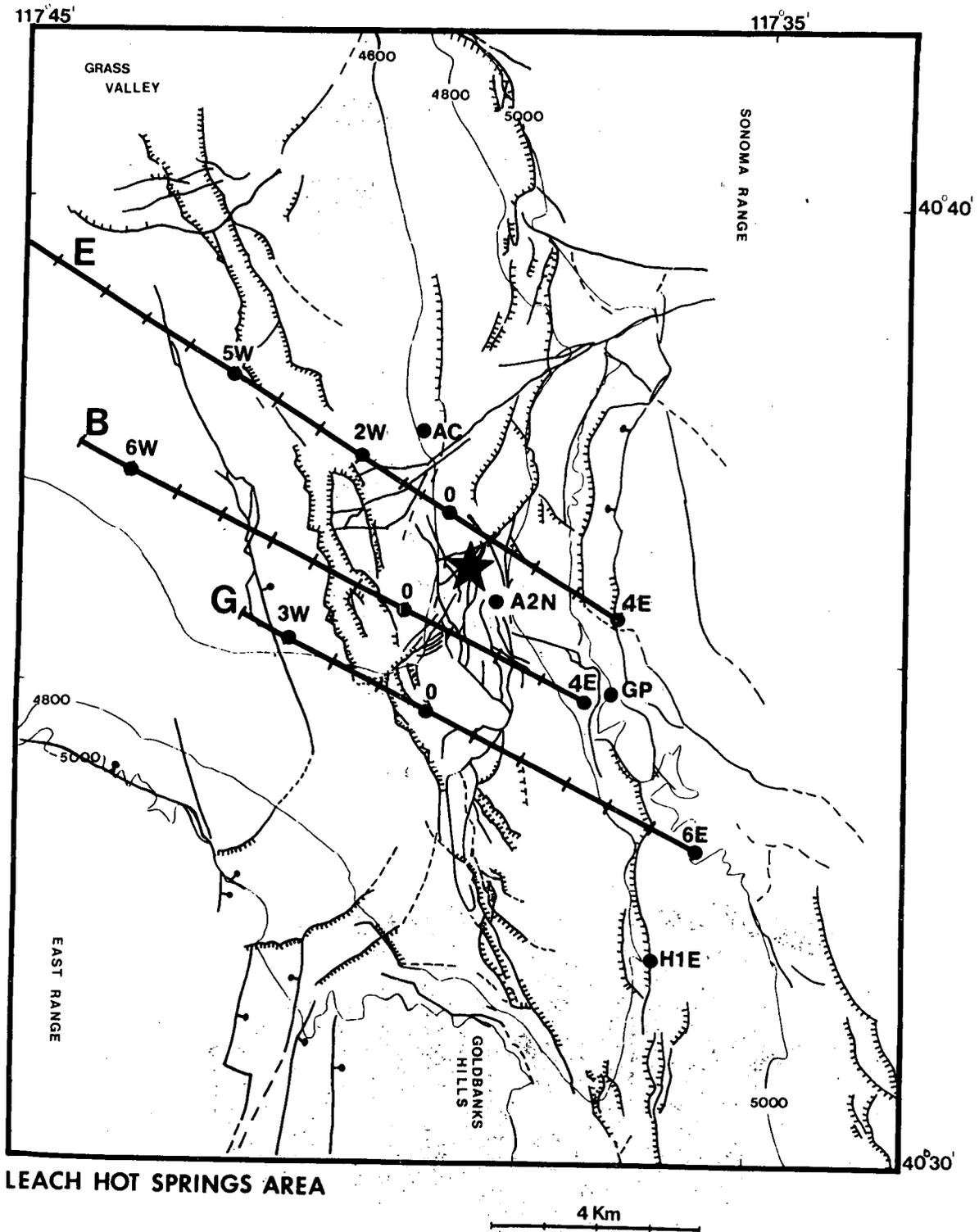
Figure 3.9 The modulus of transfer function of data acquisition, playback, and digitizing systems. The gain of the amplifier is 120 dB.  $\zeta$  represents the geophone damping factor.  $H_p(f)$  and  $H_{RC}(f)$  are not included in the response shown (see text).  $R_D / (R_S + R_D)$  of  $H_g(f)$  is assumed to be unity in calculating these responses.



## Hot Springs in Northwestern Nevada

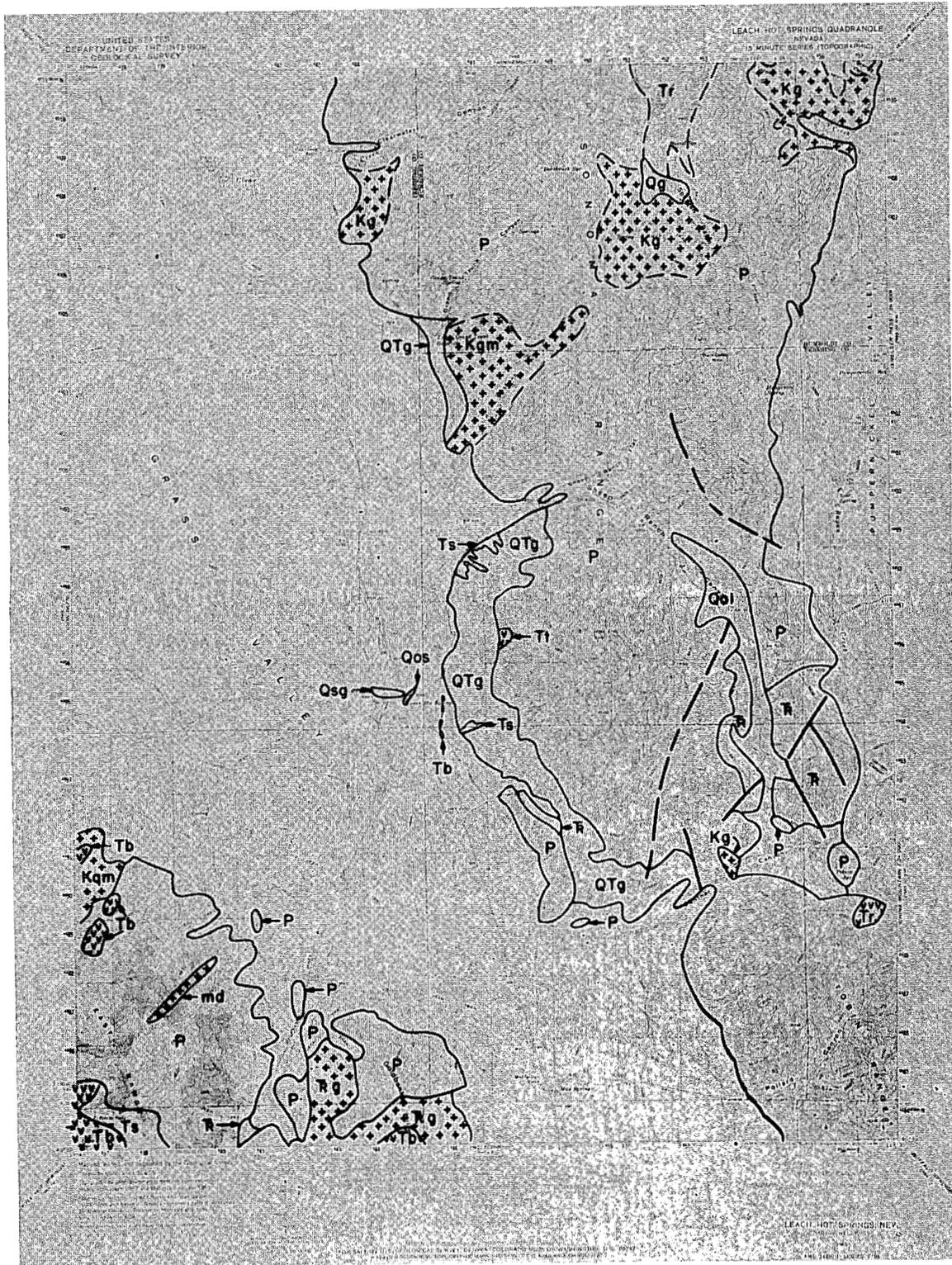
XBL 735 676

Figure 3.1



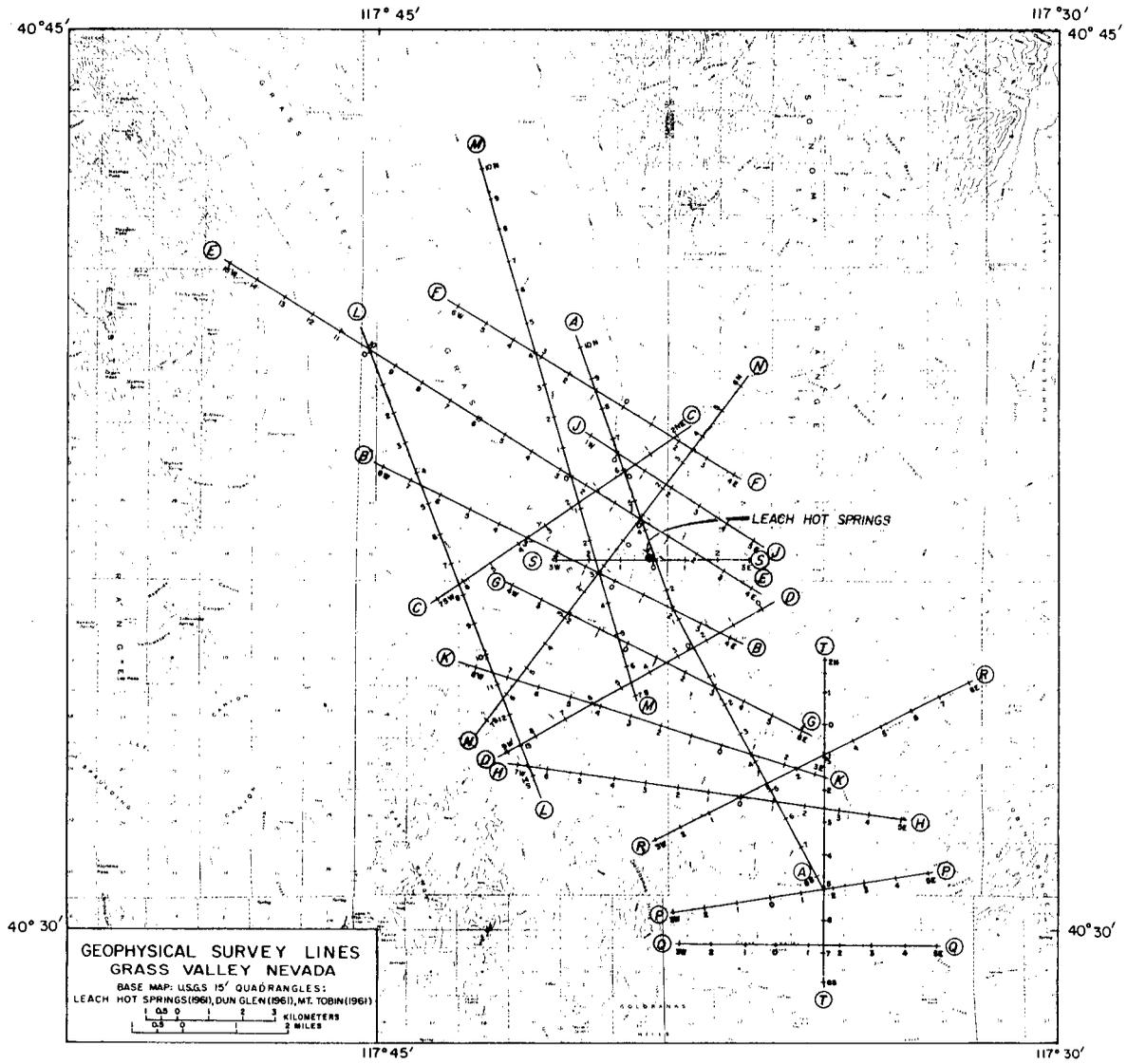
XBL 771-7423

Figure 3.2



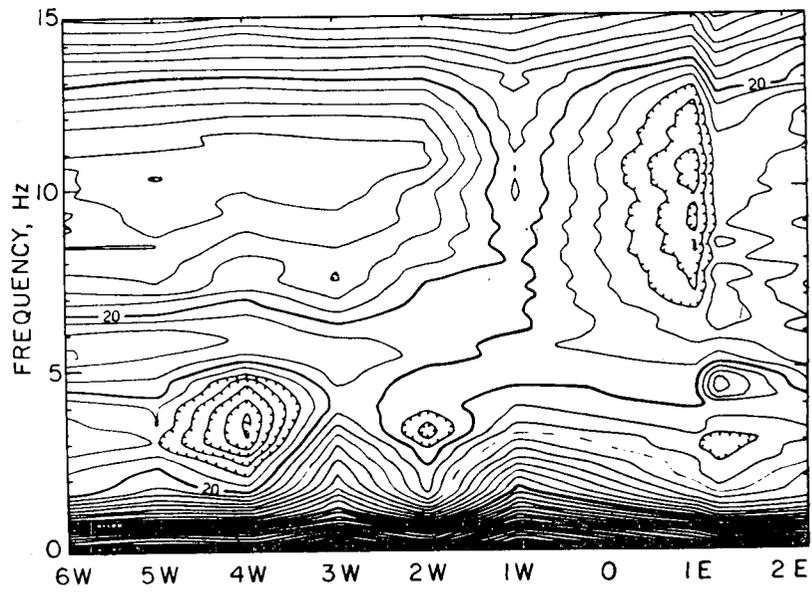
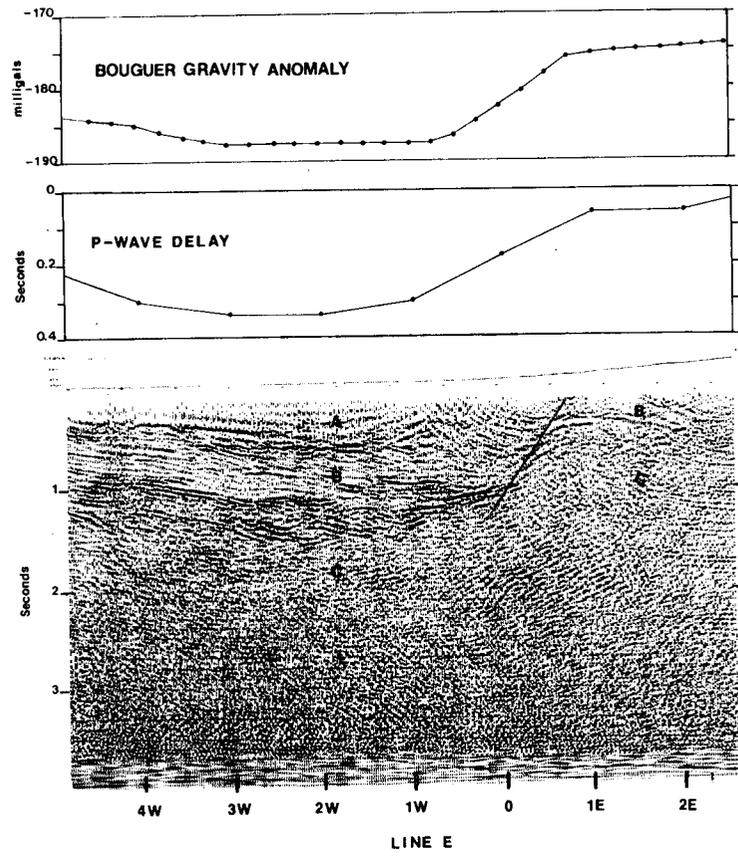
CBB 751-49

Figure 3.3



XBL 758-3669-A

Figure 3.4



XBL 777-9469A

Figure 3.5

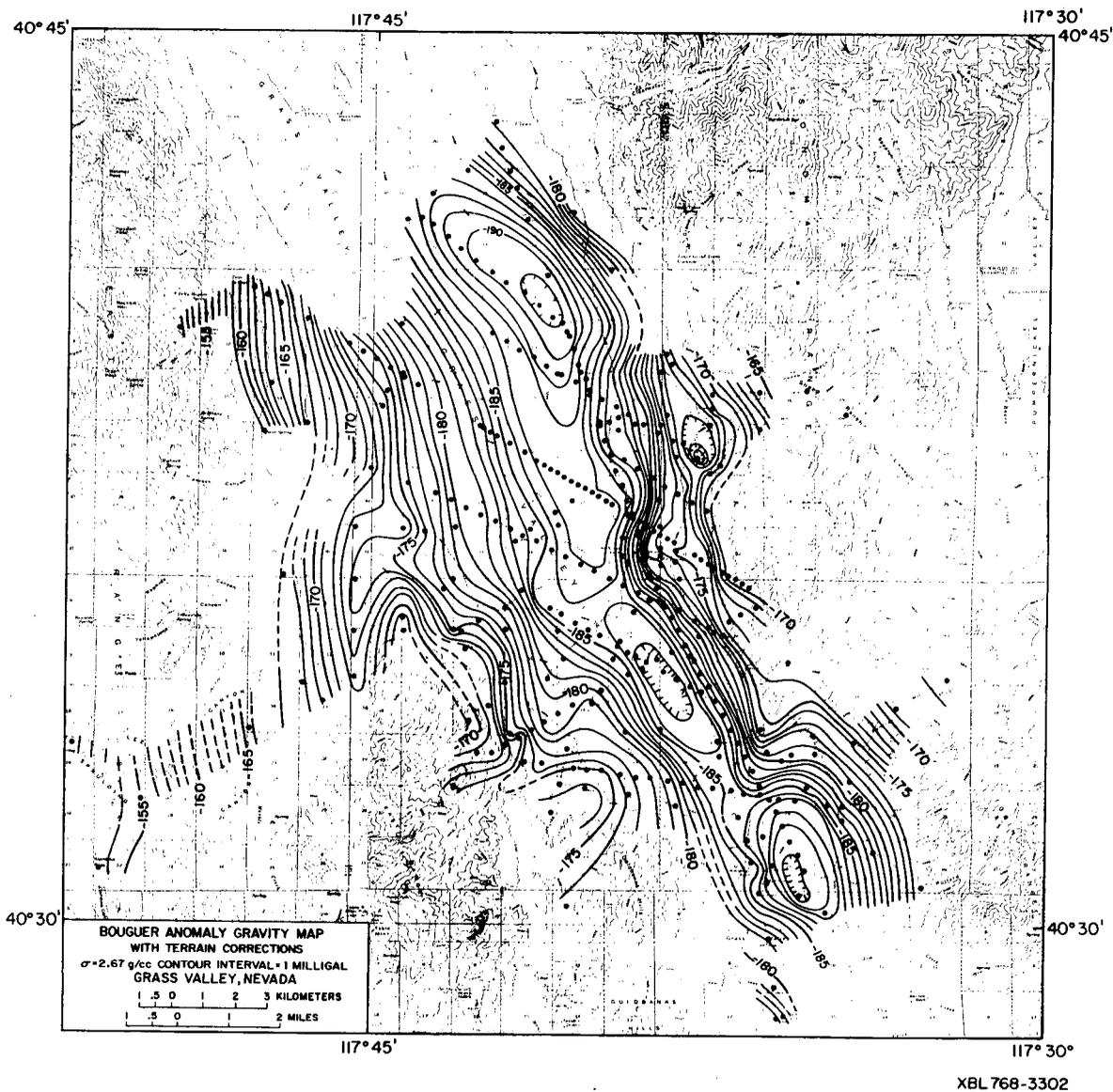


Figure 3.6

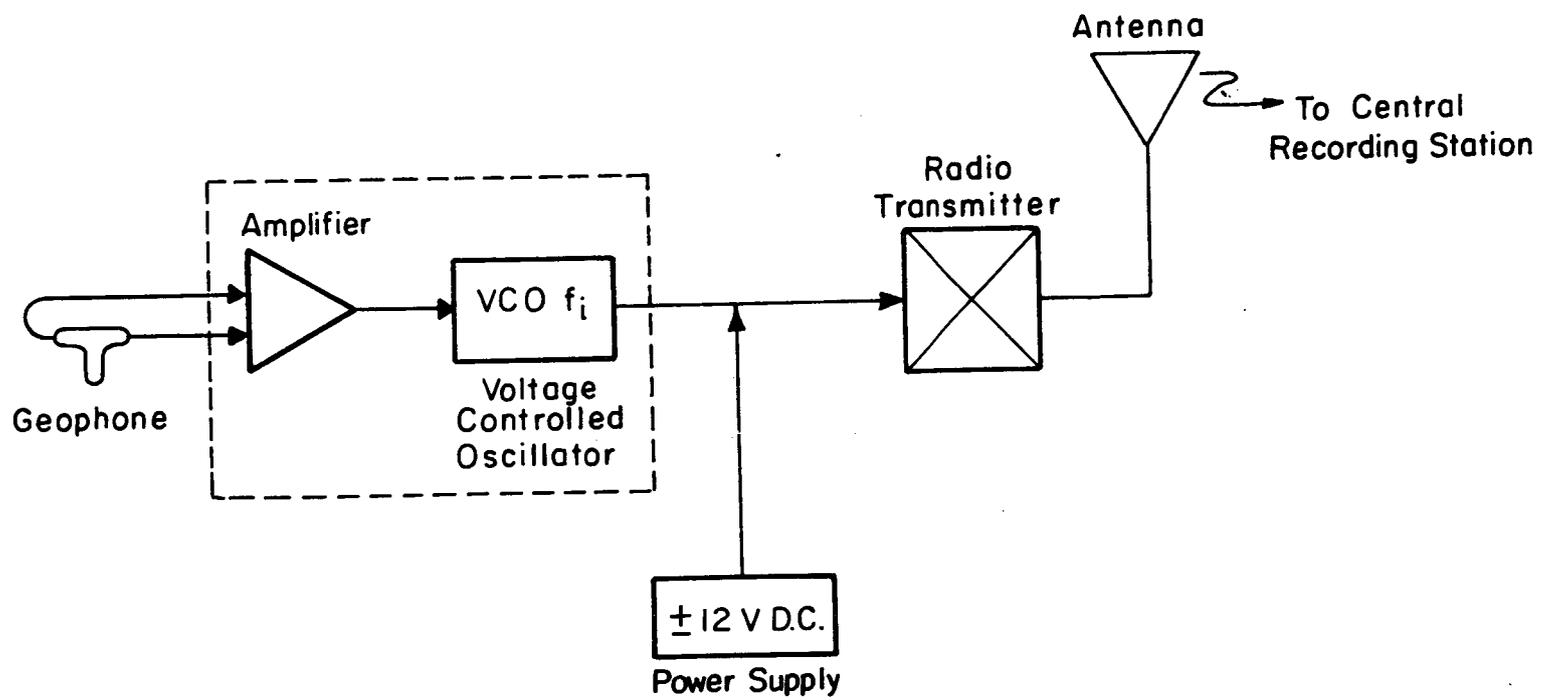
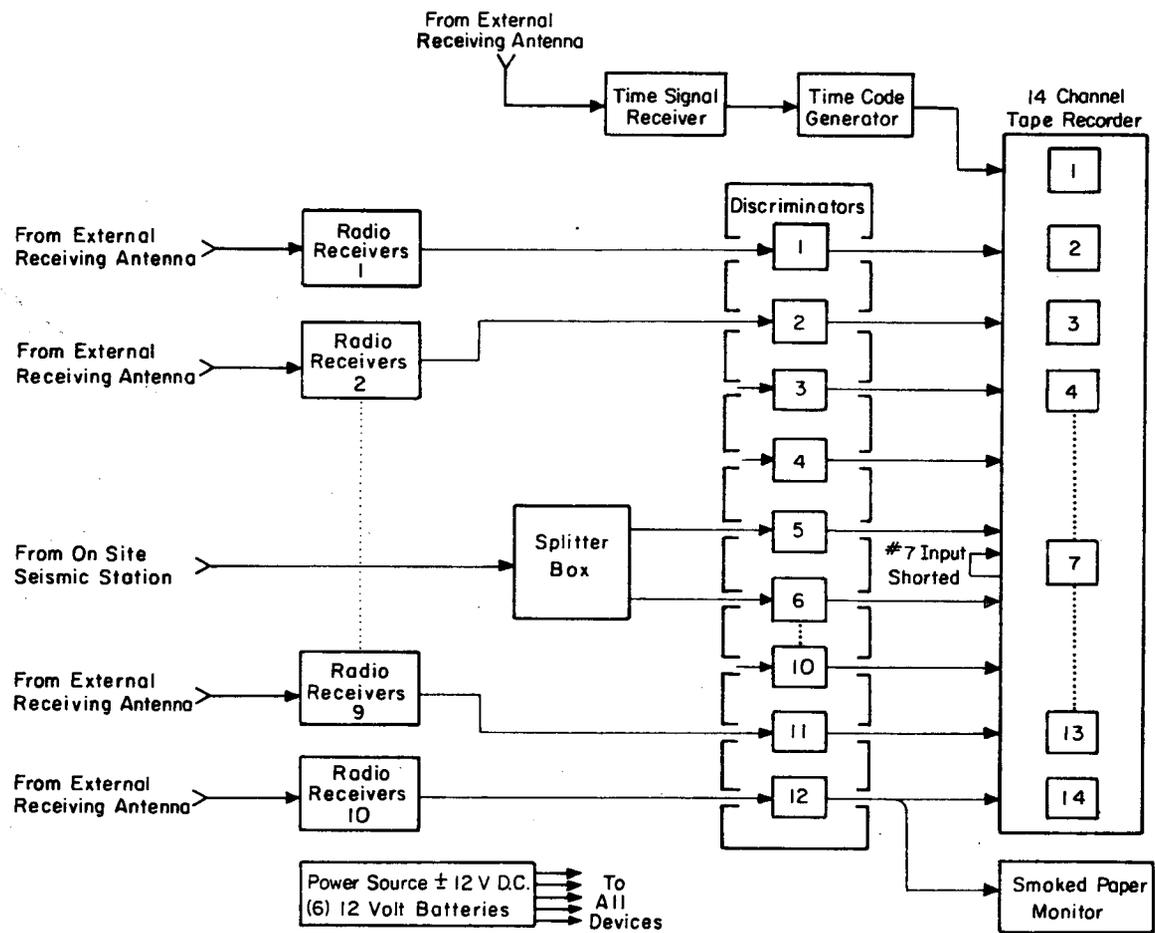


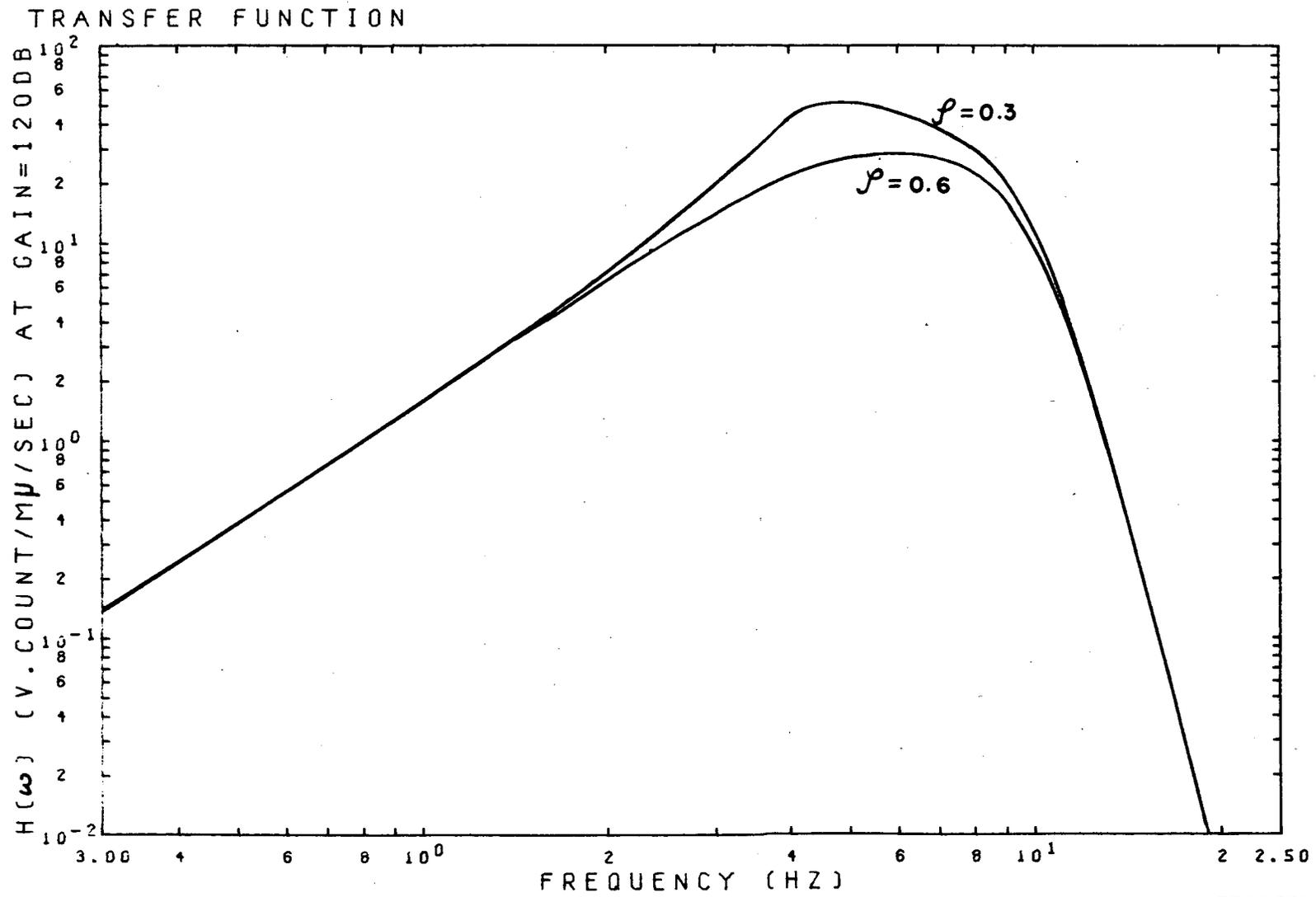
Figure 3.7

XBL 778-5997



XBL778-5996

Figure 3.8



XBL 772-7638

Figure 3.9

#### IV. AMPLITUDE VARIATIONS OF GROUND NOISE

##### 4.1 Introduction

The mapping of ground noise amplitude in selected frequency bands has often been used to locate presumed noise sources in geothermal areas. Such surveys, however, rarely give repeatable or easily interpretable results. We have investigated this exploration technique in Grass Valley by first establishing the general characteristics of the microseismic field, and then using this knowledge to design and execute an appropriate survey for more quantitative measurements. In this chapter, the field procedures, data processing techniques, and the observed temporal and spatial variations of seismic noise in Grass Valley are discussed. Weaknesses of the amplitude mapping technique in delineating buried noise sources are revealed in the discussion. In the final section of this chapter, the characteristic site-responses along line E and line G are compared for seismic noise and for waves from mine explosions. This comparison illustrates the amplification effects by the valley alluvium with respect to bedrock sites for shallow surface waves and for body waves.

##### 4.2 Field Procedures

To study the spatial variations of ground noise amplitude, we occupied a reference site at E2W (see Figure 3.2) throughout the survey period. Normally, we recorded overnight, with stations spaced at 1 km intervals along the survey lines. The smoked-paper monitor record was checked every morning to verify the occurrence of low seismic noise level at the reference site; otherwise, the sites were re-occupied

another night, until low-noise conditions prevailed. Geophones were buried about one foot below the surface. Before and after a survey, all geophones were buried in a common hole to verify uniformity of their responses.

#### 4.3 Data processing techniques: frequency spectrum estimation

It is well known that a stationary random process can be characterized by means of a power spectral density function. This function provides information on the power as a function of frequency for the process. For random processes, there are two power spectral density estimating techniques widely used, the autocorrelation method and the modified periodogram method.

##### 4.3.1 Autocorrelation method

The power spectrum,  $S_{nn}(f)$  of a function  $\phi_n(t)$  is defined as the Fourier transform of its autocorrelation function,  $C_{nn}(\tau)$ . To estimate the power spectrum, we filter, pre-whiten and detrend the time series, and then calculate the unbiased autocorrelation function for the  $i$ th data window of length  $L$  samples by the discrete formula,

$${}^i C_{nn}(\tau) = \frac{1}{L-\tau} \sum_{\ell=0}^{L-1-\tau} {}^i \phi_n(\ell) {}^i \phi_n(\ell+\tau), \quad \begin{array}{l} \tau=0, 1, \dots, J-1, \\ i=1, 2, \dots, I. \end{array} \quad (4.1)$$

The autocorrelation functions are averaged to give

$$\hat{C}_{nn}(\tau) = \frac{1}{I} \sum_{i=1}^I {}^i C_{nn}(\tau), \quad \tau=0, 1, \dots, J-1. \quad (4.2)$$

We utilize the Fast Fourier transform to obtain the power spectral density estimate by:

$$\hat{S}_{mn}(f_R) = \sum_{\tau=0}^{J-1} w(\tau) \hat{C}_{mn}(\tau) \exp(-j2\pi f_R \tau), \quad (4.3)$$

where  $f_R = \frac{R}{L}$  frequencies of discrete Fourier transform,  $-\frac{1}{2}, \dots, 0, \dots, \frac{1}{2}$ , and  $w(\tau)$  is a data window function, or so-called lag window.

The autocorrelation method is a reasonable technique to estimate the smooth spectrum. However, problems may arise in obtaining the power spectral density from the autocorrelation function due to certain kinds of the data window  $w(\tau)$ . Unless the transform of the window is entirely positive, there is a possibility that the computed power spectral density may be negative, a highly undesirable result, if the spectrum has a sharp peak. This is because the computed power spectral density is the convolution of the window transform and the transform of the estimated correlation functions. Unless the transform of the window is positive for all frequencies, the possibility exists that, due to statistical variation in estimating the correlation function, the resulting convolutions may produce negative values for the power spectrum at some frequencies. There are windows whose transforms are entirely positive, e.g., the triangular window, etc., and such windows should be used in cases where other windows lead to the undesirable result.

#### 4.3.2 Method of modified periodogram

The second and entirely equivalent estimation technique is based on the relation:

$$S_{nn}(f) = \lim_{T \rightarrow \infty} \frac{1}{2T} \left| \int_{-T}^T \phi(t) \exp(-j2\pi ft) dt \right|^2. \quad (4.4)$$

For a finite duration of time series, Equation (4.4) can be approximated to be

$$i_{S_{nn}}(f) = \frac{1}{T} \left| \int_0^T i_{\phi_n}(t) \exp(-j2\pi ft) dt \right|^2 \quad (4.5)$$

The left hand side of Equation (4.5) results from the modulus-squared of the finite Fourier transform, which is called the periodogram. In terms of discrete time series, the periodogram can be expressed as

$$i_{\phi_n}(f_R) = \sum_{\ell=0}^{L-1} w(\ell) i_{\phi_n}(\ell) \exp[-j2\pi \ell \frac{R}{L}] \quad (4.6)$$

$$= \sum_{\ell=0}^{L-1} w(\ell) i_{\phi_n}(\ell) \exp[-j2\pi \ell f_R], \quad R = -\frac{L}{2}, \dots, -1, 0, 1, \dots, \frac{L}{2}, \quad (4.7)$$

where  $w(\ell)$  is a data window whose purpose is to reduce the side-lobe amplitude hence reducing the effect of spectral leakage. The transform of Equation (4.7) is effected through the Fast Fourier transform algorithm, which requires the correction factor:

$$i_{X_n}(f) = 2(\Delta t) i_{\phi_n}(f_R), \quad (4.8)$$

where  $\Delta t$  is the sampling interval of the discrete time series.

The disadvantage of the periodogram as an estimate of the power spectrum is that the variance of  $i_{\phi_n}(f_R)$  is approximately  $S_{nn}^2(f)$ , under conditions of reasonable regularity, even when based on a lengthy stretch of data (Brillinger, 1975). One way to reduce the variance of the estimate is to average several statistically independent periodograms according to

$$\hat{S}_{nn}(f) = \frac{1}{U} \frac{1}{T} \sum_{i=1}^I |i_{X_n}(f)|^2, \quad (4.9)$$

where  $I$  is the number of periodograms, and  $U$  is the energy in the data window, defined as

$$U = \left[ \frac{1}{L} \sum_{\ell=0}^{L-1} w^2(\ell) \right] T. \quad (4.10)$$

where  $T = \Delta t(L-1)$ .

As  $L$  approaches infinity  $\hat{S}_{nn}(f)$  approaches asymptotically  $S(f)\chi_{\nu}^2/\nu$  if  $2\pi f \equiv 0 \pmod{\pi}$ , and asymptotically  $S(f)\chi_{\nu}^2 / (\nu/2)$  if  $2\pi f = \pm\pi, \pm 3\pi, \dots$  etc. (Brillinger, 1975), where  $\chi_{\nu}^2$  is a chi-squared variate with  $\nu$  degrees of freedom. The number of degrees of freedom  $\nu$  of the smoothed estimator is

$$\nu = 2 \cdot I \cdot b_1, \quad (4.11)$$

where  $b_1$  is the standardized bandwidth of data window (Jenkins and Watts, 1968) given by

$$b_1 = \frac{L}{\sum_{\ell=0}^{L-1} w^2(\ell)}. \quad (4.12)$$

$b_1 = 1$  for rectangular data window. This leads to the 100 $\gamma$  percent confidence interval for  $S(f)$  to be:

$$\frac{\nu \hat{S}_{nn}(f)}{\chi_{\nu}^2 \left( \frac{1+\gamma}{2} \right)} < S(f) < \frac{\nu \hat{S}_{nn}(f)}{\chi_{\nu}^2 \left( \frac{1-\gamma}{2} \right)} \text{ for } 2\pi f \not\equiv 0 \pmod{\pi}, \quad (4.13)$$

where  $\gamma = 0.9$  for 90% confidence limits. The velocity spectral density, VSD or  $V_n(f)$ , is obtained by taking the square root of the power spectral density and correcting it for system response  $H_D(f)$  as follows:

$$V_n(f) = \frac{\sqrt{\hat{S}_{nn}(f)}}{|H_D(f)|} \text{ m sec}^{-1} \text{ Hz}^{-1/2}, \quad (4.14)$$

where  $H_D(f)$  is given in Equation(3.6). We normally present VSD in units of millimicrons (nanometers,  $10^{-9}$ m) $\text{sec}^{-1}\text{Hz}^{-1/2}$ .

#### 4.3.3 Grass Valley data processing method

Data were selected from the quietest recording period in the early morning hours. At least 28 simultaneous blocks of data were chosen from each of the recording stations, avoiding any spurious transient events. Each data block of length 12.8 seconds was filtered and digitized. The resulting 512-point records were tapered by a 10% cosine data window and Fourier transformed. The power spectral density function is estimated by the method of modified periodogram.

#### 4.4 Temporal variation of ground noise

The total seismic noise amplitude  $\sigma(x,y,t,f)$  can be modeled very generally as the sum of three sorts of noises,

$$\sigma(x,y,t,f) = \sigma_i(x,y,t,f) + \sigma_m(x,y,t,f) + \sigma_l(x,y,t,f), \quad (4.15)$$

where

$\sigma_i(x,y,t,f)$  is the intrinsic noise at the site, including geothermal noise,

$\sigma_m(x,y,t,f)$  is the microseismic component from distant sources, and

$\sigma_l(x,y,t,f)$  is the noise generated locally at the surface by human activity and atmospheric disturbances.

If we are interested only in intrinsic noise, the sampling and processing procedures must exclude the effect of the other two noise sources. To minimize local sources,  $\sigma_l(x,y,t,f)$ , the data must be taken between midnight and dawn, because normally the noise level is low.

Figure 4.1 presents the diurnal variation of seismic noise at the reference site E2W. To construct this figure, transient-free noise data were chosen to estimate the VSD every hour for a 30 hour period. Roughly 6 minutes of seismic noise actually went into each hourly average. The spectral density was then contoured as a function of time and frequency. It can be seen that noise is high over the whole band of analysis from 9 A.M. to 7 P.M., the result of more disturbed daytime meteorological conditions and cultural activity. This confirms a well known result that seismic noise VSD is minimum at 2-4 A.M. local time, and this was the period we sampled for the best data.

We found, in Grass Valley, that the time of minimum ground noise at the reference site, E2W, coincides with the quietest period at all other sites in the region. For example, simultaneous data were sampled every hour from stations E2W and K1.5E, located approximately 11 km apart. The time-varying VSD over a 12 hour period from 9 P.M. to 9 A.M. next morning are presented in Figure 4.2. The figure indicates that at 2 A.M. the data are the quietest at both sides.

A typical survey is carried out over a period of several days, so that long term secular variations are apparent in the data. The nature of this variation over a 9 day period at the reference site, E2W, is shown in Figure 4.3. We estimate one VSD every 24 hours, using the quietest data during early morning hours, and contour the VSD from day 211 to day 219. In this figure, the high amplitude seismic noise appears from day 214 to day 216 and is related to regional weather conditions. On those three days, there were thunderstorms starting in the afternoon and ending in the early evening throughout the region.

In order to eliminate temporal variations of the observed microseisms, the band-limited power of seismic noise at each site, obtained by integrating VSD over the frequency band of interest, is normalized by the simultaneous power in the same frequency band at the reference site, provided that data are both sampled from the quiet period in early morning. Mapping the normalized power gives the spatial distribution of relative intrinsic noise power level.

#### 4.5 Spatial variation of ground noise

Estimation of ground noise VSD from simultaneous sampling in the early morning, with stations at 1 km spacing, yields relative intrinsic noise power contour maps as illustrated for the frequency band of 2-4 Hz (Figure 4.4A), 5-7 Hz (Figure 4.4B) and 10-12 Hz (Figure 4.4C). High noise levels are found at Leach Hot Springs and near the center of Grass Valley, as anticipated, but there are also local anomalies such as in the areas around G2W and G3W, H1E and H2E (see Figure 3.2 for site locations). Those ground noise anomalies, especially in 5-7 Hz band, correlating spatially with the occurrence of Bouger gravity anomalies (Figure 3.6), imply the occurrence of thickest alluvial deposits. The long-term stability of these anomalies is reproducible as indicated by close agreement with the results of a preliminary survey carried out in the summer of 1975, a year earlier than the time at which data shown were taken.

Leach Hot Springs clearly generates seismic noise, but the noise is localized and does not propagate unattenuated more than a few km. In the vicinity of the springs, noise spectra show the high amplitude

seismic noise over a wide frequency band; 500 meters northwest of the hot springs (A3.7N) the amplitude of the noise at all frequencies greater than 1 Hz has attenuated nearly 20 dB. The spectrum of the hot springs noise can be seen in Figure 4.5 (top) which compares the hot springs site with A3.7N and the valley edge site, AC. Note the wide-band nature of the hot springs noise.

In the valley center, station E5W, the noise has a distinctive broad peak around 5.5 Hz, as can be seen at the bottom of Figure 4.5. The character of the broad valley peak varies from site to site, probably as a consequence of changes in near surface properties. In Figure 4.4B, the areas of high amplitude seismic noise in 5-7 Hz band generally correspond to the areas of thick alluvium. The details of noise variation across the valley are illustrated by data for three typical survey lines, E, B, and G, shown in Figures 4.6, 4.7, and 4.8.

The instantaneous ground noise level along 8.25 km of line E is presented in Figure 4.6 for three different times of recording. Data were taken simultaneously from sites at E6W, 5W, 4W, 3W, 2W, 1W, 1E, 1.25E, and 2.25E. In this figure there is a clear peak at 5.5 Hz extending westward. The source of this well-defined and band-limited peak is not clearly understood, though it is doubtless related to near-surface properties. (We know it is a surface wave with a wavelength of about 50m). A wide-band ridge of relatively large amplitude noise appears at E3W, and is frequently seen to extend to 1W. Maximum valley fill and lowest topography occurs around 2W. A remarkable feature seen in the figure is the dramatic 10 dB contrast between points 1E and 1.25E, spanning the Hot Springs fault. It seems the

local noise field, generated by hot springs, is less attenuated east of the fault than west of it, probably due to high-Q surface rocks on the east being in faulted contact with alluvium west of the fault. This geological feature can be seen in the faults map (Figure 3.2) as well as in the Bouguer gravity anomaly, the P-wave delay profiles, and the seismic reflection section (Figure 3.5).

Asymmetrical ridges of wide-band noise with sharp gradients to the east are seen near 2W in line B (Figure 4.7) and near 1E in line G (Figure 4.8). These ridges in the noise contours, as was the case for line E, correspond in position to the location of the minimum Bouguer anomaly along each line and to the location of the thickest alluvium (Beyer, et al., 1976). The positions of high gradients in ground noise east of the noise ridges on lines B and G apparently correlate with locations of shallow faults. The prominent broad peak of 6.5 to 7 Hz, seen at G3W in Figure 4.8 and C2.5W and C4.5W in Figure 4.9, are probably also related to properties of shallow alluvium.

At the south end of Grass Valley, the ground noise level is generally lower than at the north, and this contrast is presumably due to larger distance from the hot springs and thinner alluvial deposits to the south. The noise profiles along those survey lines in the southern part of the valley, e.g. line H (Figure 4.10), line K (Figure 4.11) and line R (Figure 4.12), do not show high gradients. The close similarity in the noise profiles along lines H,K is not surprising since the shallow geology is similar along the lines. The consistent anomaly appearing in the vicinity of H2E, K1.5E and R1E in the frequency band of 3 to 6 Hz may result from the localized occurrence of thick alluvium.

We see strong evidence, then, that the conventional ground noise survey reveals anomalies due not only to radiating sources, but also to variation in shallow geological structure, even after diurnal and secular variations are carefully eliminated. The noise power mapping technique cannot discriminate between the anomaly due to a buried seismic source and that associated with alluvial response. The method does provide an alternate way to map shallow geology and to detect lateral variations of near-surface structure.

#### 4.6 Site-response characteristics

We conclude from the previous section that the spatial distribution of microseismic amplitude in a particular frequency band is strongly affected by the properties of near-surface material. The VSD of a bedrock site, away from active sources, always shows a smooth spectrum with no dominant peaks, whereas the VSD at an alluvium site always shows a well-defined peak in the spectrum. The spectral peak doubtless results from frequency-selective amplification related to the propagation characteristics of surface waves in the section of alluvium. Similar site-responses have been reported by Kanai and Tanaka (1961) and Katz (1976). Kanai and Tanaka (1961) suggested that at a given site the microtremor response correlates with the period distribution curve of local earthquakes. In Japan, microtremor recording is used extensively to determine the predominant frequency associated with various subsoil structures (Kanai, et al., 1966). The results of such measurement have been used to determine "foundation coefficients" in earthquake-resistant construction. In a similar study in the San Francisco Bay area,

Borcherdt (1970) compared the microseism-derived site-responses with the spectral amplification factors derived from nuclear explosion data. Borcherdt observed that the spectral peaks of the horizontal component of the microseisms agree with those of the nuclear explosion data, but that the predominant frequencies of horizontal microseisms do not always coincide with those of the vertical component. In contrast to the above authors, Udawadia and Trifunac (1973) found no correlation between the spectra of microtremors and the ground's response to earthquakes recorded in El Centro, California.

In Grass Valley, we have investigated the correlation between earth noise spectra and seismic event spectra using signals generated by mine blasts at the Duval Mine, some 50 km to the east. Vertical-component data were recorded simultaneously along survey lines. The VSD of the explosion arrival at each site is estimated from a data block of 6.4 sec (256 data points) and averaged over 5 adjacent frequency components. The explosion arrivals along line E are presented in Figure 4.13. The corresponding background noise data taken a few minutes before the explosion arrivals are shown in Figure 4.14. The analog records of the explosion do not show significant differences in amplitude and frequency characteristics over the line, while the background noise in the valley sites (E2W, E3W, E4W, E5W, and EbW) is apparently different from that in hill sites (E2, 25E, E1, 25E, E1E, E1W). The spectra of explosion arrivals and background noise, presented in Figure 4.15, show little similarity. We removed the source characteristics of mine blast arrivals by normalizing the VSD at each recording site, using the VSD of site E1E as a reference. The identical

normalization was applied to the VSD of background noise. The normalized results along line E are shown in Figure 4.16, where the resonant effect of valley fill in the band around 5 Hz is clearly shown on the noise spectrum but not on the blast spectrum. There is very little similarity between variation in the mine blast arrival and variation in background noise. The similar data set along the survey line G also shows very little correlation between blast arrivals and noise (Figure 4.17).

FIGURE CAPTIONS

- Figure 4.1 Diurnal variation of ground noise level at reference site E2W, from hour 10, day 212 to hour 16, day 213, in 1976. Noise levels are normalized with respect to  $10^{-11}$  m/sec/ $\sqrt{\text{Hz}}$ , (0dB), and contour interval is 2 dB. Note the minimum noise level (hachured) between 2-4 A.M. for all frequencies greater than 2 Hz.
- Figure 4.2 The simultaneous temporal variation of ground noise over a 12 hour period at sites E2W and K1.5E, approximately 11 km apart. Noise levels are normalized with respect to  $10^{-11}$  m/sec/ $\sqrt{\text{Hz}}$ , (0dB), and contour interval is 5 dB. Note that the minimum noise level is coincident at both sites in the early morning.
- Figure 4.3 Secular variation of early morning quiet ground noise level from day 211 to day 219 of 1976 at E2W with respect to  $10^{-11}$  m/sec/ $\sqrt{\text{Hz}}$ , (0dB). Contour interval is 2 dB. Thunderstorm and unsettled regional weather characterized days 214-216, the period of greatest early morning noise level.
- Figure 4.4 The power contours of relative intrinsic noise with respect to reference site E2W in three frequency bands. Contour interval is 3 dB. Solid circles indicate sampling points. (A) 2-4 Hz, (B) 5-7 Hz, (C) 10-12 Hz.
- Figure 4.5 Velocity spectral density (VSD) of ground noise at Hot Springs and at site A3.7N, 500 m NW of the hot springs (upper) and at E5W, at center of the valley (lower)

compared to bedrock site AC, at the valley edge. The error bars for A3.7N and AC sites are 95% confidence limits for the estimated VSD. Data represent spectral averages over 32 data blocks of 12.8 sec length, for each site.

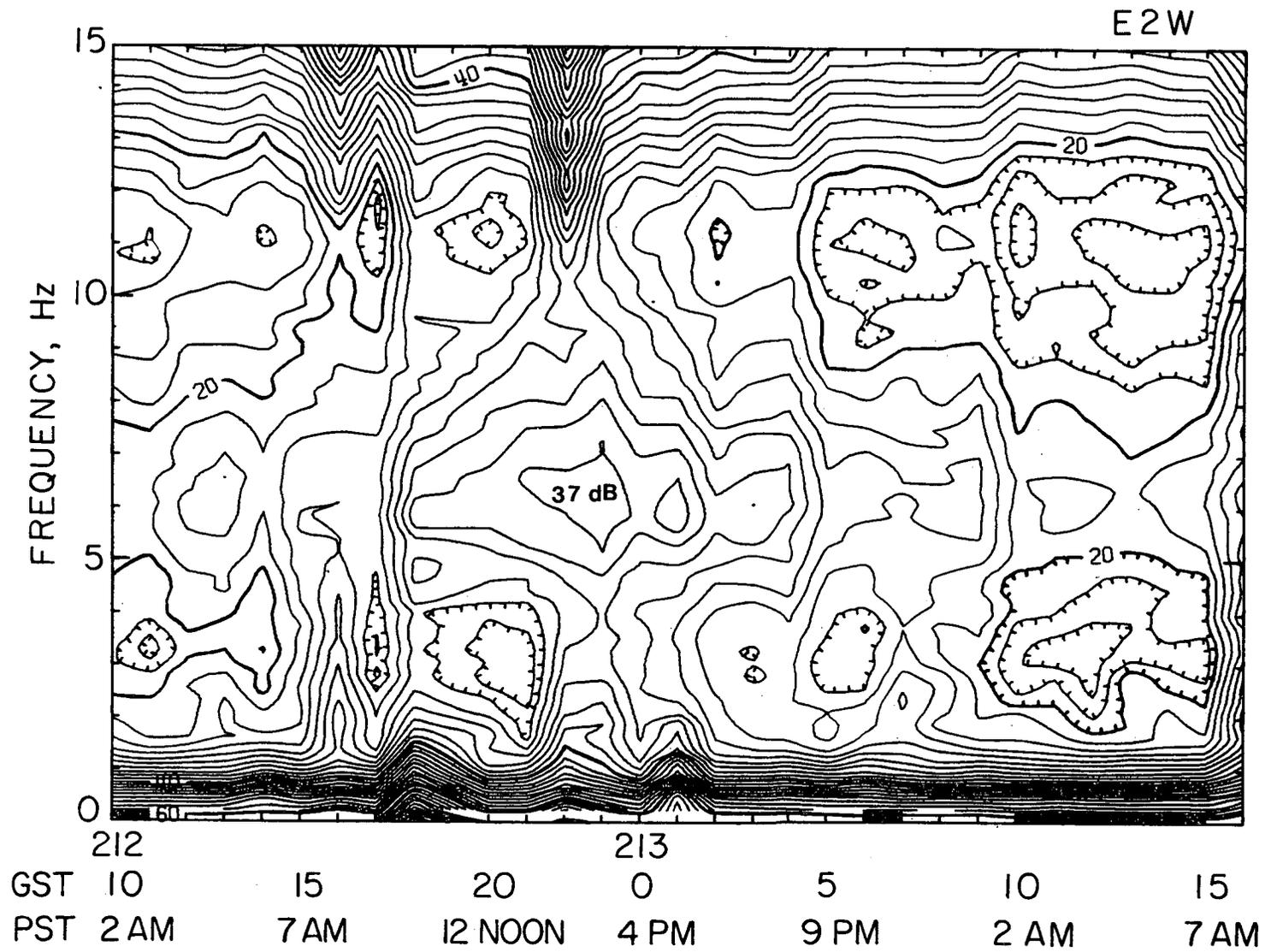
- Figure 4.6 Instantaneous noise field along survey line E for three different quiet periods. Abscissa is station location, with 1 km spacing, and ordinate is frequency. Contour interval is 2 dB. Note the high wide-band noise level at 1W (A and B) and 3W (C), the region of thick alluvium, and also the consistently sharp gradient across the valley margin fault traced to 1E. Note also the typical valley resonant peak near 6 Hz to the west.
- Figure 4.7 Instantaneous noise field along survey line B. Abscissa is station location, with 1 km spacing, and ordinate is frequency. Contour interval is 2 dB. Note the high wide-band noise level at the valley center near 2W.
- Figure 4.8 Instantaneous noise field along survey line G. Abscissa is station location, with 1 km spacing, and ordinate is frequency. Contour interval is 2 dB. Note the high wide-band noise level at the valley center near 1E. Sharp gradients may indicate valley faults.
- Figure 4.9 Instantaneous noise field along portion of survey line C. Abscissa is station location, with 1 km spacing, and ordinate is frequency. Contour interval is 2 dB. Note the band-limited peaks at both 2.5 W and 4.5 W, probably the result of alluvial resonant effects.

- Figure 4.10 Instantaneous noise field along survey line H. Abscissa is station location, with 1 km spacing, and ordinate is frequency. Contour interval is 2 dB. Note the gentle increase in noise level toward the east reaching a peak at 2E, which may correspond to resonance of the alluvial layer.
- Figure 4.11 Instantaneous noise field along survey line K. Abscissa is station location, with 1 km spacing, and ordinate is frequency. Contour interval is 2 dB. Note the gentle increase in noise level toward the east, reaching a high frequency ridge at 1W and a low frequency peak at 1.5E, which corresponds to local thick alluvial layer.
- Figure 4.12 Instantaneous noise field along survey line R. Abscissa is station location, with 1 km spacing, and ordinate is frequency. Contour interval is 5 dB. Note the quiet nature of the noise field along the line, the low frequency peak at 1E corresponds to the same features at H1E, H2E and K1.5E.
- Figure 4.13 Mine blast arrivals recorded at sites along survey line E. Note the early arrival at E1E, in the vicinity of Leach Hot Springs. The scale factor is  $1230 \times 10^{-9}$  m/sec per cm of the amplitude of displayed record at 4.5 Hz. The geophones of identical response were used.
- Figure 4.14 Typical seismic noise data at sites along survey line E. The scale factor is  $129 \times 10^{-9}$  m/sec per cm of the amplitude of displayed record at 4.5 Hz. The geophones of identical response were used.

Figure 4.15 Velocity spectral densities (VSD) of the mine explosion event (upper curve) and the seismic noise (lower curve) at each recording site along line E. Abscissa is frequency and ordinate is VSD in  $10^{-9}$  m/sec/ $\sqrt{\text{Hz}}$ . Scale is indicated in Frame H.

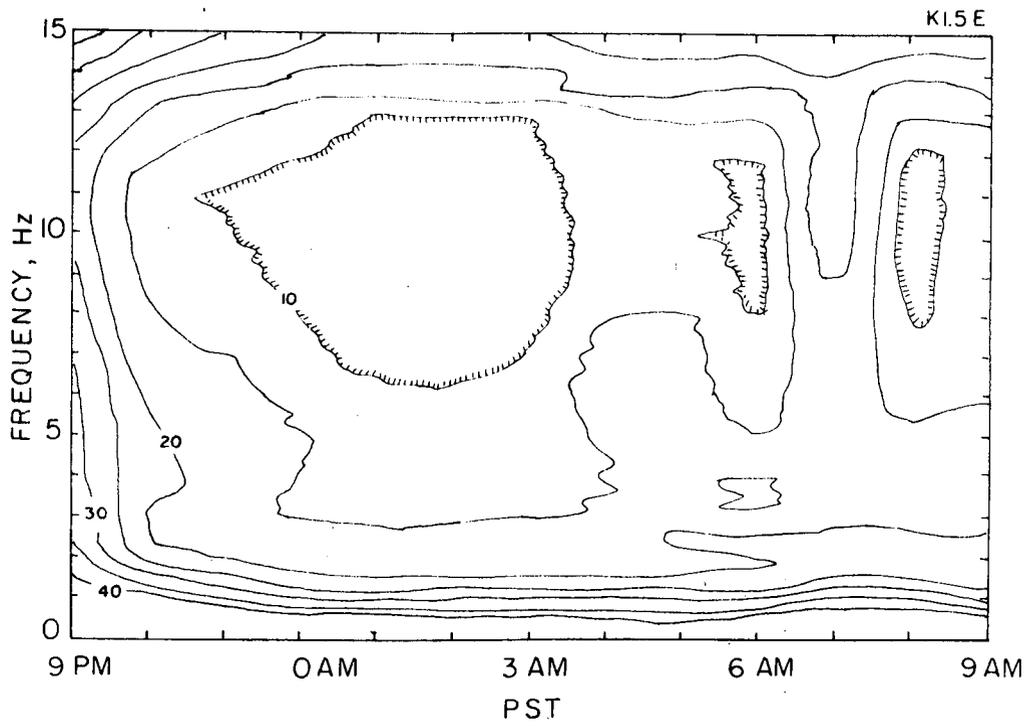
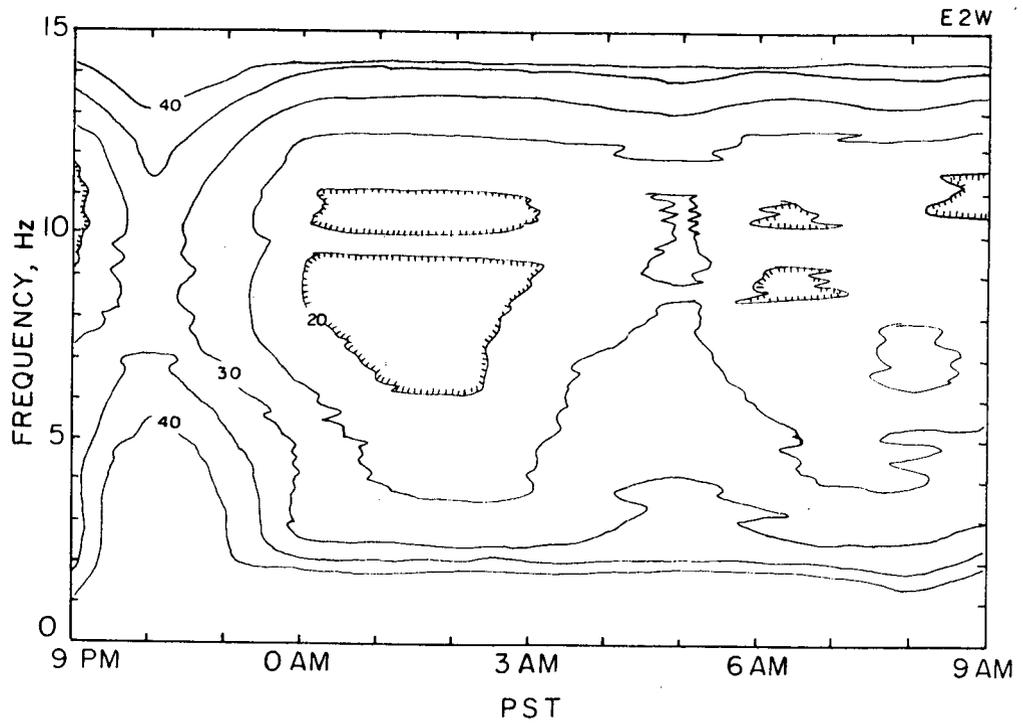
Figure 4.16 Relative VSD at sites along survey line E. VSD in Figure 4.15 are normalized with respect to site E1E. Solid curves are mine blast arrivals and the dotted curves are seismic noise.

Figure 4.17 Velocity spectral densities (VSD) of the mine explosion event (upper curve) and the seismic noise (lower curve) at each recording site along line G. Abscissa is frequency and ordinate is VSD in  $10^{-9}$  m/sec/ $\sqrt{\text{Hz}}$ .



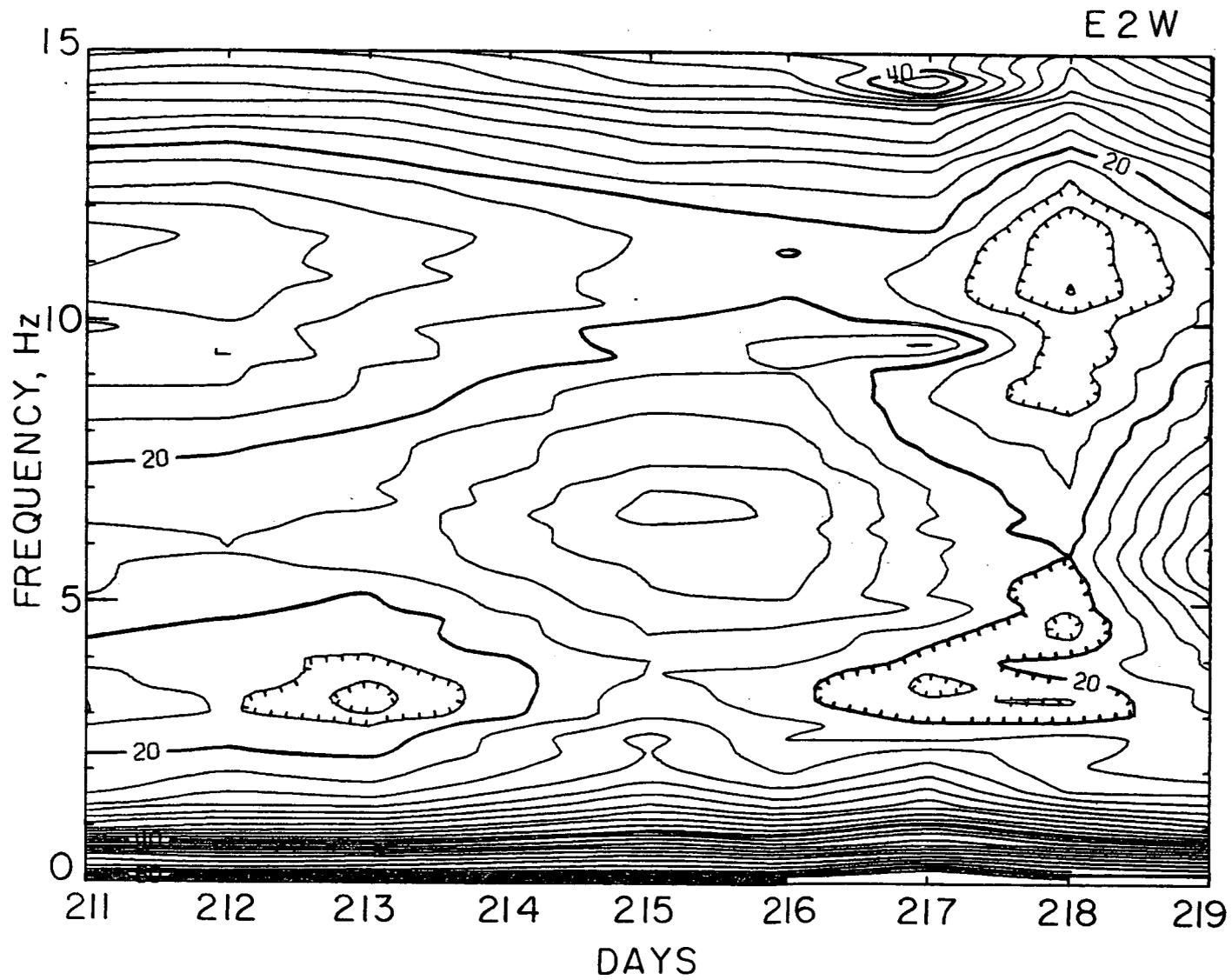
XBL 771-5001

Figure 4.1



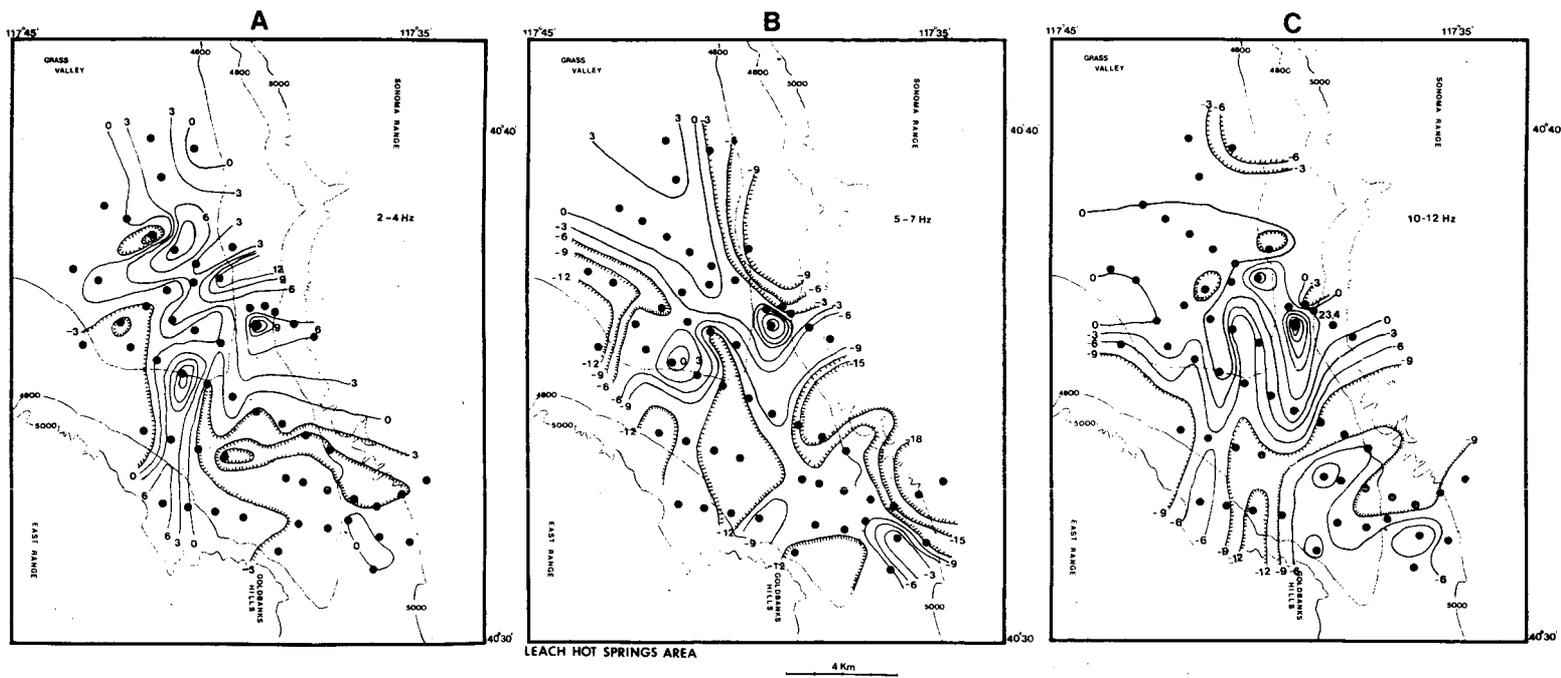
XBL 777-5862

Figure 4.2



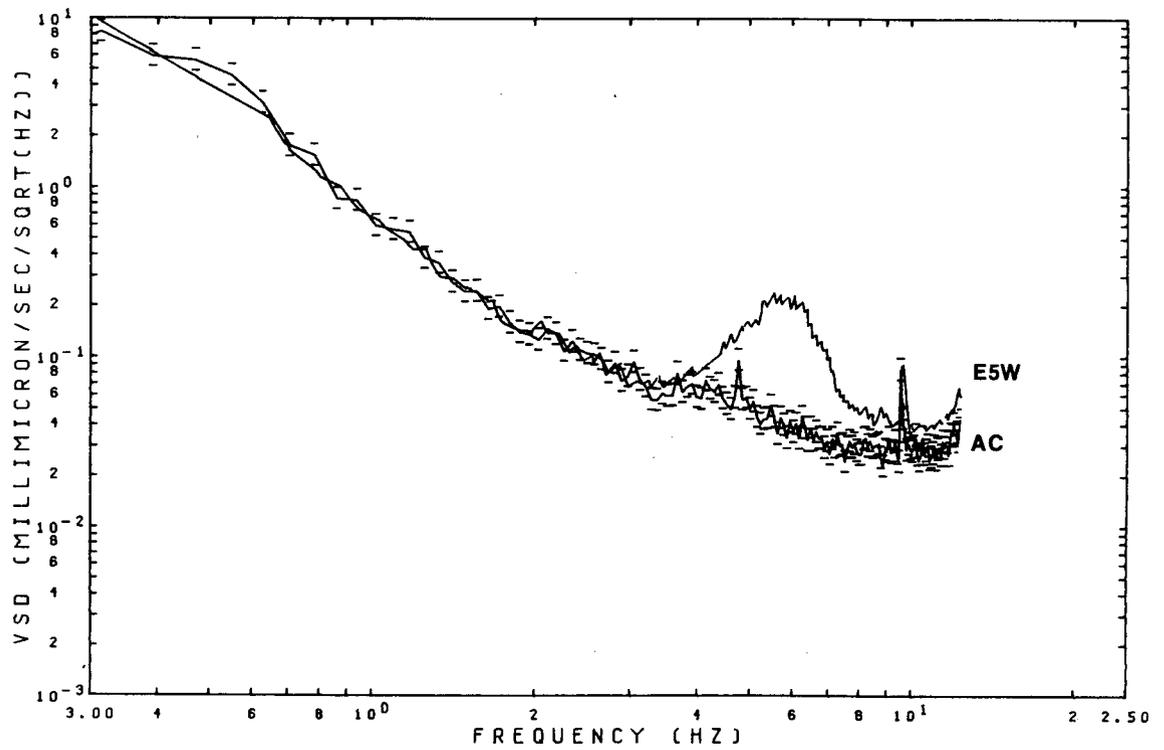
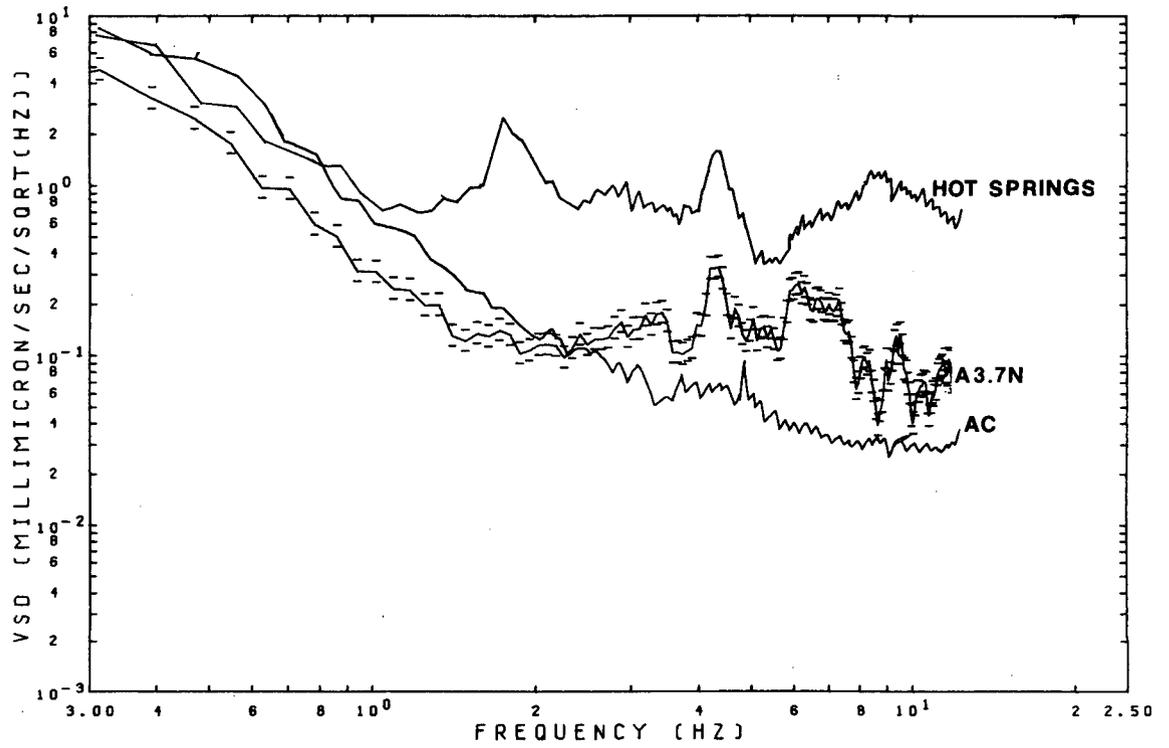
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Figure 4.3



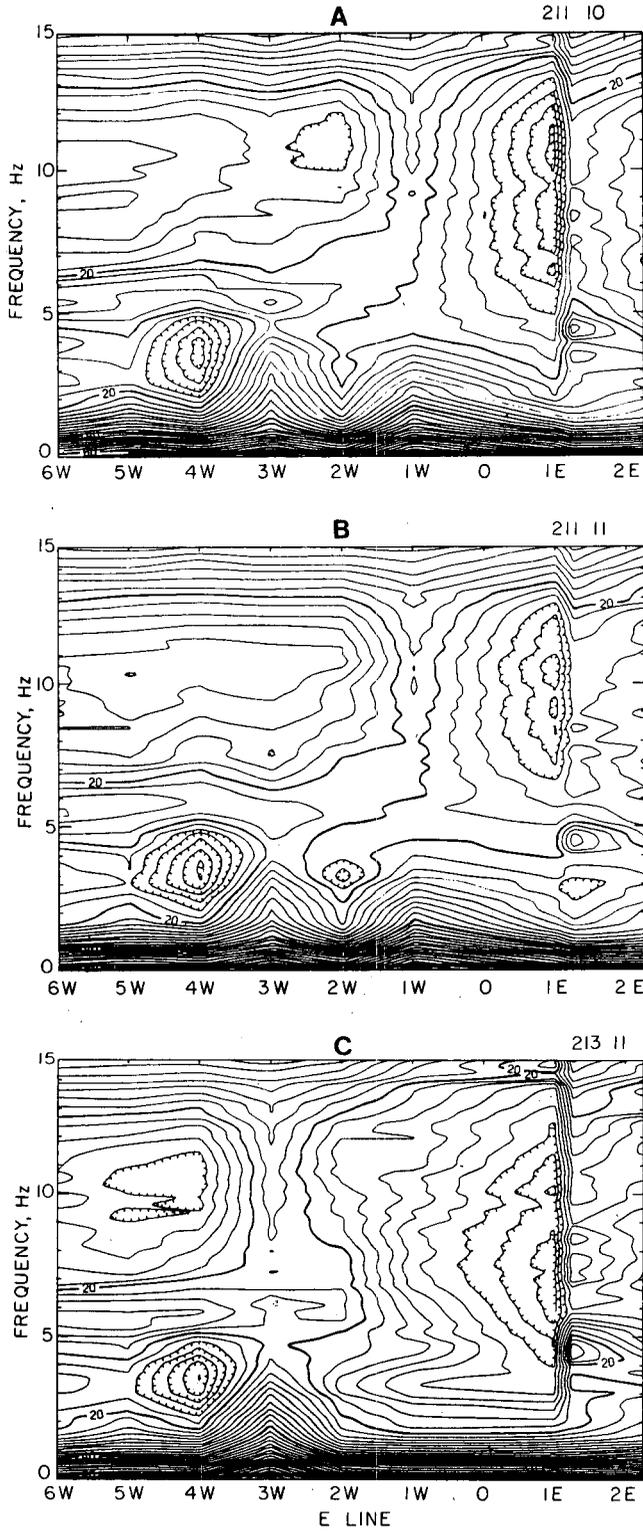
XBL 776-9499

Figure 4.4



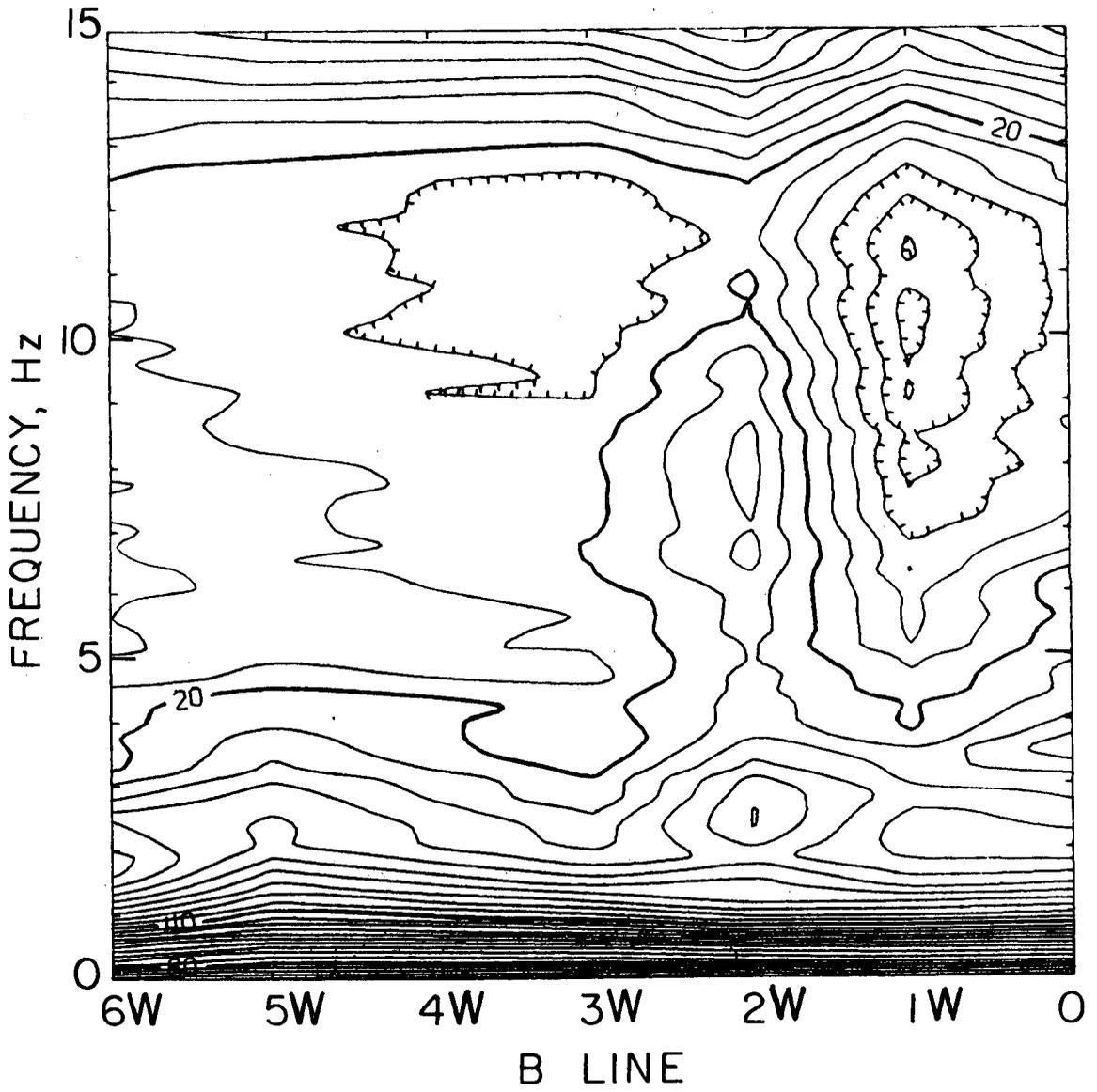
XBL 776-9466

Figure 4.5



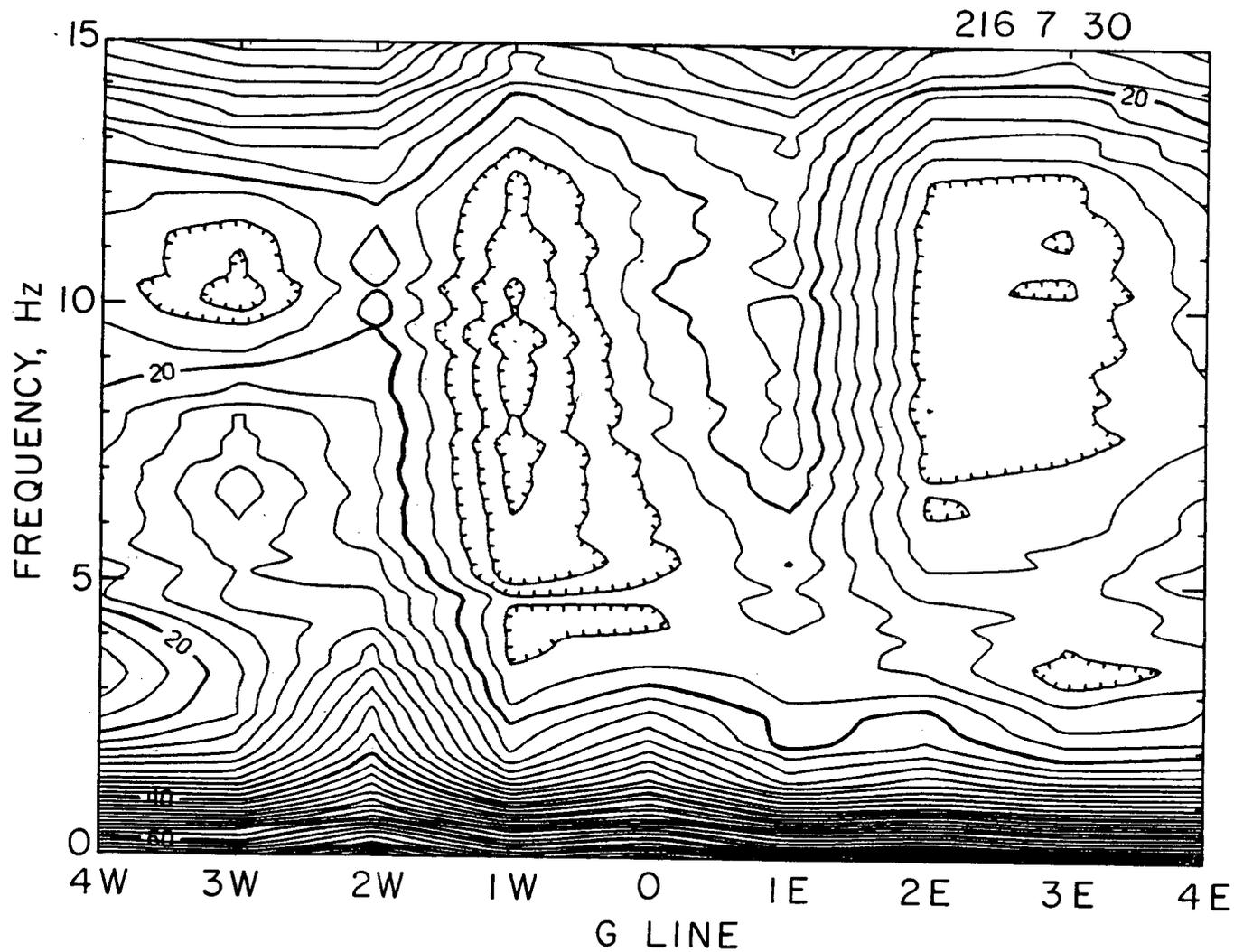
XBL 771-4993A

Figure 4.6



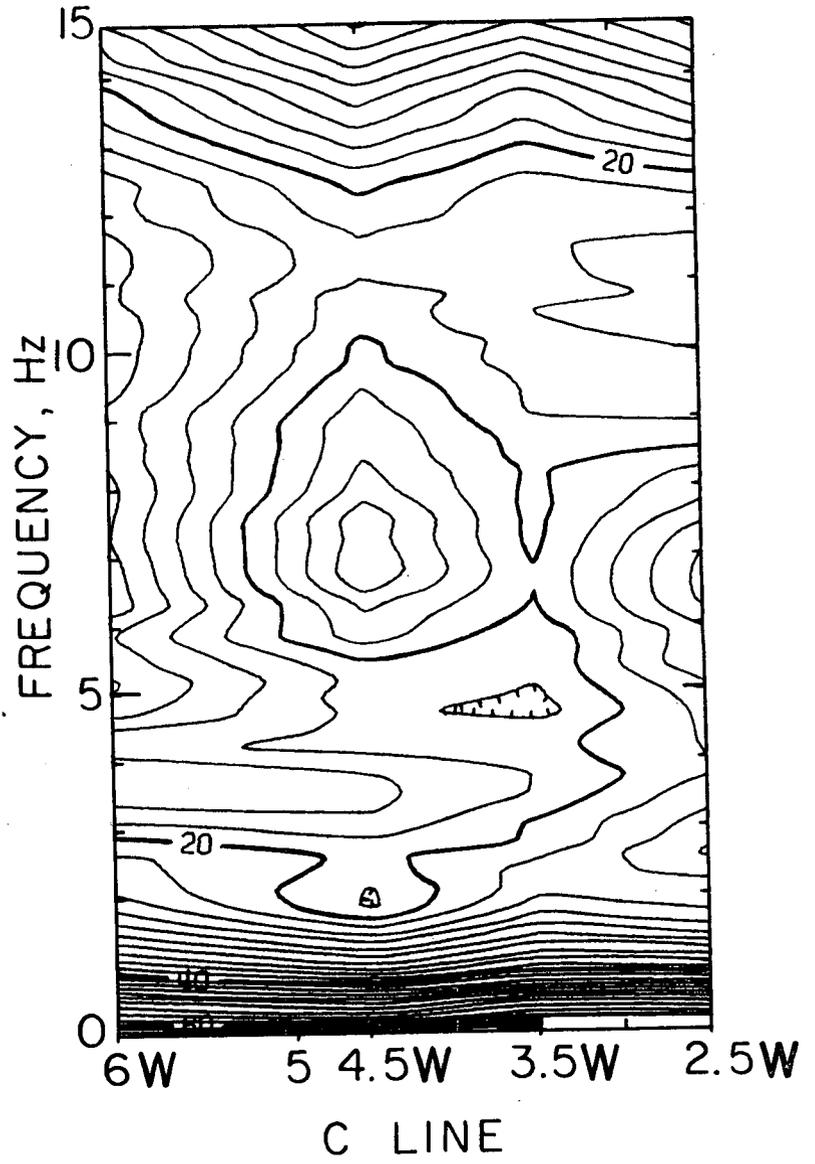
XBL 771-4995

Figure 4.7



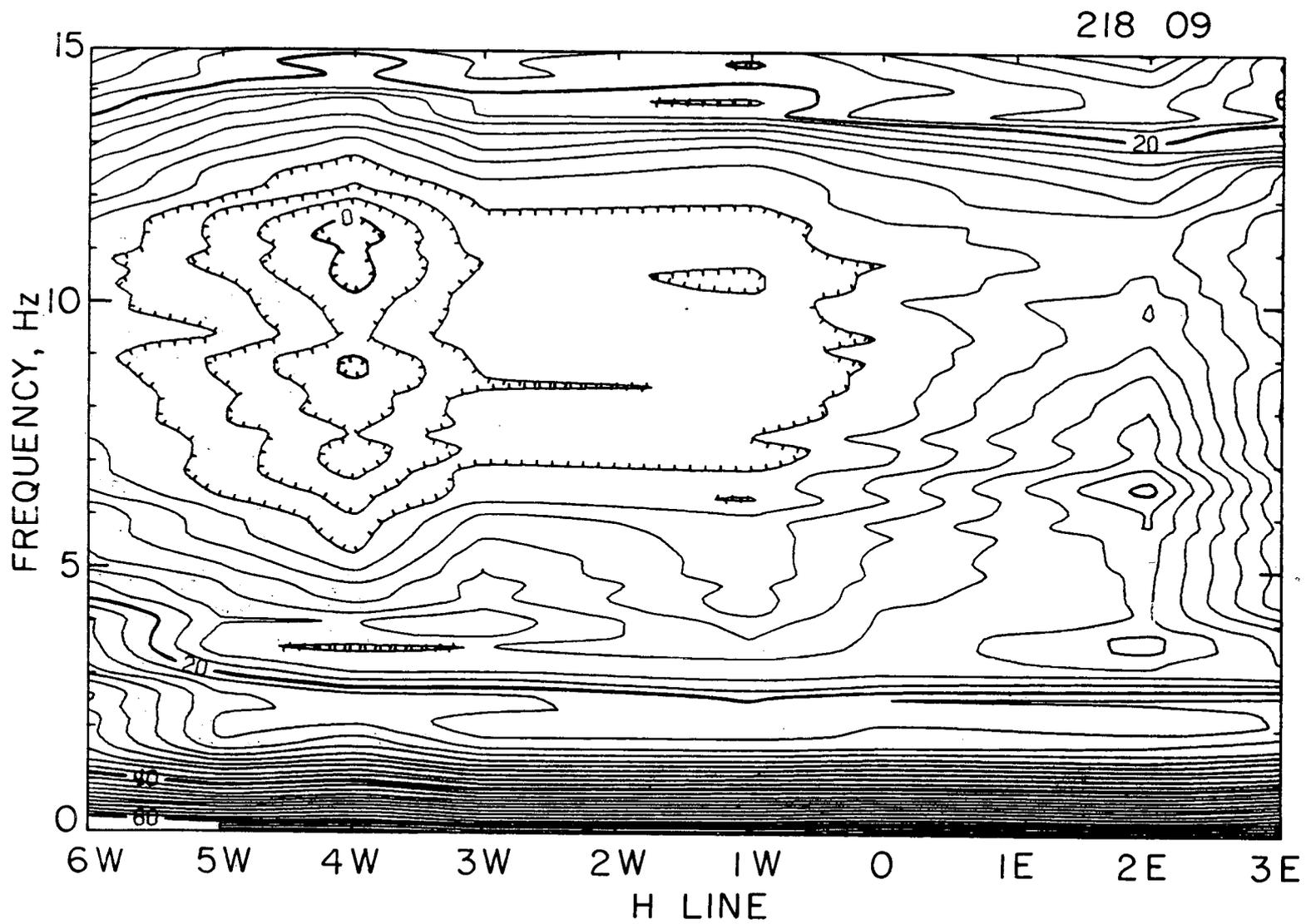
XBL 771-4997

Figure 4.8



XBL 771-4996

Figure 4.9



-63-

XBL 771-4998

Figure 4.10

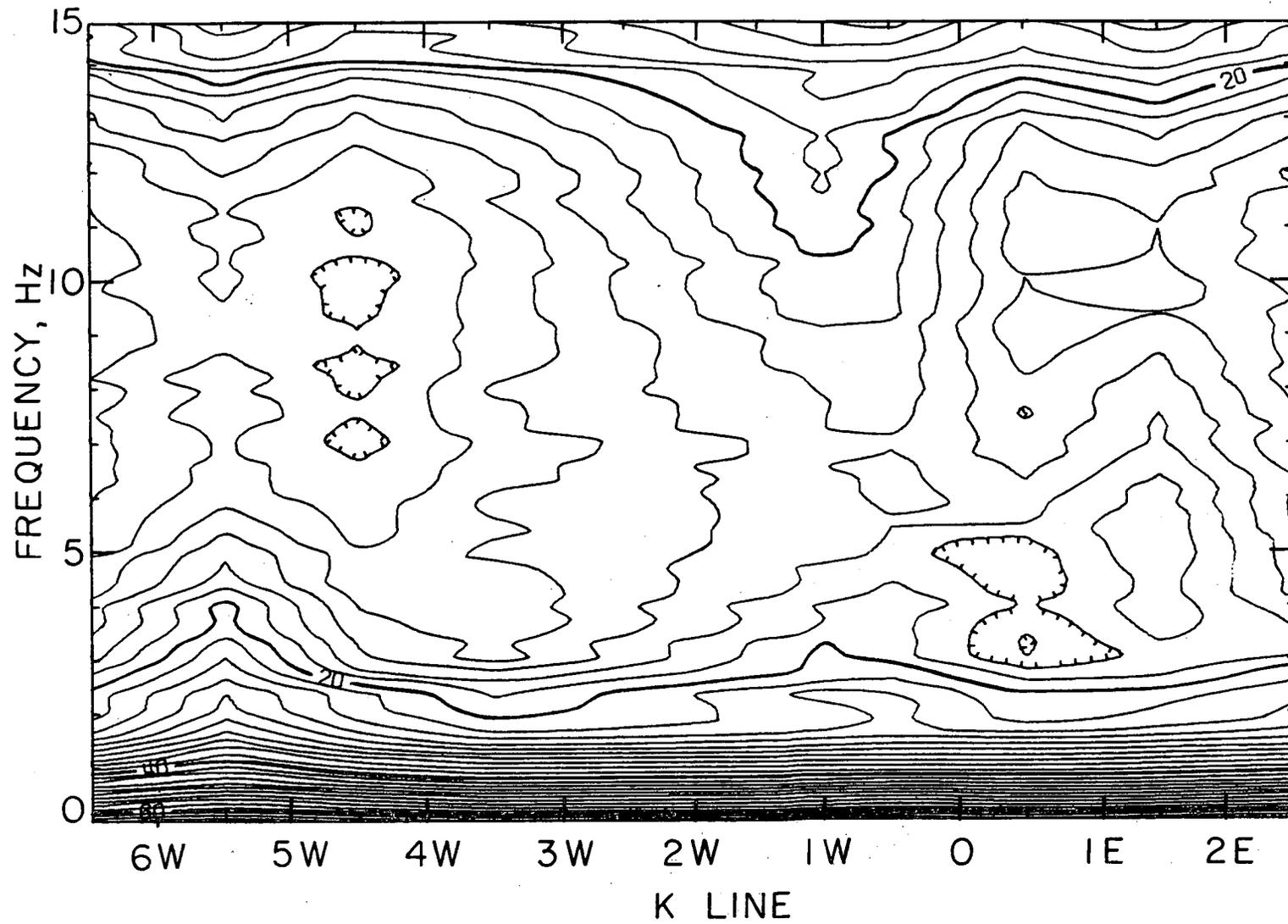


Figure 4.11

XBL 771-4999

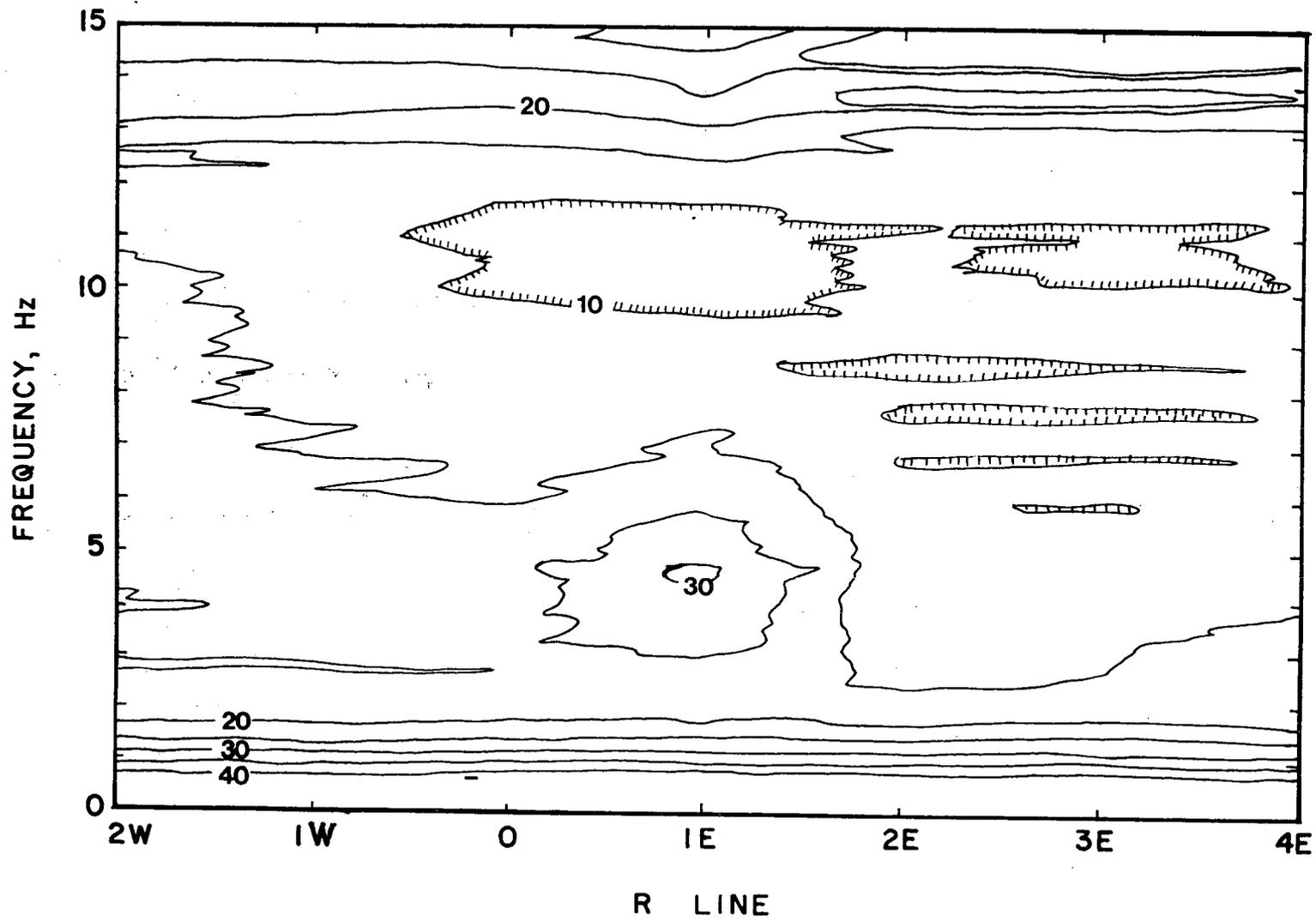


Figure 4.12

XBL 779-2472

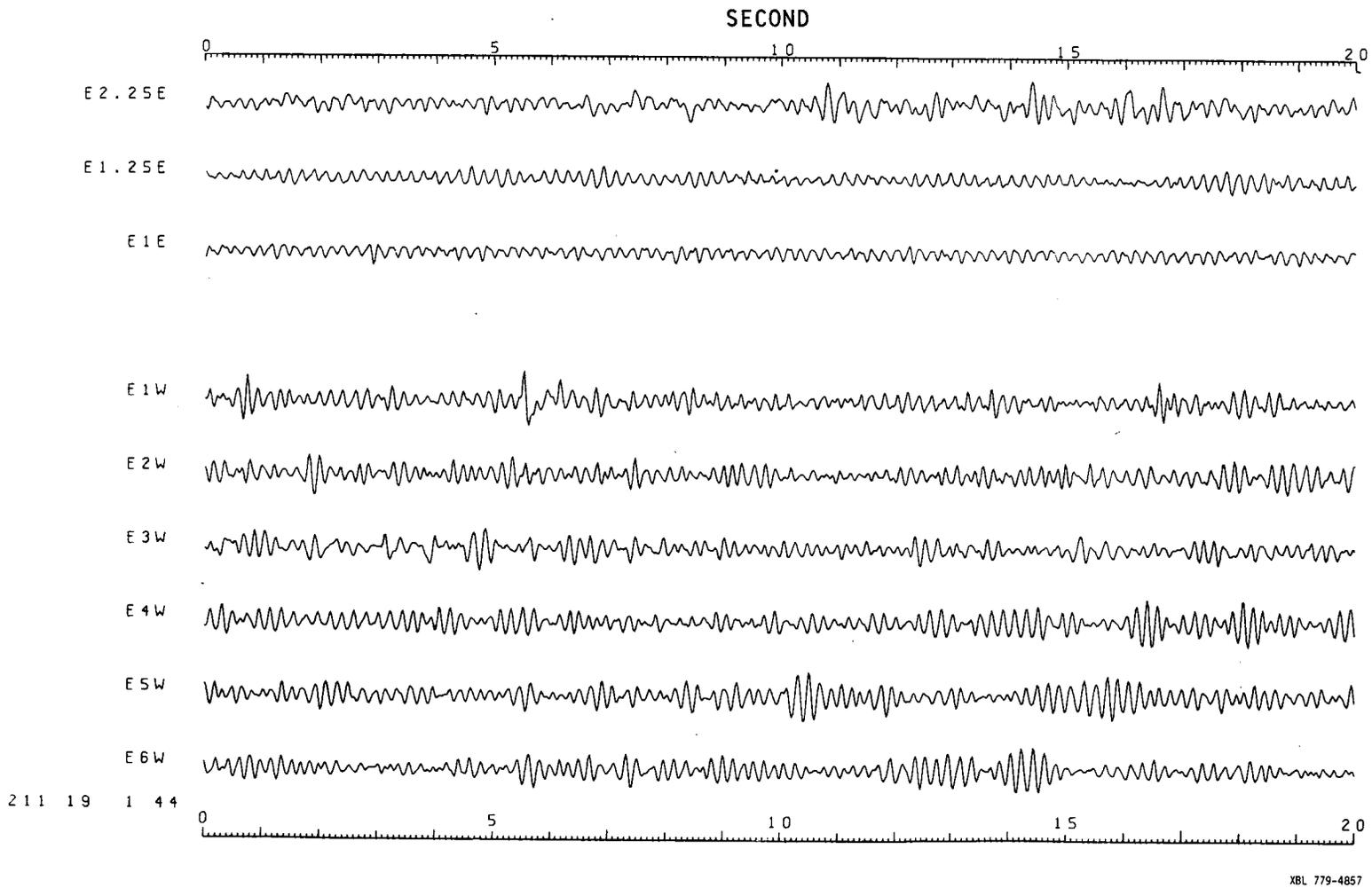


Figure 4.13

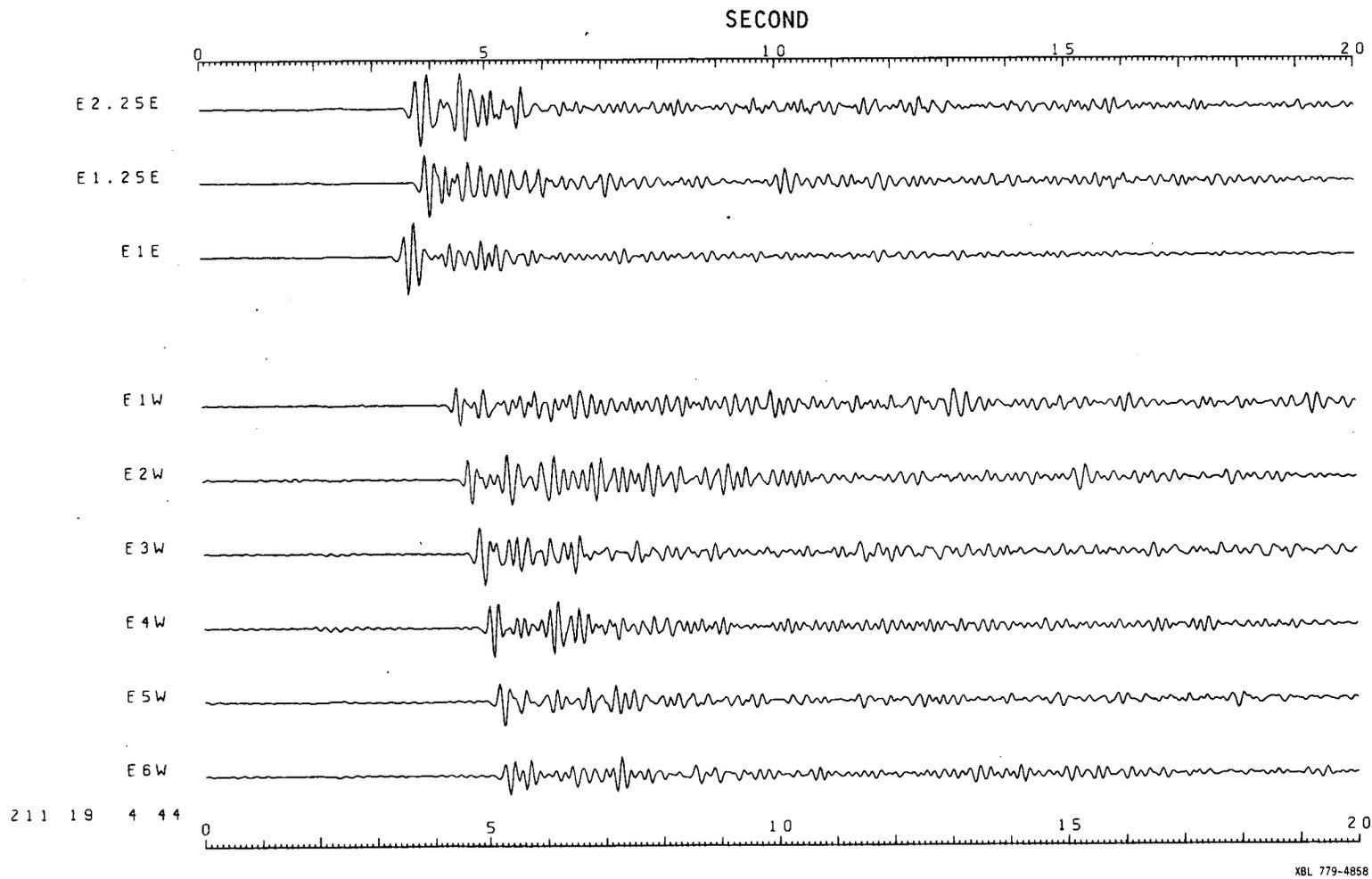
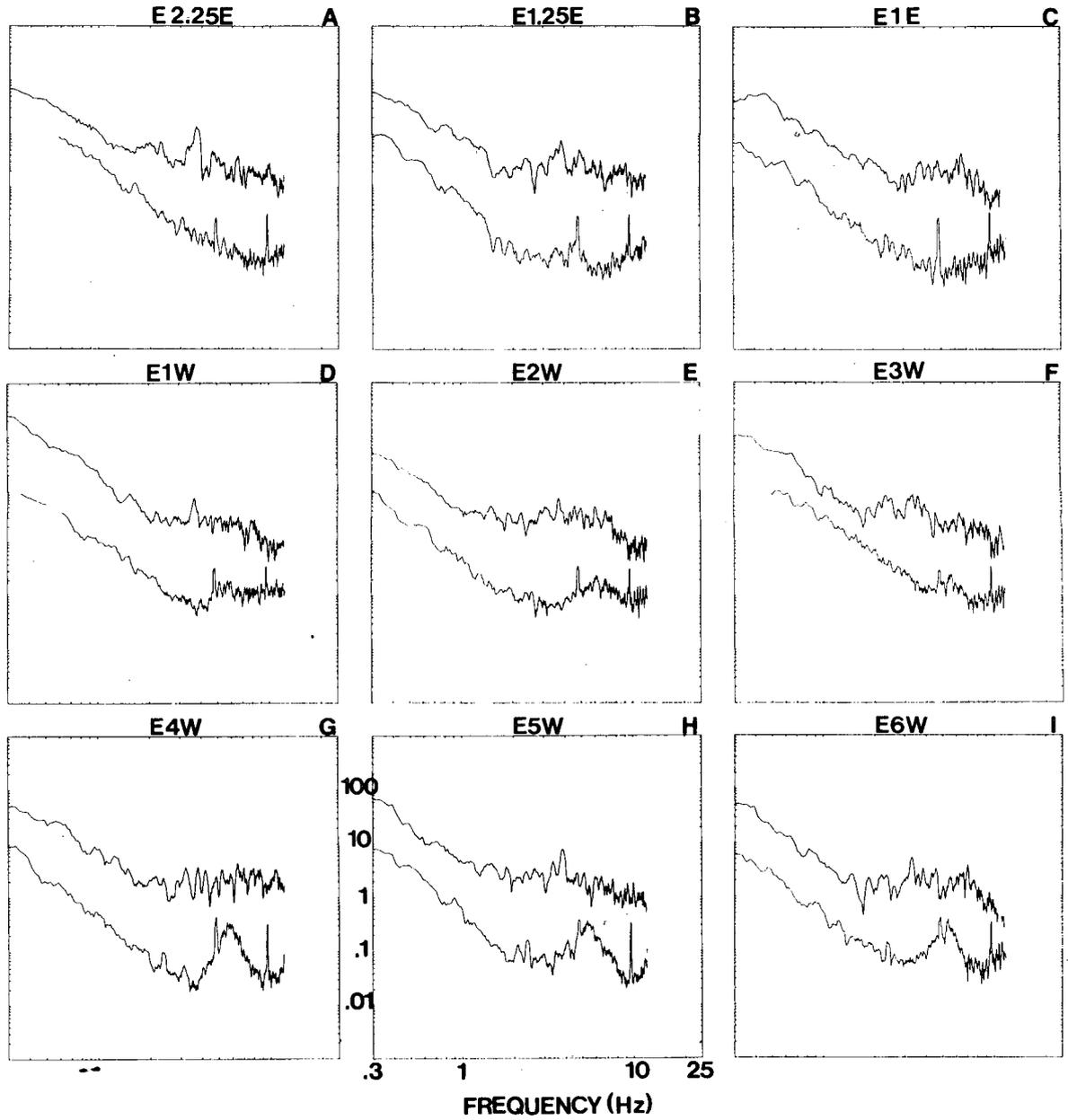


Figure 4.14



XBL 778-9828

Figure 4.15

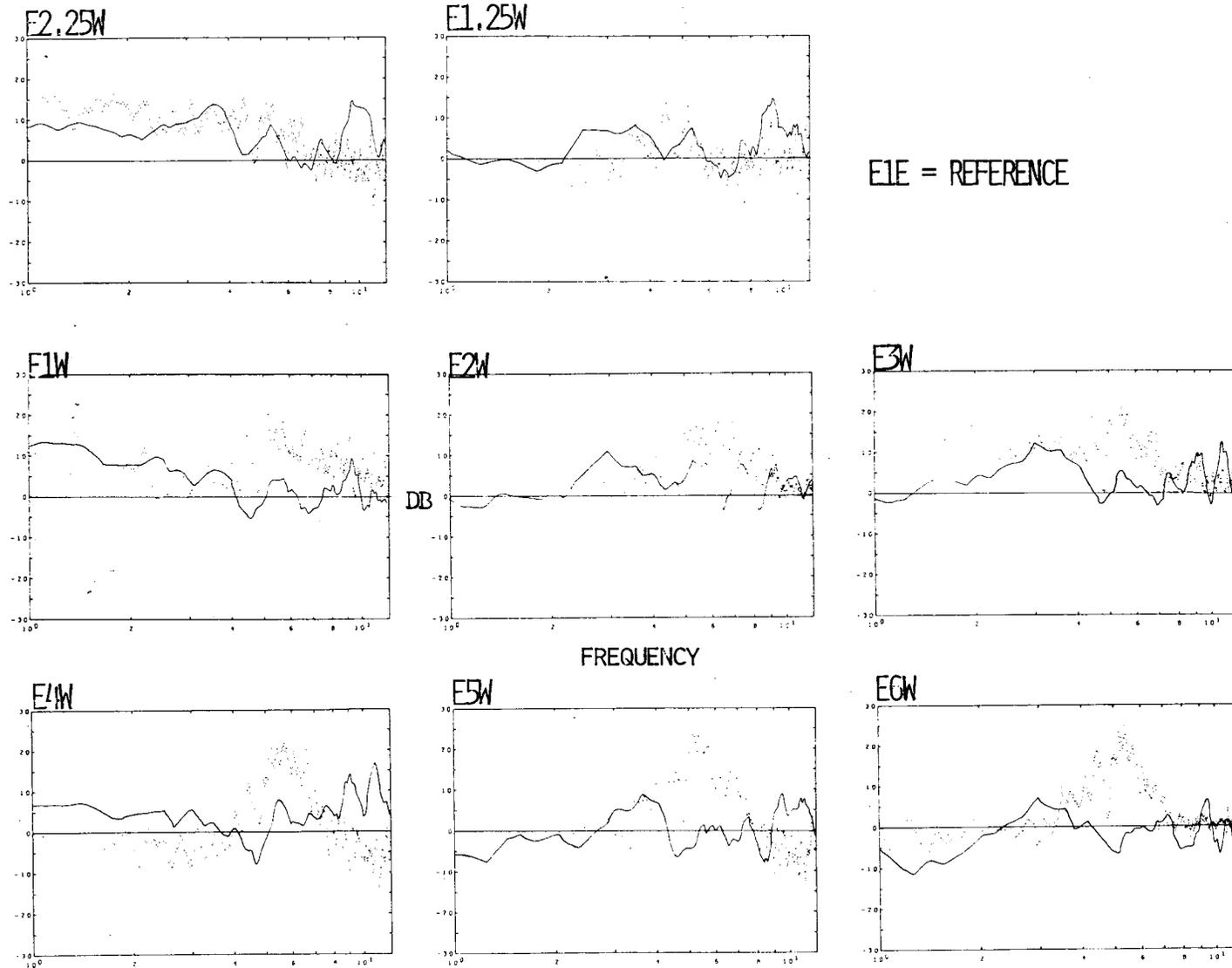
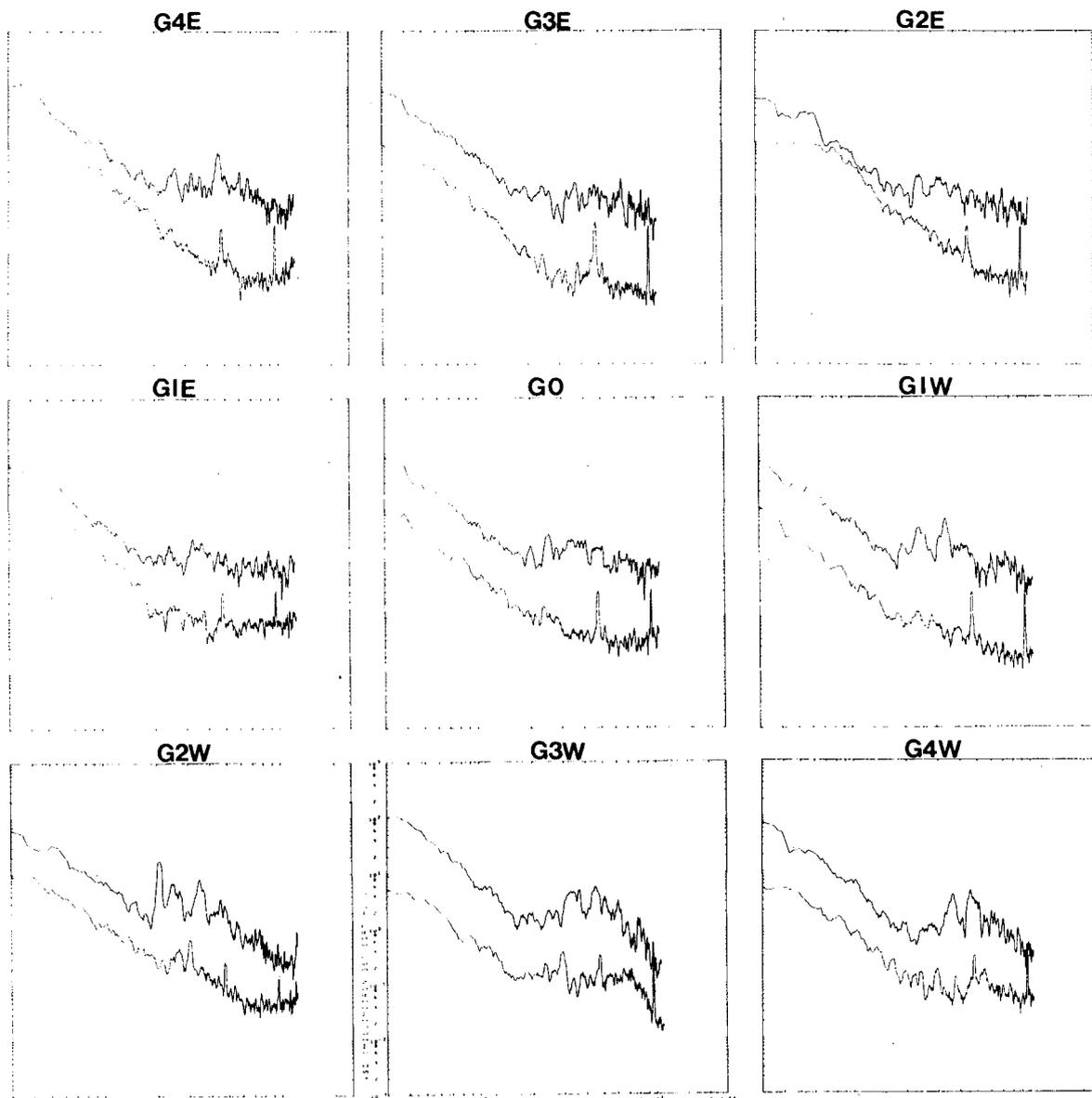


Figure 4.16



XBL 779-2473

Figure 4.17

## V. PROPAGATION CHARACTERISTICS OF GROUND NOISE

### 5.1 Introduction

We have shown in Chapter 4 that conventional noise power mapping techniques provide little information to make it possible to differentiate the effects of shallow geology from those due to buried sources. Seismic waves emitted from a buried reservoir source should be amenable to detection through local array measurements, because such vertically incident body waves have high apparent phase velocity and time-invariant propagation direction. The measurement of apparent phase velocity and azimuth is complicated for microseisms by multipath arrivals of both body waves and ambient surface waves. The reliable estimation of these propagation parameters therefore requires a properly designed array in conjunction with frequency-wavenumber (f-k) analysis. The horizontal propagating surface waves will appear in the f-k diagram as low velocity waves crossing the array, differing from the vertically incident body waves. The accurate estimation of the propagation parameters can provide the added data necessary to interpret the ambiguous noise power anomaly obtained by the method described in the previous chapter.

This chapter opens with a coherence study of seismic noise in Grass Valley conducted in order to design the array. The array design and the field program are presented in the following section. To my knowledge this work represents a first attempt at mapping the propagation parameters of microseisms in a geothermal area by using a non-aliasing roving array and advanced array processing techniques. The frequency-wavenumber (f-k) processing schemes used are the frequency domain

beam-forming method, BFM (Lacoss et al., 1969), and the maximum-likelihood method, MLM (Capon, 1969). We compare the two array processing techniques in terms of their response to identical input data. The f-k analyses of Grass Valley data are presented in the last two sections.

## 5.2 Coherence of ground noise

To design a full-scale ground noise survey, we first deployed a pilot 12-element L-shaped array on alluvium and at the hot springs to determine the coherence properties of the noise field as a function of sensor separation.

The coherence function estimate,  $\hat{R}_{mn}(f)$ , between signals m and n is given by

$$\hat{R}_{mn}(f) = \frac{|\hat{S}_{mn}(f)|}{\sqrt{\hat{S}_{mm}(f)\hat{S}_{nn}(f)}} \quad , \quad (5.1)$$

where  $\hat{S}_{mn}(f)$  is the estimate of cross-power spectral density.  $\hat{S}_{mn}(f)$  and  $\hat{S}_{nn}(f)$  are estimates of auto-power spectral density. This was defined in Equation (4.9). The 100% confidence interval for coherence estimates in the range of  $0.59 \leq R_{mn}(f) \leq 0.97$  with  $2I \geq 40$  degrees of freedom is calculated, according to Equation (6.111) of Bendat and Piersol (1971), by

$$\begin{aligned} & [\tanh\{g(f) - (2I - 2)^{-1} - \sigma_g Z_{\gamma/2}\}]^{1/2} < R_{mn}(f) \\ & [\tanh\{g(f) - (2I - 2)^{-1} + \sigma_g Z_{\gamma/2}\}]^{1/2} \quad , \end{aligned} \quad (5.2)$$

where  $Z_{\gamma/2}$  is a standard normal variate and

$$g(f) = \frac{1}{2} \ln \left[ \frac{1 + \hat{R}_{mn}(f)}{1 - \hat{R}_{mn}(f)} \right] = \tanh^{-1} \hat{R}_{mn}(f) , \quad (5.3)$$

and I = number of data blocks being averaged, and

$$\sigma_g = \sqrt{\frac{2I}{2I - 2}} \quad (5.4)$$

with 2I degrees of freedom.

Coherence of ground noise in Grass Valley decreases as the geophone spacing increases, the relation between coherence and geophone spacing varying from site to site. In the vicinity of Leach Hot Springs, high coherence between two geophones 10 m apart is seen in the frequency bands of 2 to 8 Hz and 10 to 20 Hz (Figure 5.1). As geophone spacing increases to 30 m, 40 m, and 60 m, the coherence decreases at high and low frequencies. At 60 m separation we can still observe high coherence in the 5-7 Hz band. There seems no obvious correlation between frequency bands with high coherence and the VSD (Figure 5.1E). In the valley, for example, at the intersection of line E and line M (EM), the frequency band of maximum coherence is 2 to 5 Hz when geophones are separated by 20 m (Figure 5.2). The coherence level and the width of the coherent frequency band decrease with increased geophone spacing to 40 m-60 m separation. The estimated coherence at nearby site E2.75W indicates some degree of coherence at 3 to 6 Hz, even at 120 m separation (Figure 5.2E). The degree of coherence seen, while variable, is generally low at valley sites, presumably the result of attenuation effects and interference from multi-path arrivals. Phase velocity and wavelength can be measured for the coherent part. Wavelengths as short as 10-20 m are present at the valley sites.

### 5.3 Roving array experiment

A small-aperture 12-geophone array was used. The array configuration and its impulse response in wavenumber space are shown in Figure 5.3. The existence of short wavelength noise components and the low coherence seen at large geophone separation, both dictated the tight array spacing used. An array of 100 m element separation or more, commonly used in ground noise studies elsewhere, would give spurious results because spatial aliasing folds the high-wavenumber noise components (which we have seen dominant in the valley alluvium) into low-wavenumber noise components. The spatial aliasing results in the detection of erroneously high-velocity microseisms, which are interpreted as body waves. We illustrate the effect of spatial aliasing due to inadequate element separation in Figure 5.4, where we processed a simulated 4 Hz plane wave with 50 m wavelength, propagating with phase velocity of 200 m/sec across four arrays. Those arrays have identical array shapes and numbers of sensors but different sensor spacing. The diameters of the arrays are 50 m, 75 m, 250 m, and 500 m, such that the sensor spacing for each array is proportional to the array size. Since the plane waves are propagating toward the azimuth of  $60^\circ$ , the folding effects are evident along the directions of  $60^\circ$  and  $240^\circ$ . Many interpretations of microseisms as body waves, based on coarse sensor separation, may well be incorrectly based on aliased low-velocity surface waves, as seen in Figure 5.4C. It is true, of course, that when the array is made small enough to accommodate the short-wavelength noise components, resolution for near-vertically incident body waves is seriously degraded, though the evidence of body waves should be visible and could be studied by appropriate array

expansion and spatial filtering.

To map the propagation parameters of the microseismic field, this 12-element array was placed each evening at a site. Sixteen sites in the region have been occupied. Data were transmitted by cable to the recording vehicle some 500 m from the array.

#### 5.4 Frequency-wavenumber power spectral density estimation

##### 5.4.1 Definition

The frequency-wavenumber power spectral density function (FKPSD) was introduced into seismology by Burg (1964) in the development of the optimum three-dimensional filter derived from the Wiener multichannel theory. Burg (1964) illustrated the optimum three-dimensional filter by a theoretical problem of P-wave enhancement in the presence of ambient Rayleigh waves. The FKPSD, a three-dimensional equivalent of ordinary spectral density function, is given by the relation

$$P(f, \underline{k}) = \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} c(\tau, \underline{\rho}) \exp[-j2\pi(f\tau + \underline{k} \cdot \underline{\rho})] d\tau d\rho_x d\rho_y, \quad (5.5)$$

where  $c(\tau, \underline{\rho})$  is the correlation function with time delay  $\tau$  and spatial lag  $\underline{\rho}$ ,  $\underline{k}$  is the vector wavenumber in cycles/km.

The correlation function of the noise field is defined as

$$c(\tau, \underline{\rho}) = E[\phi(t, \underline{r}) \cdot \phi(t+\tau, \underline{r}+\underline{\rho})], \quad (5.6)$$

where  $E$  denotes the averaged value, or expectation, and  $\phi(t, \underline{r})$  is the time series of the noise field at seismometer locations  $\underline{r} = (x, y)$ .

The array processing is normally carried out in the frequency domain rather than in the time domain because of computation time considerations.

With carefully designed filter coefficients, however, the time domain operation has the advantage of better resolution than the frequency domain operation (Capon, et al., 1967 and Lacoss, et al., 1969).

In the following sections, I compare the techniques for estimating FKPSD using the conventional and maximum-likelihood method in the frequency domain.

#### 5.4.2 Conventional method (BFM)

The conventional method, commonly known as the frequency domain beam forming method (BFM), estimates the FKPSD,  $P(f, \underline{k})$ , by the relation

$$\hat{P}(f, \underline{k}) = \frac{1}{N^2} \underline{a}' \cdot \underline{\hat{S}} \cdot \underline{a}, \quad (5.7)$$

where  $N$  is the number of geophones in the array and  $\underline{a}'$ , the conjugate transpose of  $\underline{a}$ , is the row vector

$$\underline{a}' = (e^{j2\pi \underline{k} \cdot \underline{r}_1}, e^{j2\pi \underline{k} \cdot \underline{r}_2}, e^{j2\pi \underline{k} \cdot \underline{r}_3}, \dots, e^{j2\pi \underline{k} \cdot \underline{r}_N}), \quad (5.8)$$

where  $\underline{r}_n$  is the coordinate of  $n^{\text{th}}$  geophone location.  $\underline{\hat{S}}$  is the estimate of the spectral matrix between sensors. Each entry of  $\hat{S}$ ,  $\hat{S}_{mn}(f)$ , is obtained from the averaged cross-power,

$$\bar{S}_{mn}(f) = \frac{1}{I} \sum_{i=1}^I \phi_m^i(f) \phi_n^{i*}(f), \quad (5.9)$$

by the normalization

$$\hat{S}_{mn}(f) = \frac{\bar{S}_{mn}(f)}{\sqrt{\bar{S}_{mm}(f) \bar{S}_{nn}(f)}}, \quad (5.9A)$$

where  $\phi_n^i(f)$  is the Fourier transform of block noise sample  $i$ , from seismometer  $n$  located at  $\underline{r}_n$ .

The BFM estimate of FKPSD, by re-arranging Equation (5.7), is

$$\hat{P}(f, \underline{k}) = \frac{1}{N^2} \sum_{m=1}^N \sum_{n=1}^N W_m W_n^* \hat{S}_{mn}(f) \exp[-j2\pi \underline{k} \cdot (\underline{r}_m - \underline{r}_n)] , \quad (5.10)$$

or

$$\hat{P}(f, \underline{k}) = \frac{1}{I} \sum_{i=1}^I \left| \frac{1}{N} \sum_{n=1}^N W_n^i \phi_n(f) \exp[-j2\pi \underline{k} \cdot \underline{r}_n] \right|^2 . \quad (5.11)$$

The term  $\exp(-j2\pi \underline{k} \cdot \underline{r}_n)$  in Equation (5.11) has the effect of advancing the phase of sinusoid observed at  $\underline{r}_n$  by the amount of the time delay with respect to the origin of the array assuming a plane wave propagating toward the azimuth  $\underline{k}$  with phase velocity  $\underline{V}$ .  $\underline{V}$  is given by

$$\underline{V} = \frac{f}{|\underline{k}|} \text{ km/sec} . \quad (5.12)$$

Taking  $W_m = W_n = 1$ , the BFM applies uniform weighting to each array element before the delay-and-sum operation. The BFM is efficient in computational time and provides an accurate estimate of azimuth and phase velocity if the noise field has high signal-to-noise ratio in a unique direction and represents a single mode of wave propagation. On the other hand, in the presence of multi-path propagation, the result in wavenumber space always shows an ambiguous pattern of peaks due to smearing of the true spectrum. The big side-lobes in the impulse response of the array (Figure 5.3) cause serious leakage in estimating spectral density.

Statistical properties of the estimator  $\hat{P}(f, \underline{k})$  were given by Capon and Goodman (1970). The a priori assumptions in deriving the probability

distribution are the following:

1) The sensor outputs comprise a multi-dimensional Gaussian random process with zero mean and stationary discrete time series, and

2) The length of the segments employed is long so that  $\hat{\phi}_m(f)$  is statistically independent of  $\hat{\phi}_n(f)$ ,  $m \neq n$ , then  $\zeta_1, \zeta_2, \zeta_3, \dots, \zeta_I$  are  $I$  independent and identically distributed  $N$ -variate complex Gaussian random variables, where  $\zeta_i$  is a column vector defined as

$$\zeta_i = \text{Col}[\phi_{1i}(f), \phi_{2i}(f), \phi_{3i}(f), \dots, \phi_{Ni}(f)] , \quad (5.13)$$

where  $\phi_{Ni}(f)$  is the Fourier transform of data in the  $i^{\text{th}}$  segment,  $n^{\text{th}}$  channel, frequency  $f$ .

The  $N \times N$  matrix-valued random variables,  $\hat{\underline{S}}(f)$  defined in Equation (5.9), have a complex Wishart distribution of dimension  $N$  and  $2I$  degree of freedom (when  $|\underline{k}_0| \neq 0$ ). The random variable  $\hat{P}(f, \underline{k})$  thus is a multiple of a Chi-square variable with mean and variance given by

$$E\{\hat{P}(f, \underline{k})\} = \int_{-f_N}^{f_N} \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} P(x, \underline{k}) |B(\underline{k} - \underline{k}_0) W_L(x - f_0)|^2 dx dk_x dk_y , \quad (5.14)$$

$$\text{var}\{\hat{P}(f, \underline{k})\} = \begin{cases} \frac{1}{I} [E\{\hat{P}(f_0, \underline{k}_0)\}^2], & f_0 \neq 0, f_N \\ \frac{2}{I} [E\{\hat{P}(f_0, \underline{k}_0)\}^2], & f_0 = 0, f_N \end{cases} , \quad (5.15)$$

where  $f_N$  is the Nyquist frequency and  $P(f, \underline{k})$  is the frequency-wavenumber power spectrum.  $|B(\underline{k})|^2$  is the array response function

$$B(\underline{k}) = \frac{1}{N} \sum_{n=1}^N \exp(j\underline{k} \cdot \underline{r}_n) . \quad (5.16)$$

$|W_L(x)|^2$  is the frequency window function

$$|W_L(x)|^2 = \frac{1}{L} \left| \frac{\sin(L/2)x}{\sin(1/2)x} \right|^2. \quad (5.17)$$

$E\{\hat{P}(f, \underline{k}_0)\}$  is obtained by means of a frequency-wavenumber window  $|B(\underline{k}-\underline{k}_0)|^2 \cdot |W_L(x-f)|^2$ . Hence  $\hat{P}(f, \underline{k})$  will be an asymptotically unbiased estimate of  $CP(f, \underline{k})$  as

$$\int_{-\pi}^{\pi} \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} |W_L(x-f) \cdot B(\underline{k}-\underline{k}_0)|^2 dx dk_x dk_y = C. \quad (5.18)$$

The var  $\{\hat{P}(f, \underline{k}_0)\}$  approaches zero as  $I$  approaches infinity such that  $\hat{P}(f, \underline{k})$  is a consistent estimate for  $CP(f, \underline{k})$ . The stability of the estimate is independent of the FKPSD, seismometer locations, or the weights  $W_n$ ,  $n = 1, \dots, N$  (Iacoss, et al., 1969) and is given by

$$v = \frac{2E\{\hat{P}(f_0, \underline{k}_0)\}^2}{\text{var}\{\hat{P}(f_0, \underline{k}_0)\}} \cdot b_1 = \begin{cases} 2Ib_1 & \text{for } f_0 \neq 0 \text{ or } f_N \\ Ib_1 & \text{for } f_0 = 0 \text{ or } f_N \end{cases}, \quad (5.19)$$

where  $b_1$  is dependent on the spectral window defined in Equation (4.12) and the  $100\gamma$  percent confidence interval for  $\hat{P}(f_0, \underline{k}_0)$  is

$$\frac{v\hat{P}(f_0, \underline{k}_0)}{\chi_v^2 \left(\frac{1+\gamma}{2}\right)} < P(f_0, \underline{k}_0) < \frac{v\hat{P}(f_0, \underline{k}_0)}{\chi_v^2 \left(\frac{1-\gamma}{2}\right)}. \quad (5.20)$$

For 24 data blocks, each block tapered by the rectangular data window,  $v = 48$ , the 90% confidence limits are about 1.58 dB above and -1.24 dB below  $\hat{P}(f_0, \underline{k}_0)$  for  $f_0 \neq 0$  or  $f_N$ .

### 5.4.3 Maximum-likelihood method (MLM)

The maximum-likelihood estimate for FKPSD is based on the application of optimal weighting functions which correspond to a maximum-likelihood filter in 2-dimensional wavenumber space to control the shape of the frequency-wavenumber window function, i.e., the beam pattern of the array impulse response. The construction of the maximum-likelihood filter is based on the coherence characteristics of the data among array sensors. These optimal weighting functions, when applied to the output of each array element, result in the maximum signal-to-noise ratio in the array signal estimation. At a selected frequency component, the maximum-likelihood filter is able to pass undistorted a monochromatic plane wave traveling at a velocity corresponding to a steering wavenumber,  $\underline{k}_0$ , and to suppress in an optimal least squares sense the power of those waves traveling at velocities corresponding to wavenumbers other than  $\underline{k}_0$ . The weighting function changes as the spectrum changes.

The maximum-likelihood estimate for FKPSD can be written as

$$\tilde{P}(f, \underline{k}) = \sum_{m=1}^N \sum_{n=1}^N A_m(f) A_n^*(f) \hat{S}_{mn}(f) \exp[-j2\pi(\underline{k} \cdot \underline{\rho})] \quad , \quad (5.21)$$

where  $A_m(f)$  and  $A_n(f)$  are weights applied to the outputs of array elements  $m$  and  $n$  respectively. Note that  $A_m(f)$  and  $A_n(f)$  are functions of frequency, wavenumber, and spatial coordinates of the array element.

To consider the optimal weighting function in the "distortionless" sense and with optimal suppression of noise, we rearrange Equations (5.21) into the form

$$\tilde{P}(f, \underline{k}) = \sum_{m=1}^N \sum_{n=1}^N A_m(f) A_n^*(f) \tilde{S}_{mn}(f) \quad , \quad (5.22)$$

where

$$\tilde{S}_{mn}(f) = \hat{S}_{mn}(f) \exp(-j2\pi \underline{k} \cdot \underline{x}_m) \exp(j2\pi \underline{k} \cdot \underline{x}_n) , \quad (5.23)$$

is the spectral matrix with delays.

Define a weighting vector

$$\underline{W} = \text{col}[A_1(f), A_2(f), \dots, A_n(f)] . \quad (5.24)$$

Equation (5.22) becomes

$$\tilde{P}(f, \underline{k}) = \underline{W}' \cdot \underline{\tilde{S}} \cdot \underline{W} , \quad (5.25)$$

where  $\underline{W}'$  is the conjugate transpose of  $\underline{W}$ .

In this optimization problem, we desire minimum array output power for  $\underline{k} \neq \underline{k}_0$ , which is equivalent to minimizing the quantity  $(\underline{W}' \cdot \underline{\tilde{S}} \cdot \underline{W})$ . The constraint of distortionless filter response at  $\underline{k} = \underline{k}_0$  requires the sum of all  $N$  coefficients to be unity over the narrow wavenumber band around  $\underline{k} = \underline{k}_0$ , i.e.

$$\sum_{n=1}^N A_n(f, \underline{k}_0) = 1 \quad (5.26)$$

or

$$\underline{C}^T \cdot \underline{W} = 1 , \quad (5.27)$$

where  $\underline{C} = \text{col}[1, 1, \dots, 1]$ . (5.28)

The problem of finding the optimum set of filter weights,  $\underline{W}_{\text{opt}}$ , is summarized by Equations (5.25) and (5.27) as

$$\text{minimize } (\underline{W}' \cdot \underline{\tilde{S}} \cdot \underline{W}) \text{ with respect to } \underline{W} \quad (5.29)$$

$$\text{subject to } \underline{C}^T \cdot \underline{W} = 1 . \quad (5.30)$$

Using the method of Lagrange multipliers in the calculus of variations,

we set up the equation as:

$$H(\underline{W}) = \frac{1}{2} \underline{W}' \cdot \underline{\tilde{S}} \cdot \underline{W} + \lambda (\underline{C}^T \underline{W} - 1) , \quad (5.31)$$

where the coefficient 1/2 is introduced to simplify later arithmetic and  $\lambda$  is an undetermined Lagrange multiplier.

Taking the gradient of Equation (5.31) with respect to  $\underline{W}$  and equating to zero,

$$\nabla_{\underline{W}} H(\underline{W}) = \underline{\tilde{S}} \cdot \underline{W} + \underline{C}\lambda = 0 \quad (5.32)$$

yielding

$$\underline{W} = -\underline{\tilde{S}}^{-1} \underline{C}\lambda , \quad (5.33)$$

substituting (5.33) into (5.30) we obtain

$$\lambda = -(\underline{C}^T \underline{\tilde{S}}^{-1} \underline{C})^{-1} . \quad (5.34)$$

From Equations (5.33) and (5.34), the optimal weighting vector is obtained as

$$\underline{W}_{\text{opt}} = (\underline{S}^{-1} \cdot \underline{C}) (\underline{C}^T \cdot \underline{\tilde{S}}^{-1} \cdot \underline{C})^{-1} , \quad (5.35)$$

and the optimal weighting functions are:

$$A_n(f) = \frac{\sum_{m=1}^N \tilde{q}_{mn}(f)}{\sum_{m=1}^N \sum_{n=1}^N \tilde{q}_{mn}(f)} , \quad n = 1, \dots, N , \quad (5.36)$$

where the matrix  $\{\tilde{q}_{mn}(f)\}$  is the inverse of the delayed spectral matrix,  $\{\tilde{S}_{mn}(f)\}^{-1} = \underline{\tilde{S}}^{-1}$ , as defined in Equation (5.23).

The filter obtained from the above optimization procedures is a distortionless form of a constrained least squares filter, or constrained

LMS filter (Frost, 1972). The right hand side of Equation (5.36) is identical to the coefficients of the maximum-likelihood filter derived by Capon (1973) using the maximum-likelihood estimation technique.

Substituting Equation (5.35) into Equation (5.25), the estimate for FKPSD is:

$$\tilde{P}(f, \underline{k}) = \left( \frac{\underline{C}^T \cdot \underline{\tilde{S}}^{-1}}{\underline{C}^T \cdot \underline{\tilde{S}}^{-1} \cdot \underline{C}} \right) \cdot \underline{\tilde{S}} \cdot \left( \frac{\underline{\tilde{S}}^{-1} \cdot \underline{C}}{\underline{C}^T \cdot \underline{\tilde{S}}^{-1} \cdot \underline{C}} \right) \quad (5.37)$$

$$= (\underline{C}^T \cdot \underline{\tilde{S}}^{-1} \cdot \underline{C})^{-1} \quad (5.38)$$

$$= (\underline{a}' \cdot \underline{\hat{S}}^{-1} \cdot \underline{a})^{-1} .$$

This is just the high-resolution estimate for  $P(f, \underline{k})$  given by Capon(1969).

The justification for the usage of  $\tilde{P}(f, \underline{k})$  as an estimation for  $P(f, \underline{k})$  is provided by an important property of the maximum-likelihood estimator. When the maximum-likelihood estimator is used in signal estimation, it is identical to the minimum-variance unbiased estimator of the signal (Capon et al., 1967; and Capon, 1973). This is the consequence of minimizing the variance of the array output,  $(\underline{W}' \cdot \underline{\tilde{S}} \cdot \underline{W})$ , in Equation (5.29). Therefore,  $\tilde{P}(f, \underline{k})$  is a minimum-variance unbiased estimate for  $P(f, \underline{k})$ .

The statistical properties of  $\tilde{P}(f, \underline{k})$  have been discussed by Capon and Goodman (1970), who show that  $\tilde{P}(f, \underline{k})$  is a multiple of chi-square variable with  $2(I - N + 1)$  degrees of freedom, if  $f_0 = 0$ , or  $f_N$ . The mean and variance of  $\tilde{P}(f, \underline{k})$  are given by

$$E\{\tilde{P}(f_0, \underline{k}_0)\} = \left(\frac{I-N+1}{I}\right) \int_{-f_N}^{f_N} \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} P(x, \underline{k}) |W_L(x-f_0) \tilde{B}(f_0, \underline{k}; \underline{k}_0)|^2 dx dk_x dk_y \quad (5.39)$$

$$\text{var}\{\tilde{P}(f_0, \underline{k}_0)\} = \left(\frac{1}{I-N+1}\right) [E\{\tilde{P}(f_0, \underline{k}_0)\}]^2, \quad f_0 \neq 0 \text{ or } f_N \quad (5.40)$$

$$\text{or} \quad \left(\frac{2}{I-N+1}\right) [E\{\tilde{P}(f_0, \underline{k}_0)\}]^2, \quad f_0 = 0 \text{ or } f_N \quad (5.41)$$

The confidence limits for  $\tilde{P}(f_0, \underline{k}_0)$  can be obtained in a manner similar to that discussed for the BFM. If  $I = 24$ ,  $N = 12$ , and the rectangular taper is used, there are 26 degrees of freedom and the 90% confidence limits are 2.28 dB above and -1.76 dB below  $\tilde{P}(f_0, \underline{k}_0)$ , for  $f_0 \neq 0$  or  $f_N$ . In the article discussing the frequency domain beam forming method, Lacoss et al. (1969) have shown that as long as the requirement  $W_{nm} = W_{nm}^*$  is satisfied, the stability of any frequency domain beam forming estimate is independent of the seismometer location, FKPSD, or the weights  $W_n$ ,  $n = 1, \dots, N$ .  $W_{nm}$  represents  $W_m W_n$  and  $W_m$  is the weighting function applied on  $m^{\text{th}}$  sensor output. The requirement  $W_{nm} = W_{nm}^*$  ensures that  $\tilde{P}(f_0, \underline{k}_0)$  be real. Since the weighting function of MLM satisfies the above requirements, the confidence limits of the MLM estimate depend only on the number of data blocks and array sensors.

Equation (5.39) indicates that  $E\{P(f_0, \underline{k}_0)\}$  is obtained by means of a frequency-wavenumber window,  $|W_L(x-f_0) \cdot \tilde{B}(f_0, \underline{k}; \underline{k}_0)|^2$  where

$$\tilde{B}(f_0, \underline{k}; \underline{k}_0) = \sum_{n=1}^N A_n(f) \exp[j(\underline{k} - \underline{k}_0) \cdot \underline{r}_n]; \quad (5.42)$$

therefore  $\tilde{P}(f_0, \underline{k}_0)$  is an asymptotically unbiased estimator for  $CP(f_0, \underline{k}_0)$ , if the window approaches a three-dimensional delta function in such a way that

$$\left(\frac{I-N+1}{I}\right) \int_{-f_N}^{f_N} \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} |W_L(x-f_0) \tilde{B}(f_0, \underline{k}; \underline{k}_0)|^2 dx dk_x dk_y = C. \quad (5.43)$$

#### 5.4.4 Comparing the FKPSD estimation techniques

The term "high resolution f-k spectrum" is commonly used for maximum-likelihood (MLM) estimation of FKPSD, because, under favorable conditions, this method always results in a spectrum with sharper peaks and lower side-lobes than that estimated by conventional (BFM) methods. Based on theoretical analyses, however, Cox (1973) has found that the MLM has a disadvantage relative to the BFM in terms of its sensitivity to measurement errors, especially in the case of channel mismatch. Mismatch may result from distortion in the waveform during propagation, from amplitude, phase and position errors in the sensors, or in the sampling and digitization. We have examined the effects of mismatch by processing simulated data with both MLM and BFM. Input data are 4 Hz plane waves superimposed upon random noise of varying amplitude level. The 4 Hz plane wave propagates with phase velocity of 200 m/sec across an array in 50 m diameter. A block of data without random noise components is shown in Figure 5.5. We have found from the results of BFM that, regardless of the signal-to-noise ratios of input data, the f-k diagrams at 4 Hz are all identical to Figure 5.6A. This figure shows a shifted array response, centered around the wavenumber of the simulated signal. The peak value of f-k spectrum at 4 Hz is independent of the signal-to-noise ratios between 48.8 dB and 10.6 dB. On the other hand, the results of MLM analysis indicate that MLM is sensitive to signal-to-noise ratio of the input data. This effect is illustrated by the f-k plots for signal-to-noise ratio of 48.8 dB (Figure 5.6B), 30.6 dB (Figure 5.6C), 25.6 dB (Figure 5.6D), 20 dB (Figure 5.6E) and 10.6 db (Figure 5.6F). Those figures clearly indicate that MLM suppresses the side-lobes more efficiently and produces higher resolutions than BFM. The high

resolution effect is a distinct advantage in the case of multiple arrivals.

The desirable characteristics of MLM are results of the combination of the side-lobe suppression and the noise rejection. We shall discuss these operations by assuming that  $\underline{\hat{S}}$  consists of signal-plus-noise, so that we may express  $\underline{\hat{S}}$  as the form

$$\underline{\hat{S}} = \sigma_{\underline{\hat{Q}}}^2 + \sigma_1^2 \underline{d} \cdot \underline{d}' \quad , \quad (5.44)$$

where

- $\sigma_{\underline{\hat{Q}}}^2$  is the noise spectral matrix,
- $\underline{\hat{Q}}$  is normalized to have its trace equal to the number of sensors,
- $\sigma_0^2$  is the input noise spectral level averaged across the sensors,
- $\sigma_1^2$  is the averaged input power spectrum of the coherent microseismic signal,
- $\underline{d}$  is a directional vector of microseismic field, defined as

$$\underline{d} = \begin{bmatrix} \exp(j2\pi \underline{k}_0 \cdot \underline{x}_1) \\ \exp(j2\pi \underline{k}_0 \cdot \underline{x}_2) \\ \vdots \\ \exp(j2\pi \underline{k}_0 \cdot \underline{x}_N) \end{bmatrix} \quad (5.45)$$

$\underline{k}_0$  is a wavenumber vector in the direction to which the microseismic field is propagating.

We can also write a general expression for estimate FKPSD as

$$P(f, \underline{k}) = \underline{\tilde{W}}' \cdot \underline{\hat{S}} \cdot \underline{\tilde{W}} \quad (5.46)$$

where  $\underline{W}$  is the vector of generalized weighting functions,

$$\underline{\tilde{W}}' = \frac{1}{N} \underline{a}' \quad (5.47)$$

for BFM, and

$$\underline{\tilde{W}} = (\underline{\tilde{S}}^{-1} \cdot \underline{C})(\underline{C}^T \cdot \underline{\tilde{S}}^{-1} \cdot \underline{C})^{-1} = (\underline{\hat{S}} \cdot \underline{a})(\underline{a}' \cdot \underline{\hat{S}}' \cdot \underline{a})^{-1} \quad (5.48)$$

for MLM.

Substituting Equation (5.44) to Equation (5.46)

$$P(f, \underline{k}) = \sigma_{\underline{O}}^2 \underline{\tilde{W}}' \cdot \underline{\hat{Q}} \cdot \underline{\tilde{W}} + \sigma_1^2 |\underline{\tilde{W}}' \cdot \underline{d}|^2 \quad (5.49)$$

We define an inner product between two column vectors  $\underline{a}$  and  $\underline{d}$  by  $\underline{a}' \cdot \underline{C} \cdot \underline{d}$ , where  $\underline{C}$  is a positive definite Hermitian matrix, and the cosine-square of the generalized angle between  $\underline{a}$  and  $\underline{d}$  to be

$$\cos^2(\underline{a}, \underline{d}; \underline{C}) = |\underline{a}^* \cdot \underline{C} \cdot \underline{d}|^2 / \{(\underline{a}' \cdot \underline{C} \cdot \underline{a})(\underline{d}' \cdot \underline{C} \cdot \underline{d})\} \quad (5.50)$$

BFM. For BFM, the output spectrum can be written as

$$\hat{P}(f, \underline{k}; \underline{k}_0) = \frac{\sigma_{\underline{O}}^2}{N^2} (\underline{a}' \cdot \underline{\hat{Q}} \cdot \underline{a}) + \sigma_1^2 \cos^2(\underline{a}, \underline{d}; \underline{I}) \quad (5.51)$$

The first term is the noise response, the second term is the signal response, and  $\cos^2(\underline{a}, \underline{d}; \underline{I})$  is the array impulse response shifted by the steering vector  $\underline{a}$  to the "direction"  $\underline{d}$ , as shown in Figure 5.6A. It is evident that signal response depends on the cosine square of the generalized angle between  $\underline{a}$  and  $\underline{d}$  in inner product space and is not affected by the noise spectral matrix  $\underline{\hat{Q}}$ . Therefore, the signal response is relatively insensitive to small mismatch between  $\underline{a}$  and  $\underline{d}$ , as well as input signal-to-noise ratio. The noise response is reduced

by increase N which is the number of sensors in the array.

MLM. The output spectrum of MLM is obtained by substituting Equation (5.48) into Equation (5.49),

$$\begin{aligned} \tilde{P}(f, \underline{k}; \underline{k}_o) = & \frac{\sigma_o^2}{\underline{a}' \cdot \underline{\hat{Q}}^{-1} \cdot \underline{a}} \left\{ 1 + \frac{[2(S/N)_{\max} + (S/N)_{\max}^2] \sin^2(\underline{a}, \underline{d}; \underline{\hat{Q}}^{-1})}{[1 + (S/N)_{\max} \sin^2(\underline{a}, \underline{d}; \underline{\hat{Q}}^{-1})]^2} \right\} \\ & + \sigma_i^2 \left( \frac{\underline{d}' \cdot \underline{\hat{Q}}^{-1} \cdot \underline{d}}{\underline{a}' \cdot \underline{\hat{Q}}^{-1} \cdot \underline{a}} \right) \left( \frac{\cos^2(\underline{a}, \underline{d}; \underline{\hat{Q}}^{-1})}{\{1 + (S/N)_{\max} \sin^2(\underline{a}, \underline{d}; \underline{\hat{Q}}^{-1})\}^2} \right), \end{aligned} \quad (5.52)$$

where  $(S/N)_{\max}$  is the maximum output signal-to-noise spectral ratio defined as:

$$(S/N)_{\max} = (\underline{d}' \cdot \underline{\hat{Q}}^{-1} \cdot \underline{d}) \frac{\sigma_1^2}{\sigma_o^2}, \quad \text{and} \quad (5.53)$$

$$\sin^2(\underline{a}, \underline{d}; \underline{\hat{Q}}^{-1}) = 1 = \cos^2(\underline{a}, \underline{d}; \underline{\hat{Q}}^{-1}) \quad (5.54)$$

The expressions for signal and noise responses appearing in Equation (5.52) are more complicated than those of the BFM. It is evident from Equation (5.52) that both  $(S/N)_{\max}$  and  $\sin^2(\underline{a}, \underline{d}; \underline{\hat{Q}}^{-1})$  play important roles in signal expression and noise rejection. The suppression of spurious side-lobes due to finite dimension array response can be seen from the behavior of the second term in Equation (5.4). As  $\underline{a}$  deviates away from  $\underline{d}$ , the quantity of  $\cos^2(\underline{a}, \underline{d}; \underline{\hat{Q}}^{-1})$  becomes smaller while  $\sin^2(\underline{a}, \underline{d}; \underline{\hat{Q}}^{-1})$  becomes larger. The signal suppression can be significant at  $\underline{a} \neq \underline{d}$  when  $(S/N)_{\max}$  is larger than unity.

Let us turn our attention to the first term of Equation (5.52), i.e. the noise response. To examine the unusual nature of the noise response, we let  $\hat{\underline{Q}} = \underline{I}$ , such that the quantity  $(\underline{a}' \cdot \hat{\underline{Q}}^{-1} \cdot \underline{a})$  is a constant. Then as  $\sin^2(\underline{a}, \underline{d}; \underline{I})$  varies from zero to one, the noise response increases from the value  $\sigma_o^2/N$  at  $\sin^2(\underline{a}, \underline{d}; \underline{I}) = 0$ , until it reaches a maximum of  $(\frac{\sigma_o^2}{N}) \frac{[2+(S/N)_{\max}]^2}{4+4(S/N)_{\max}}$  at  $\sin^2(\underline{a}, \underline{d}; \underline{I}) = 1/(2+(S/N)_{\max})$ .

This peculiar noise response is due to the fact that the MLM estimator treats the mismatched signal as an unwanted interference and performs a compromise between suppressing the signal and rejecting the real noise. The stronger the mismatched signal, the more importance the estimator puts on suppressing it. In suppressing the mismatched signal it accepts a lesser rejection of noise. Near the point  $\sin^2(\underline{a}, \underline{d}; \underline{I}) = 0$ , the signal suppression is minimum. As  $\sin^2(\underline{a}, \underline{d}; \underline{I})$  increases, the greater suppression of the signal is possible with a corresponding increasing penalty in noise rejection. Eventually the mismatch reaches the point where the signal suppression is sufficient so that a further penalty in noise response is not justified. The processor then reverses the trend and places greater emphasis on rejecting the noise.

The MLM, however, is not an optimal array processor. The terms in large braces in Equation (5.52) deviate the output spectrum of MLM away from that of an optimal array processor. We will discuss the optimal array processor later in this section. Look back at Equation (5.36) where we used the cross-power spectral matrix of the signal-plus-noise to construct the weighting function for MLM.

The optimal array processor, on the other hand, uses the cross-power spectral

matrix of noise alone to construct the weighting functions in order to enhance the transient signal, such as earthquake or explosion arrivals. Unfortunately, such an optimal array processor is not applicable to microseismic data because of the difficulty in separating signal from noise in microseismic data. It appears that, at present, MLM is the best way to process two-dimensional array data for high resolution in the presence of multi-path interference, the normal situation in ground noise studies. It is worthwhile, nevertheless, to consider the responses of the optimal array processor.

Optimal Array Processor. The output spectrum given by Cox (1973) is

$$P(f, \underline{k}; \underline{k}_0) = \frac{\sigma_0^2}{(\underline{a}' \cdot \underline{\hat{Q}}^{-1} \cdot \underline{a})} + \sigma_1^2 \left( \frac{\underline{d}' \cdot \underline{\hat{Q}}^{-1} \cdot \underline{d}}{\underline{a}' \cdot \underline{\hat{Q}}^{-1} \cdot \underline{a}} \right) \cos^2(\underline{a}, \underline{d}; \underline{\hat{Q}}^{-1}) \quad (5.55)$$

which results from substituting  $\tilde{\underline{W}}' = (\underline{\hat{Q}}^{-1} \cdot \underline{a})(\underline{a}' \cdot \underline{\hat{Q}}^{-1} \cdot \underline{a})^{-1}$  into Equation (5.49). Again, the first term corresponds to the noise response and the second term corresponds to the signal. Notice that Equation (5.55) reduces to Equation (5.51) for the case of spatially uncorrelated noise,  $\underline{\hat{Q}} = \underline{I}$ , and  $\underline{a} = \underline{d}$ .

In order to realize the effects of optimization, one must understand the effect of the matrix  $\underline{\hat{Q}}^{-1}$ , since we have defined  $\underline{\hat{Q}}$  as a noise spectral matrix in Equation (5.44), the small eigenvalues of  $\underline{\hat{Q}}$  correspond to the elements of array with less noise. The effect of introducing  $\underline{\hat{Q}}^{-1}$  in  $\cos^2(\underline{a}, \underline{d}; \underline{\hat{Q}}^{-1})$  can be compared with  $\cos^2(\underline{a}, \underline{d}; \underline{I})$  of Equation (5.51) where there is no optimization involved. Cox (1973) pointed out that  $\underline{\hat{Q}}^{-1}$  in  $\cos^2(\underline{a}, \underline{d}; \underline{\hat{Q}}^{-1})$  is equivalent to  $\cos^2(\underline{a}, \underline{d}; \underline{I})$  with multiplications of scaling factors to the projections of  $\underline{a}$  and  $\underline{d}$  on the eigenvectors of  $\underline{\hat{Q}}$ .

Since  $\hat{\underline{Q}}$  has been normalized to have its trace equal to  $N$ ,  $\sum_{n=1}^N \lambda_n = N$ . The scaling factors associated with  $\hat{\underline{Q}}^{-1}$  are  $\sqrt{1/\lambda_n}$  which emphasizes components of  $\underline{a}$  and  $\underline{d}$  corresponding to small eigenvalues of  $\hat{\underline{Q}}$ , and de-emphasizes components corresponding to large eigenvalues of  $\hat{\underline{Q}}$ . Therefore the quantity  $\left\{ \frac{\underline{d}' \cdot \hat{\underline{Q}}^{-1} \cdot \underline{d}}{\underline{a}' \cdot \hat{\underline{Q}}^{-1} \cdot \underline{a}} \right\} \cos^2(\underline{a}, \underline{d}; \hat{\underline{Q}}^{-1})$  gives rise to signal gain, for the assumed signal direction.  $\left\{ \frac{\underline{d}' \cdot \hat{\underline{Q}}^{-1} \cdot \underline{d}}{\underline{a}' \cdot \hat{\underline{Q}}^{-1} \cdot \underline{a}} \right\}$  defines the maximum possible gain, and  $(\underline{a}' \cdot \hat{\underline{Q}}^{-1} \cdot \underline{a})$  suppresses the noise. The choice of  $\tilde{\underline{W}}$  results in the optimal estimator which provides the maximum gain for  $\underline{a} = \underline{d}$  and optimal suppression for  $\underline{a} \neq \underline{d}$ .

#### 5.4.5 Grass Valley data processing method

Transient-free quiet-interval data blocks from each of the 12 elements of the array are selected for processing. The number and length of the data blocks are selected on the basis of resolution and statistical stability of the estimated power spectral density. A MLM comparison for different numbers and lengths of data blocks, holding the total number of data points constant, is illustrated with array data from site E5.9W (Figure 5.7). The results of processing the identical data using three different lengths are shown in Figure 5.8 for 12 blocks x 128 points, 24 blocks x 64 points, and 48 blocks x 32 points. The 12 blocks x 128 points (Figure 5.8A) provides the highest resolution, indicating the multiple directions of propagation. However, the greater number of data points in the time domain requires the analysis of more frequency components and the FKPSD estimated in such a way is statistically less stable than the smaller number of data points. The 48 blocks x 32 points (Figure 5.8C) has only 16 discrete frequency

components from DC to Nyquist frequency. In this case, the coarse frequency interval may result in an erroneous phase velocity estimate. I, therefore, have selected 24 blocks x 64 data points (Figure 5.8B) to process Grass Valley array data; this combination provides adequate resolution in wavenumber space and reasonable stability in the statistical estimation. In the figures, FKPSD are estimated for a desired frequency component at each of 41 x 41 grid points in wavenumber space, then normalized with respect to peak value and contoured in dB. Normally, data are processed in a wide frequency band and the maximum FKPSD of each frequency component is plotted as shown in Figure 5.9 for data from sites E5.9W, GP, and A2N. Those frequencies corresponding to the FKPSD maxima are selected for interpretation. The wavenumber and frequency of a FKPSD maximum provides the estimate of apparent phase velocity and direction of propagation for the most coherent propagating plane wave in the data sample. In case the microseismic field has very low coherence across the array sensors, the plot of maximum FKPSD over a wide frequency band will show a low normalized curve without distinct peaks. Accordingly, in Figure 5.9, the coherence of array data recorded at site A2N is higher than that recorded at site GP in the frequency band of 2.54 Hz to 12.9 Hz. A comparison of BFM and MLM is provided in Figures 5.8B and 5.8D for the case of 24 blocks x 64 points. The greater resolution in MLM is apparent in resolving the multiple directions of propagation. Consequently, our processing method was normally MLM using 24 blocks x 64 points.

### 5.5 Grass Valley data interpretation

Valley sites. The noise anomaly in the center of the valley, e.g. E5.9W at 5 to 7 Hz (Figure 4.4B), can be explained by the superposition of multi-path surface waves propagating in the shallow alluvial section. The absence of a unique and time-invariant propagation direction, as seen in Figure 5.10, indicates clearly that the high amplitude ground noise at this site is not due to a local buried source. Further, the uniform propagation velocity,  $|k| \approx 16$  cycles/km, seen at all azimuths suggests a surface wave nature of the noise field. Similar multi-azimuth surface waves are seen also in the f-k results at 5.71 Hz for the array data at sites E0.5W and M2.9N (see Figure 3.4 for site locations).

Leach Hot Springs. Time-invariant azimuths of propagating noise fields are seen at sites A2N, B0.35W, E2.9E, and GP (see Figure 3.2 for site locations) in the vicinity of Leach Hot Springs. Typical noise data recorded in this area show highly coherent energy in the array. Data from site A2N are shown in Figure 5.11. Except at B0.35W, the dominant frequency of the propagating noise field in the area is 4.4 Hz. The f-k plots at this frequency are shown in Figures 5.12A, 5.12C, and 5.12E. The unique azimuth in each plot is in a direction away from the hot springs. At B0.35W the dominant frequency of noise field is 2.8 Hz. In the same frequency band, we also see distinct 2.5 Hz noise components at sites A2N and GP. Noise in this lower frequency band propagates constantly during quiet recording periods at an azimuth around  $210^\circ$ , as indicated in Figures 5.12B, 5.12D, and 5.12F for sites A2N, GP, and B0.35N. The phase velocities estimated from these plots

indicate that the microseisms are apparently fundamental mode Rayleigh waves.

Another interesting feature appears at C6W and at the intersection of lines L and H (site LH) where f-k analyses indicate that the noise fields at both sites propagate away from the hills (Figures 5.13C, 5.13D, 5.13E and 5.13F). There is a possibility that the hills respond to gusty winds and generate ground noise.

#### 5.6 Dispersion characteristics and shallow structure

On the assumption that the microseismic field consists of surface waves, the f-k analysis technique allows direct measurement of the local dispersion curve by selecting phase velocities corresponding to the frequencies with peak FKPSD. As an example, in Figure 5.14 we show phase velocities so estimated, along with computed fundamental and first higher mode Rayleigh wave dispersion curves, for a model based on P-wave velocities from a shallow refraction survey in the area. The effect of the very shallow velocity structure is illustrated clearly. Lateral variations in the upper 10-20 m will control the surface wave propagation characteristics. In estimating dispersion curves, we do not restrict sampling to the quiet periods, since larger microseisms are very coherent across the array. The dispersion measurements, besides providing local observations of phase velocity for shallow structure mapping, also provide a method of verifying the wave nature of the microseisms. It is clear that waves with periods of 1 sec and greater must be analyzed for structural information at geothermal target depths, if the microseisms are fundamental mode Rayleigh waves (see, for example, McEvelly and Stauder, 1965).

FIGURE CAPTIONS

- Figure 5.1 Estimated coherence near Leach Hot Springs for various geophone spacings of (A) 10 m, (B) 30 m, (C) 40 m, and (D) 60 m in a single line. Note the coherence decreases as geophone spacing increases, along with narrower coherent frequency bands. Frame E is the typical VSD of this site.
- Figure 5.2 Estimated coherence at valley sites, EM (A-D) and E2.75W (E), for linear geophone spacings of (A) 20 m (NS direction), (B) 20 m (EW direction), (C) 40 m, (D) 60 m, and (E) 120 m.
- Figure 5.3 Array configuration and its contoured impulse response in wavenumber space, plotted to  $k_x$  and  $k_y = 71$  cycles/km. The effective Nyquist wavenumber can be seen to vary with azimuth in the range of about 50 - 70 cycles/km. The interior square outlines the standard f-k plot range of 35.7 cycles/km, used in subsequent figures. Radii of the concentric circles in array are indicated.
- Figure 5.4 High-resolution f-k power spectral density estimates for a simulated 4 Hz plane wave signal propagating  $N60^\circ E$  across the array at horizontal phase velocity 200 m/s ( $k=20$  cycles/km), to illustrate spatial aliasing. The array configuration is the same as shown in Figure 5.3. The array dimension scales are (A) 1 time, (B) 1.5 times, (C) 5 times, (D) 10 times, the radii values indicated in Figure 5.3. The maximum  $k_x$  and  $k_y$  values in the plots are (A) 71.4, (B) 47.6, (C) 14.3, (D) 7.1 cycles/km corresponding approximately to the effective Nyquist wavenumbers for the arrays.

The f-k power spectral density contours are -1.0, -3.0, -6.0, -9.0, -12.0 dB below the main peak. Circles indicate the constant velocities shown, expanding with array size. Aliasing is apparent in the high phase velocities in Frames (B), (C), and (D), easily misinterpreted as detected body waves. The 90% confidence limits on the estimated FKPSD are  $\pm 1.9$  dB, based on the multiple of chi-square variable with approximately 26 degrees of freedom.

Figure 5.5 Noise-free simulated 4 Hz plane waves crossing the array of Figure 5.3 with horizontal phase velocity of 200 m/sec.

Figure 5.6 Comparison of f-k resolutions using BFM (A) and MLM (B to F) on the simulated plane waves shown in Figure 5.5, with noise added. MLM resolution deteriorates as the signal-to-noise ratio of the input data varies from (B) 48.8 dB, (C) 30.6 dB, (D) 25.6 dB, (E) 20 dB, to (F) 10.6 dB, whereas BFM is insensitive to signal-to-noise ratio in the same range. The BFM resolutions for all noise levels are all identical to (A). Contoured lines are at -1, -3, -6, -9, and -12 dB below the main peak.

Figure 5.7 Typical seismic noise data in the valley, at site E5.9W, recorded by the array shown in Figure 5.3. The difference in signal amplitude among traces is a consequence of the difference in geophone damping resistor of each channel. The scale factor, in  $10^{-9}$  m/sec per cm of displayed record amplitude, at 4.5 Hz, are shown for each trace on the right margin.

Figure 5.8 f-k results for site E5.9W for different data block lengths, comparing MLM and BFM.

- (A) 12 data blocks, each with 128 points, processed by MLM, with 90% confidence limits of +7.7 dB.
- (B) 24 data blocks, each with 64 points, processed by MLM, with 90% confidence limits of +1.96 dB,
- (C) 48 data blocks, each with 64 points, processed by MLM with 90% confidence limits of +1.16 dB,
- (D) 24 data blocks, each with 64 points, processed by BFM, with 90% confidence limits of +1.4 dB.

The frequency on each frame corresponds to a maximum FKPSD (see Figure 5.9). The range of wavenumber plotted is 35.7 cycles/km. Contour lines are at -1, -3, -6, -9, and -12 dB below the main peak.

Figure 5.9 Examples of maximum FKPSD plots as a function of frequency, independent of the wavenumber. The maximum FKPSD at each frequency is normalized with respect to a common, reference power level. Array data are taken from sites A2N, E5.9W and GP.

Figure 5.10 The time-variant nature of propagating seismic noise at valley site E5.9W. Data are selected from the 4th, 5th, 7th, 8th, 9th, and 10th hours of the recording period, as indicated, and processed by MLM using 24 blocks with 64 points per data block. Circles indicate velocities of peaks given along with azimuth in each plot.

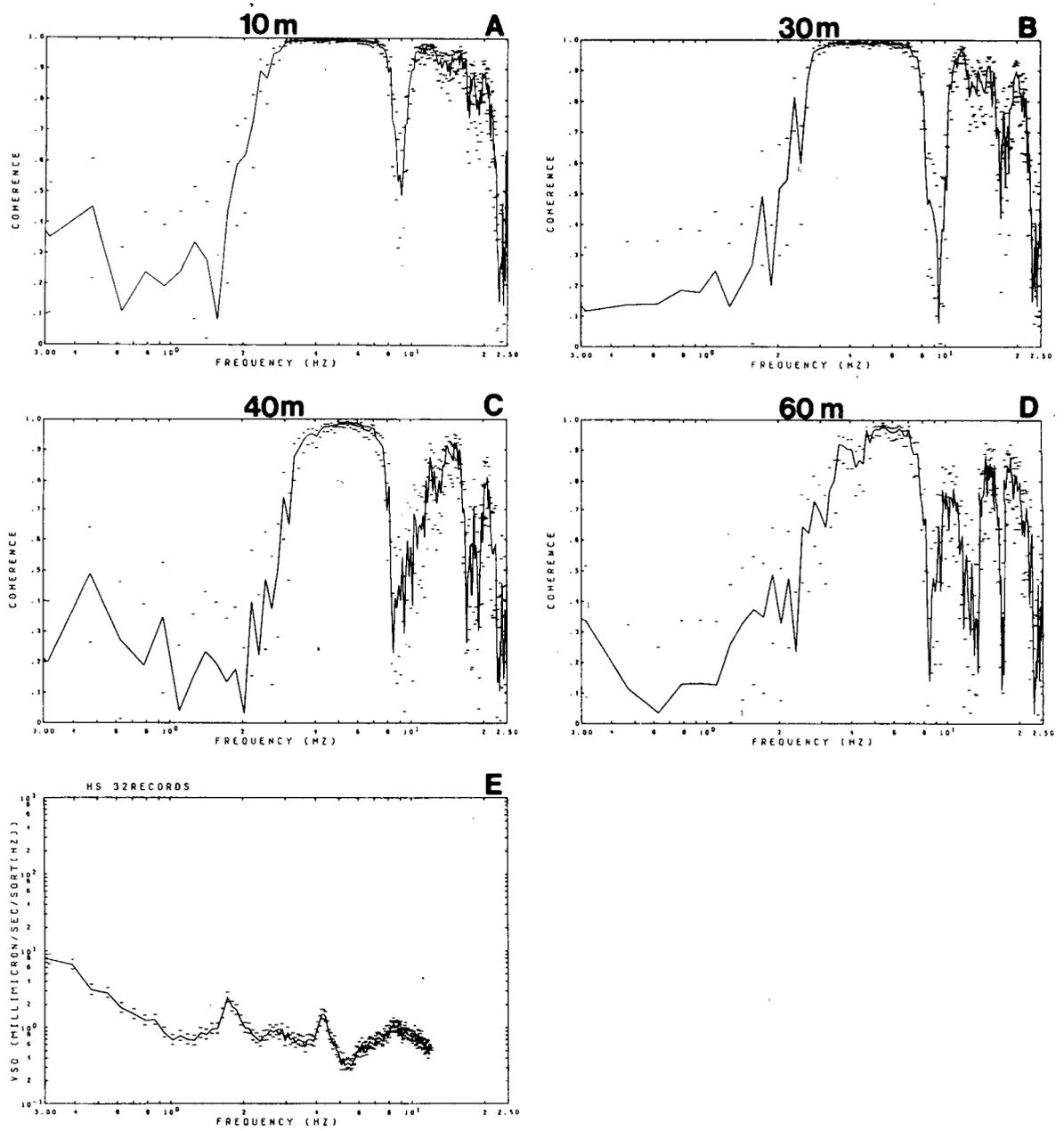
Figure 5.11 Typical seismic noise data at site A2N, 1 km SE of Leach Hot Springs, recorded by the array shown in Figure 5.3. Note the high coherence across the array. The difference in signal amplitude among traces is resulted from the different geophone damping resistor of each channel. The scale factor, in  $10^{-9}$  m/sec per cm of displayed record amplitude, at 4.5 Hz, are shown for each trace on the right margin.

Figure 5.12 High-resolution f-k results in wavenumber space for sites A2N, GP, E2.9E, and B0.35W located near Leach Hot Springs. The frequency on each frame corresponds to a maximum in the maximum FKPSD plot (Figure 5.9). The maximum wavenumber plotted is 35.7 cycles/km, except for frame (F), where it is 17.9 cycles/km. The contoured lines are -1, -3, -6, -9, and -12 dB below the main peak. The horizontal phase velocity and the time-invariant direction of the propagating seismic noise field are indicated on each frame. These noise components are apparently fundamental mode Rayleigh wave generated at the hot springs, where near-surface velocities exceed 4 km/sec.

Figure 5.13 High-resolution f-k results in wavenumber space for sites B2.6W, C6W and LH. The maximum wavenumber plotted in 35.7 cycles/km. The contoured lines are -1, -3, -6, -9, and -12 dB below the main peak.

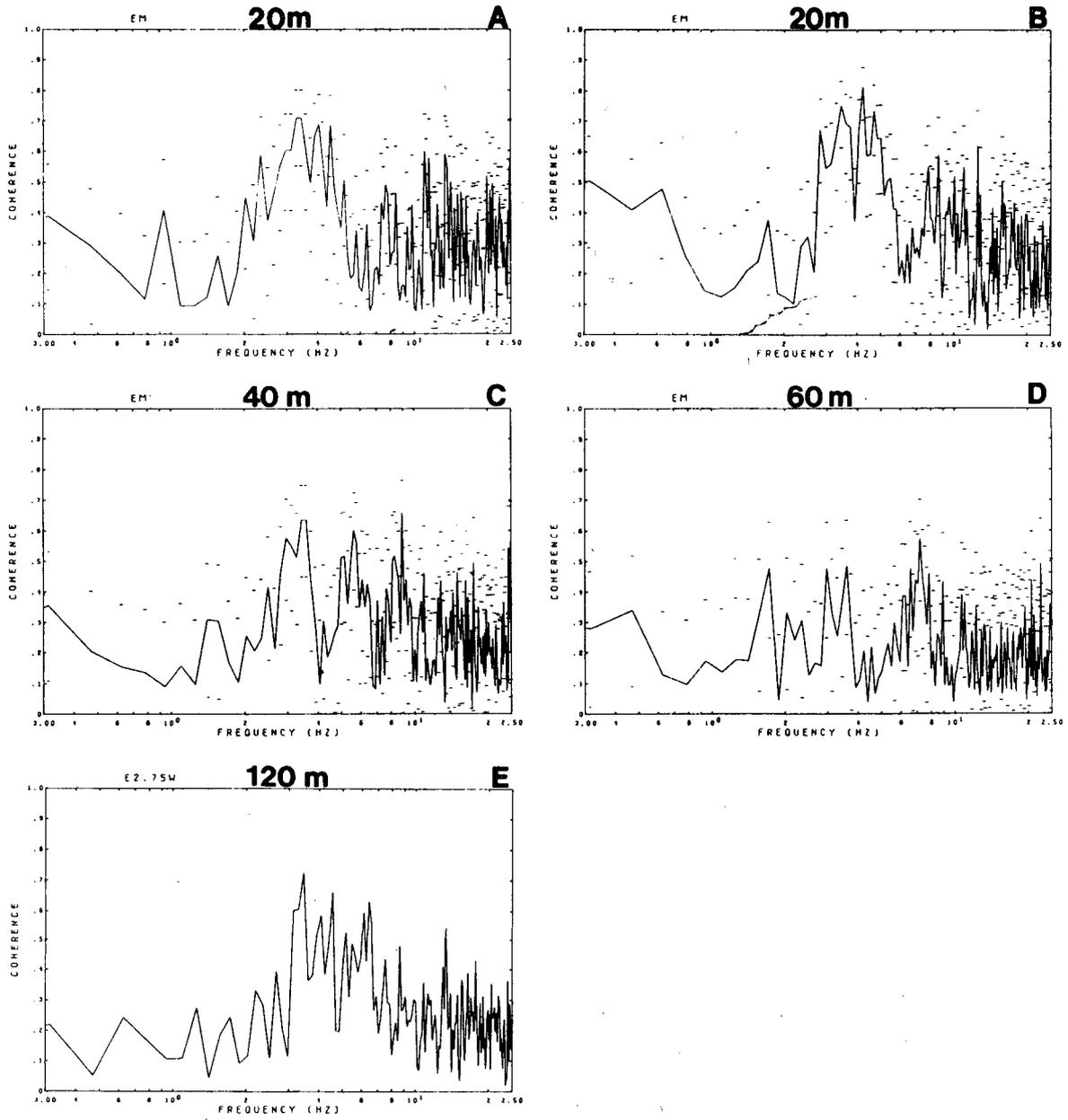
Figure 5.14 Rayleigh wave dispersion curves for fundamental and first higher modes, computed for the model shown, compared with observed ground noise phase velocities at site E5.9W.

The observed phase velocities were determined at various times by f-k analysis, the hour indicated by symbol type.



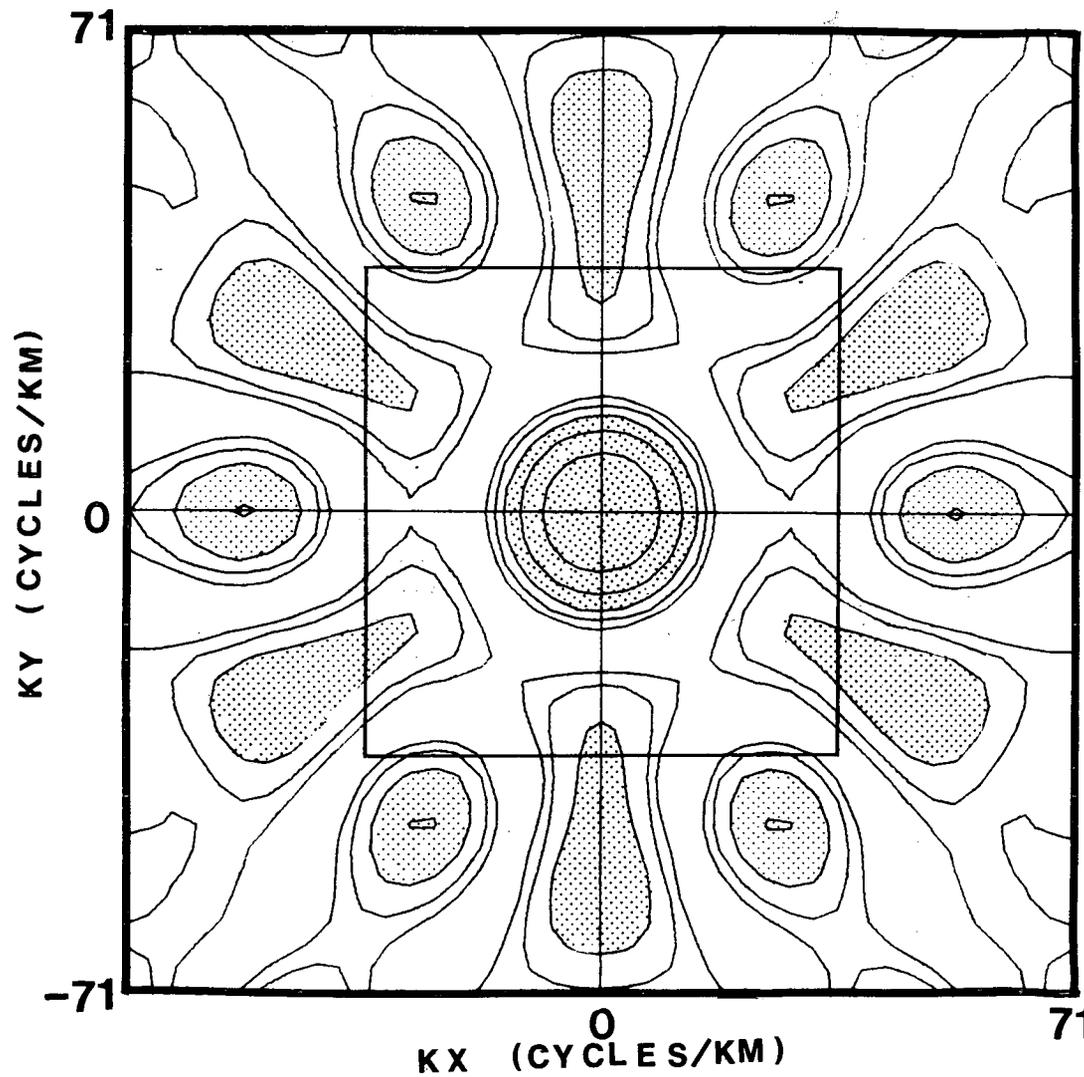
XBL 778-9827

Figure 5.1

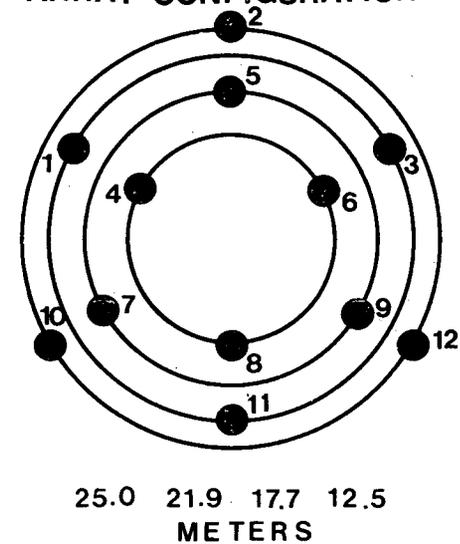


XBL 778 9802

Figure 5.2

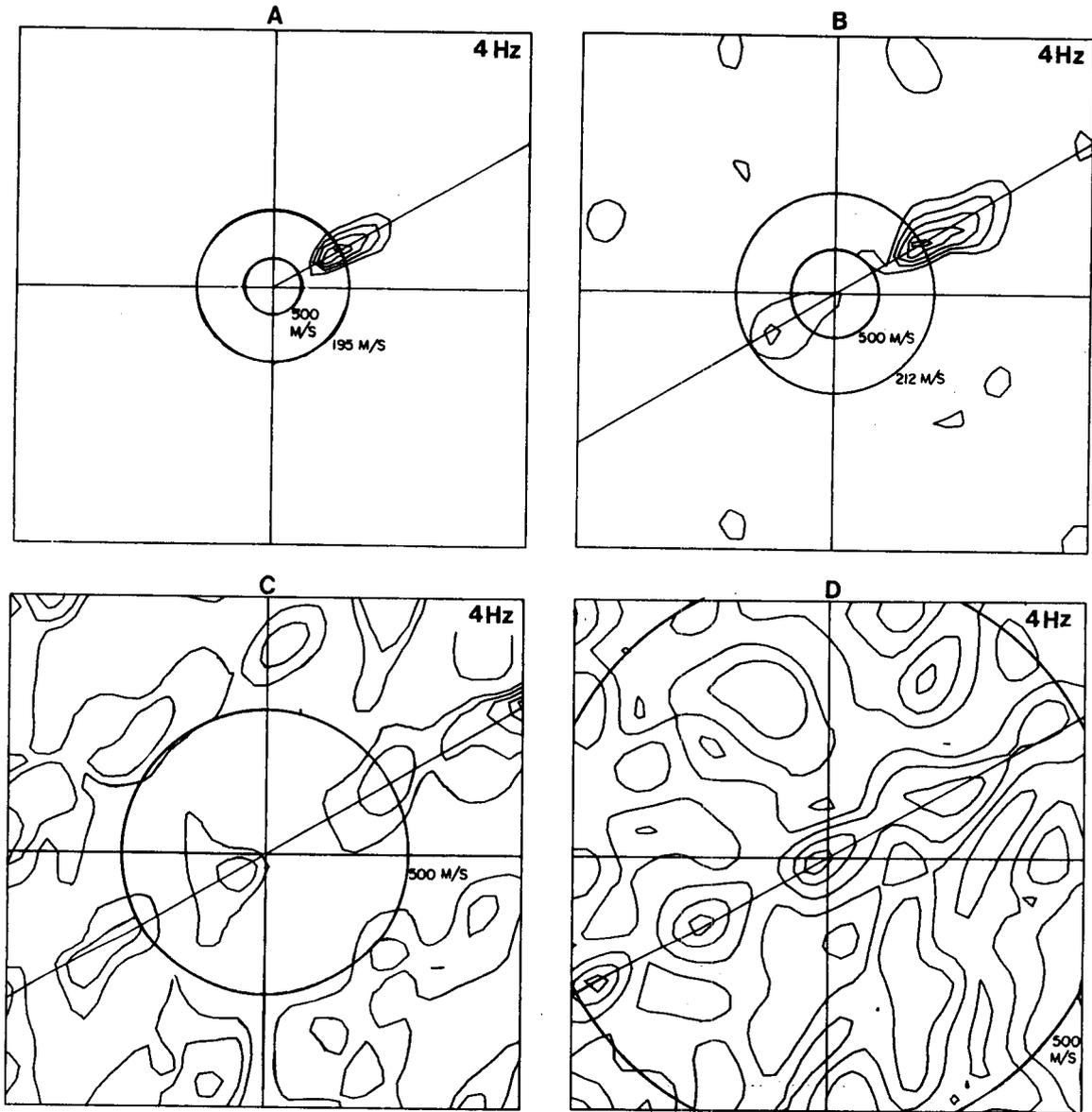


ARRAY CONFIGURATION



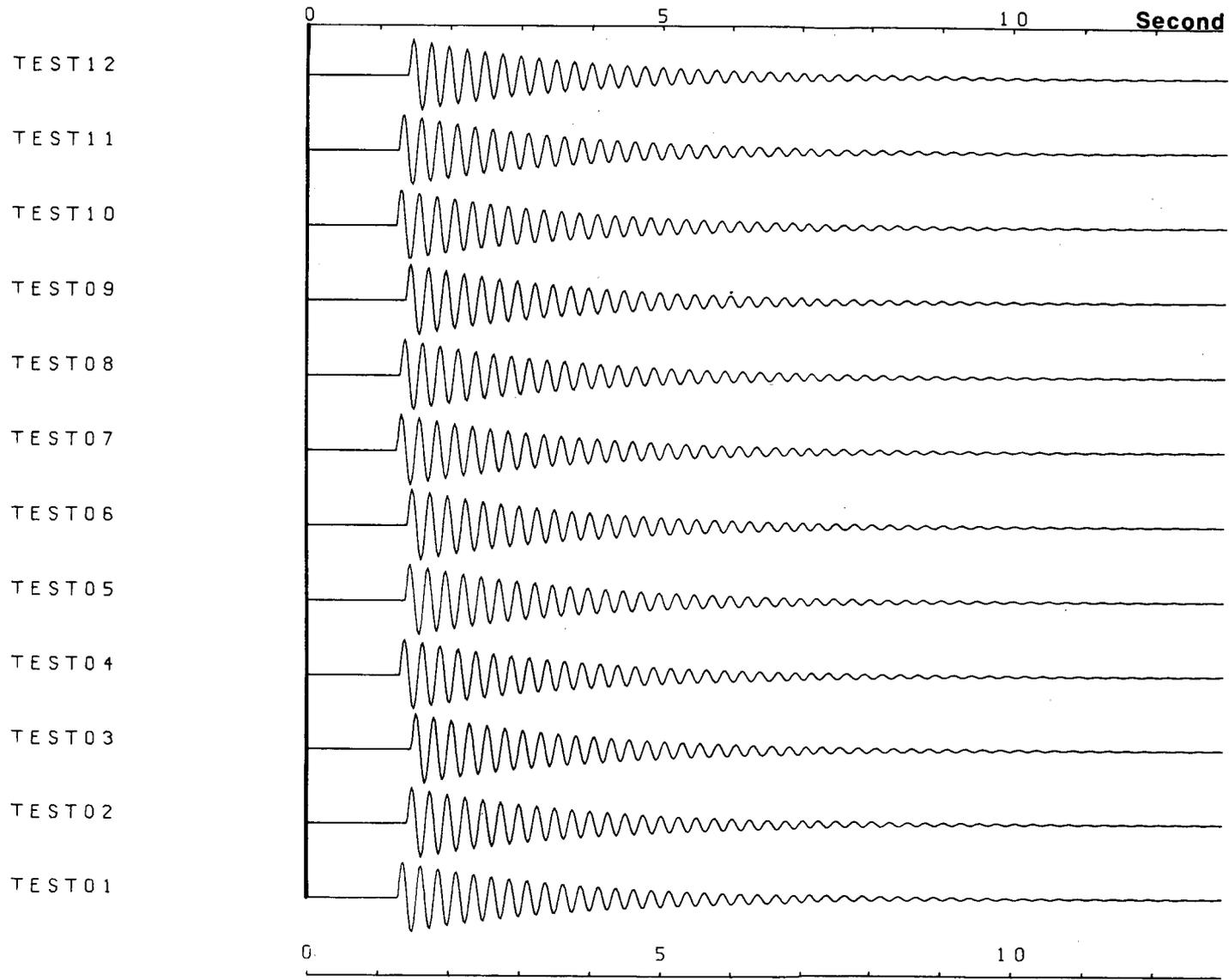
XBL 7610- 4980

Figure 5.3



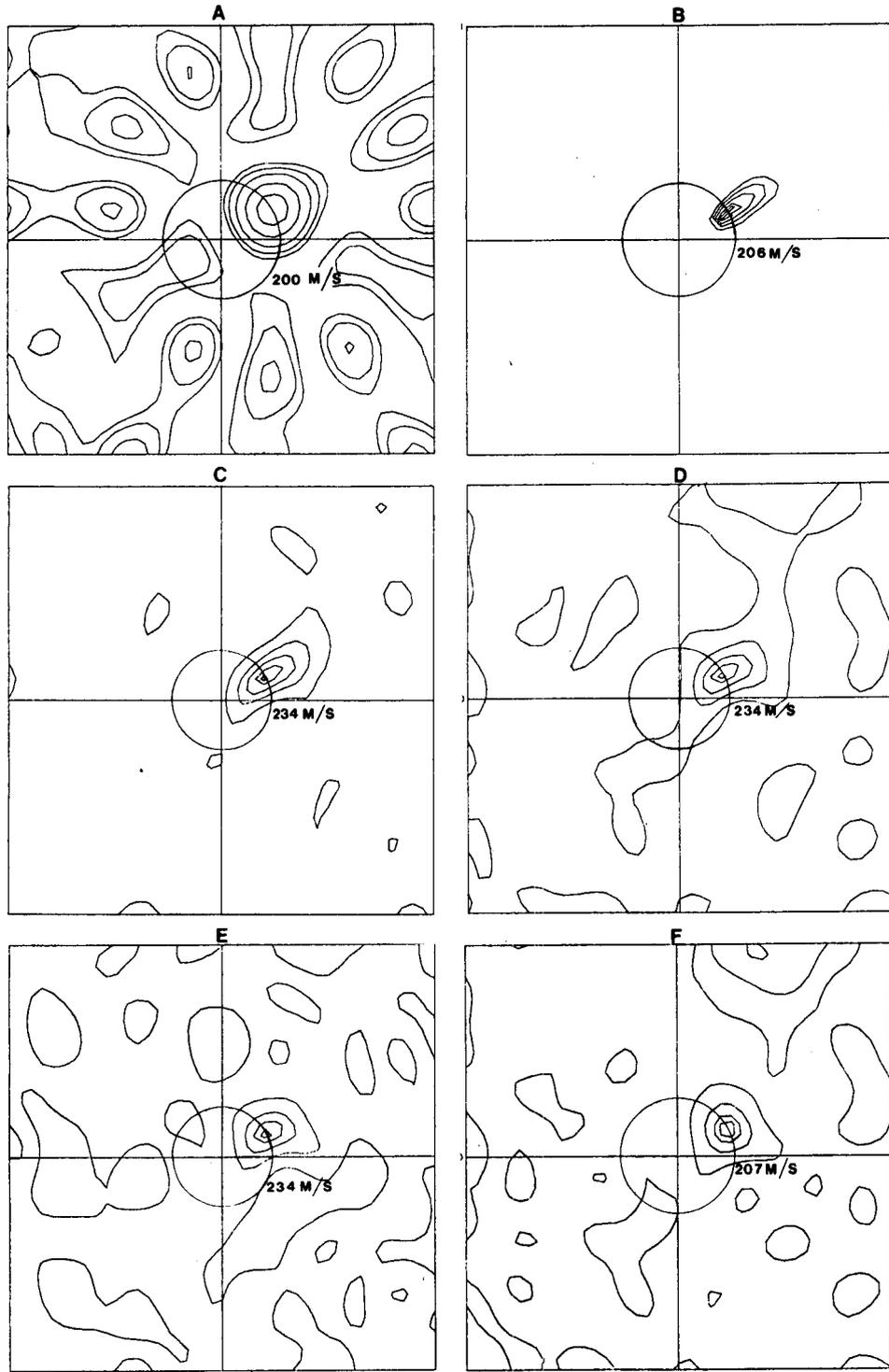
XBL 776=9467

Figure 5.4



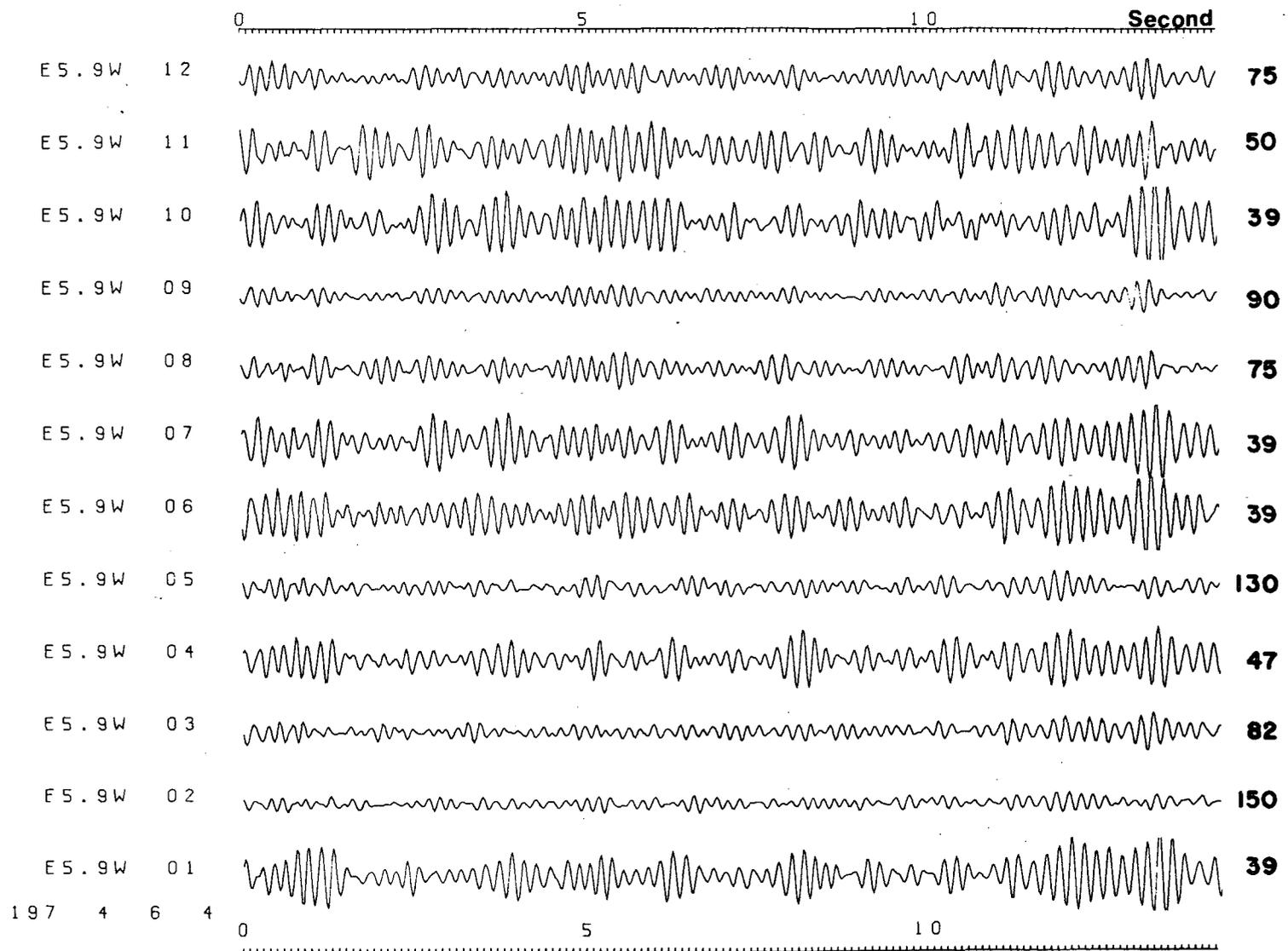
XBL 773-8115

Figure 5.5



XBL 777-9618

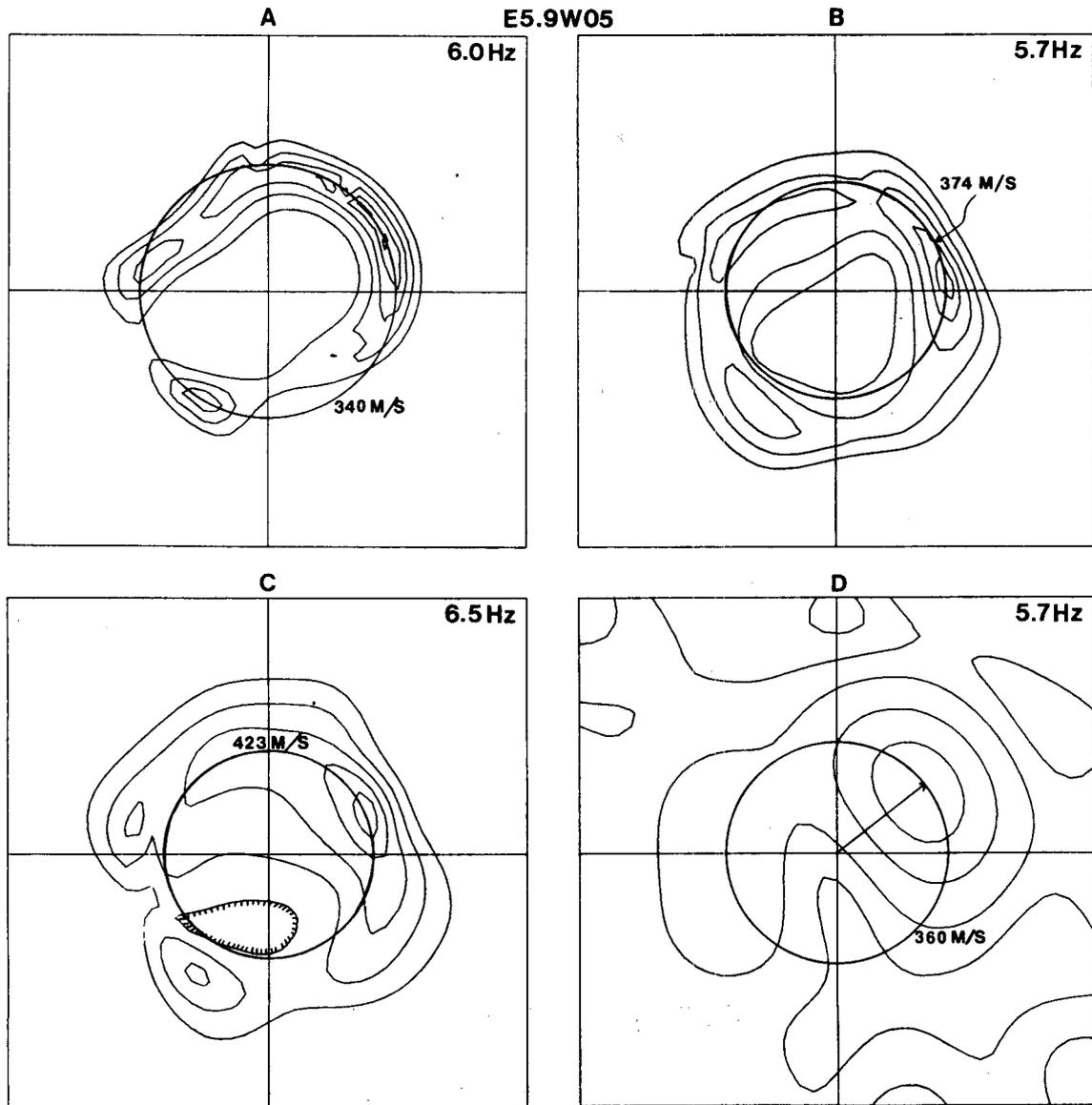
Figure 5.6



197 4 6 4

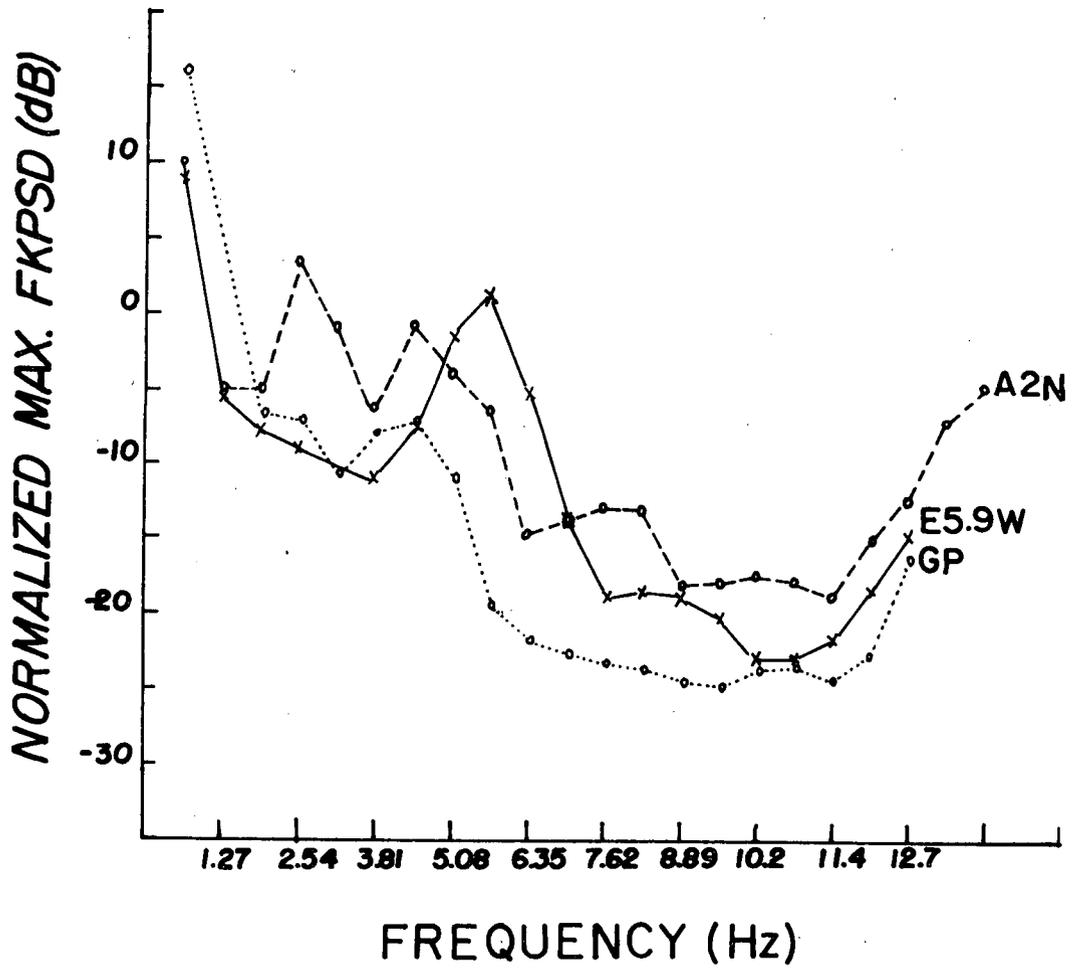
XBL 773-8116

Figure 5.7



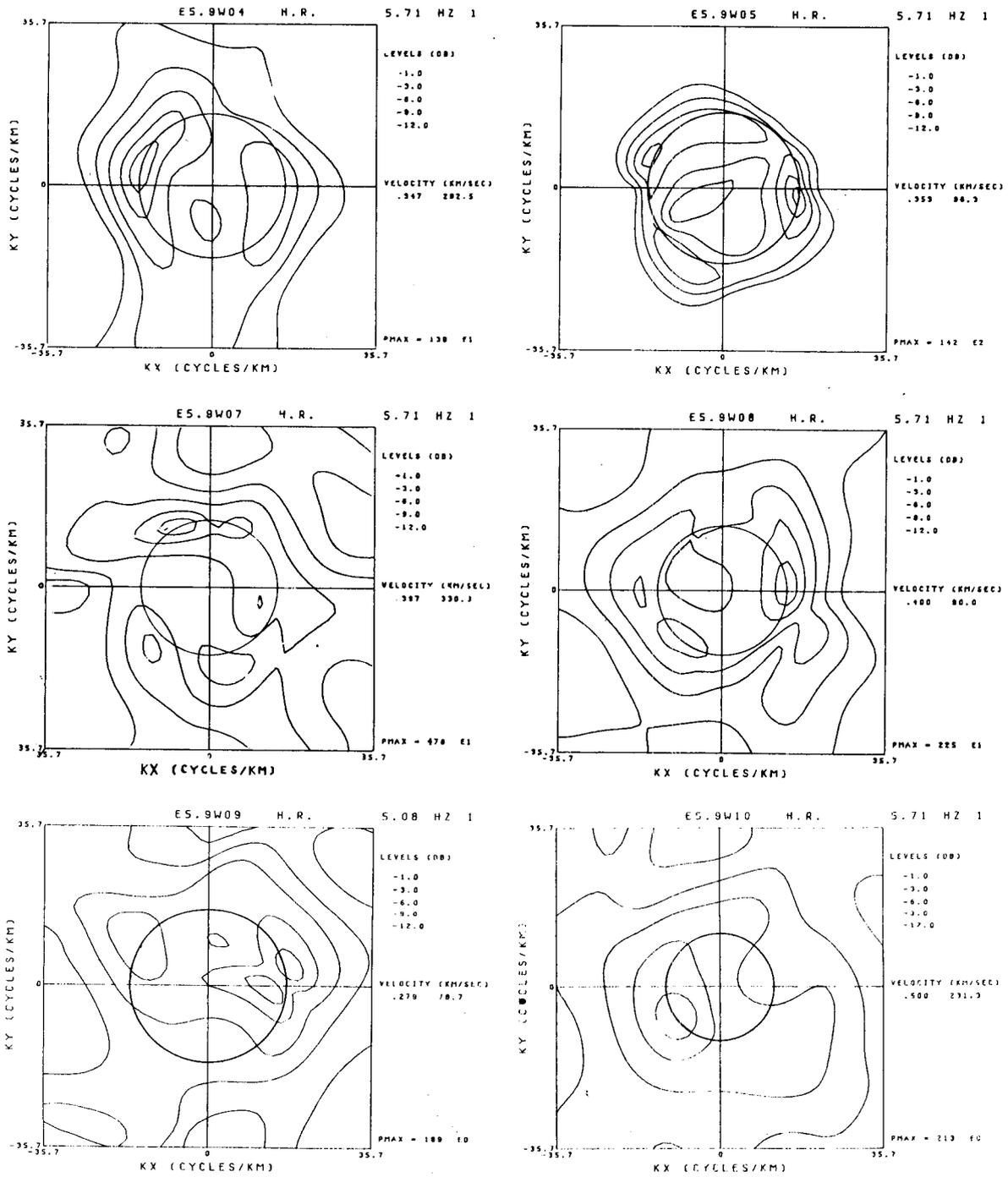
XBL 776-9468A

Figure 5.8



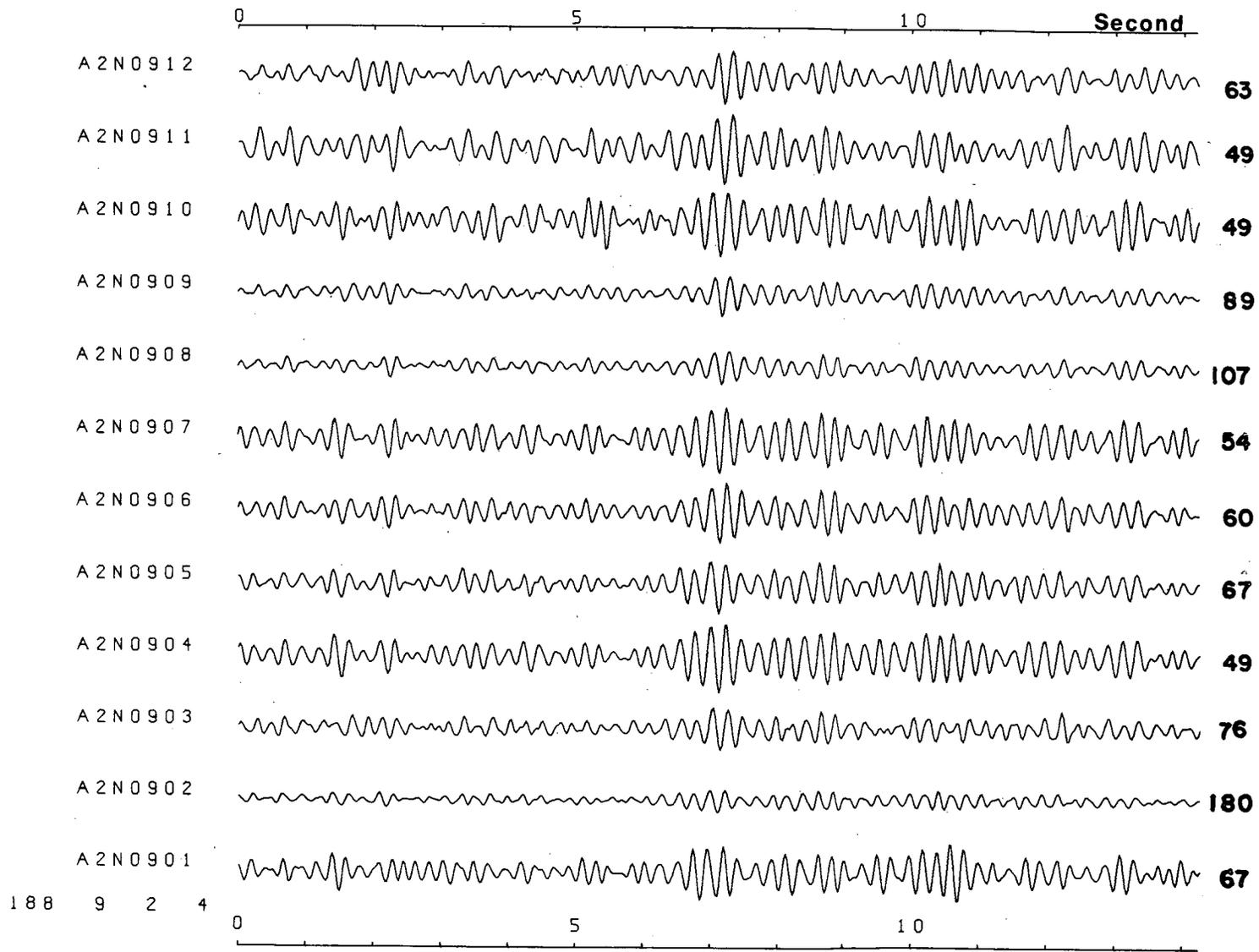
XBL 779-2463

Figure 5.9



XBL 772-7631

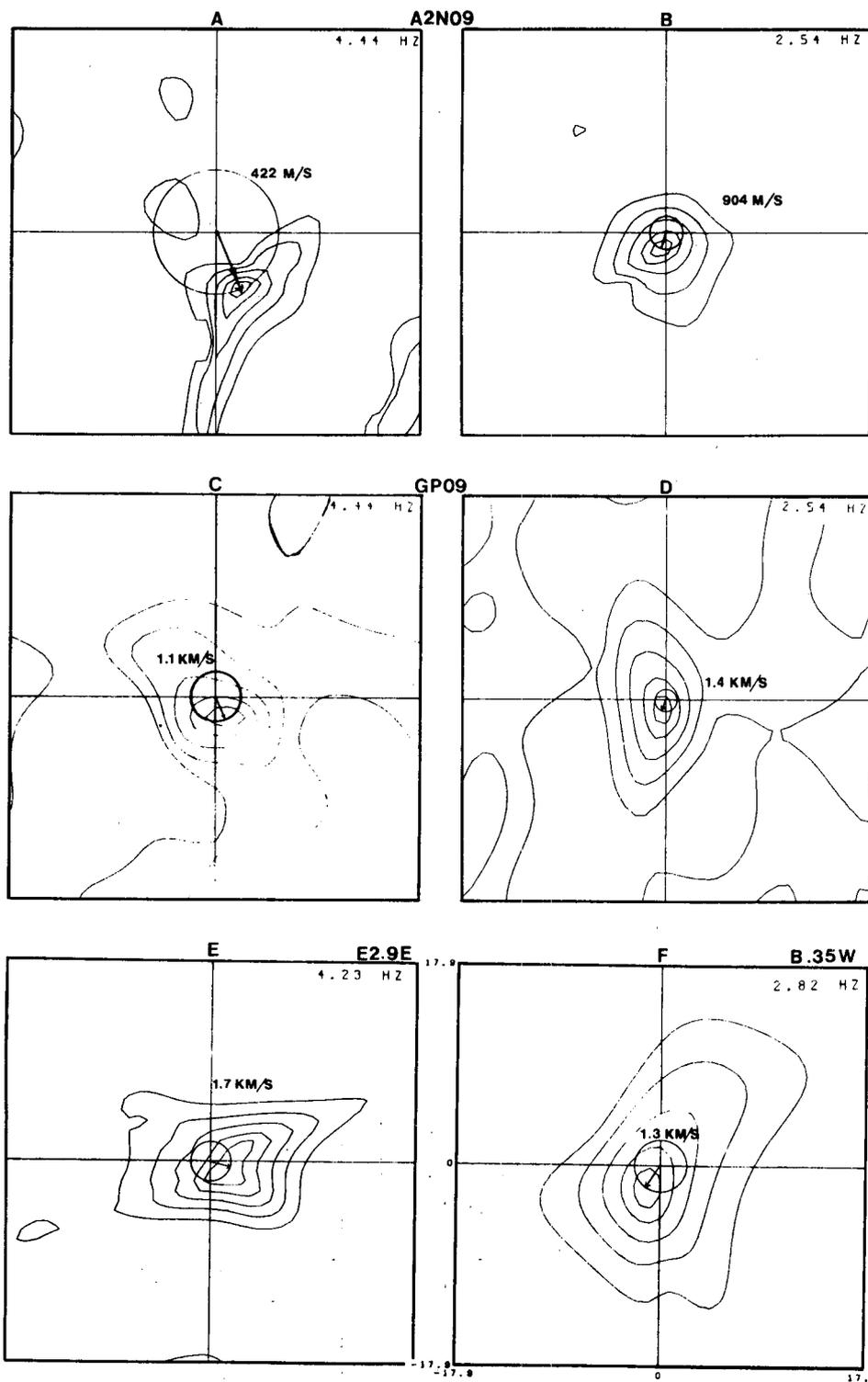
Figure 5.10



-110-

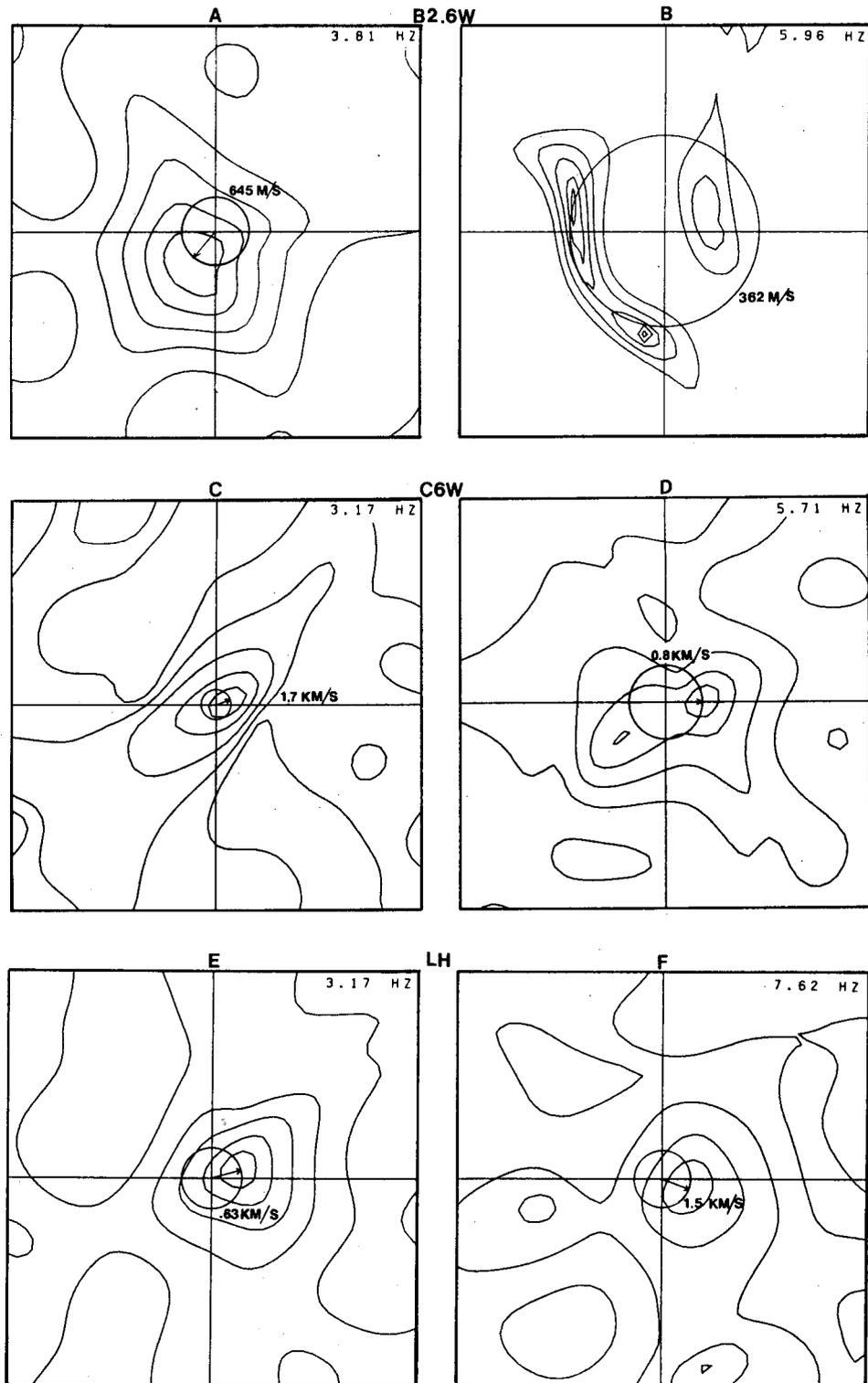
XBL 773-8117

Figure 5.11



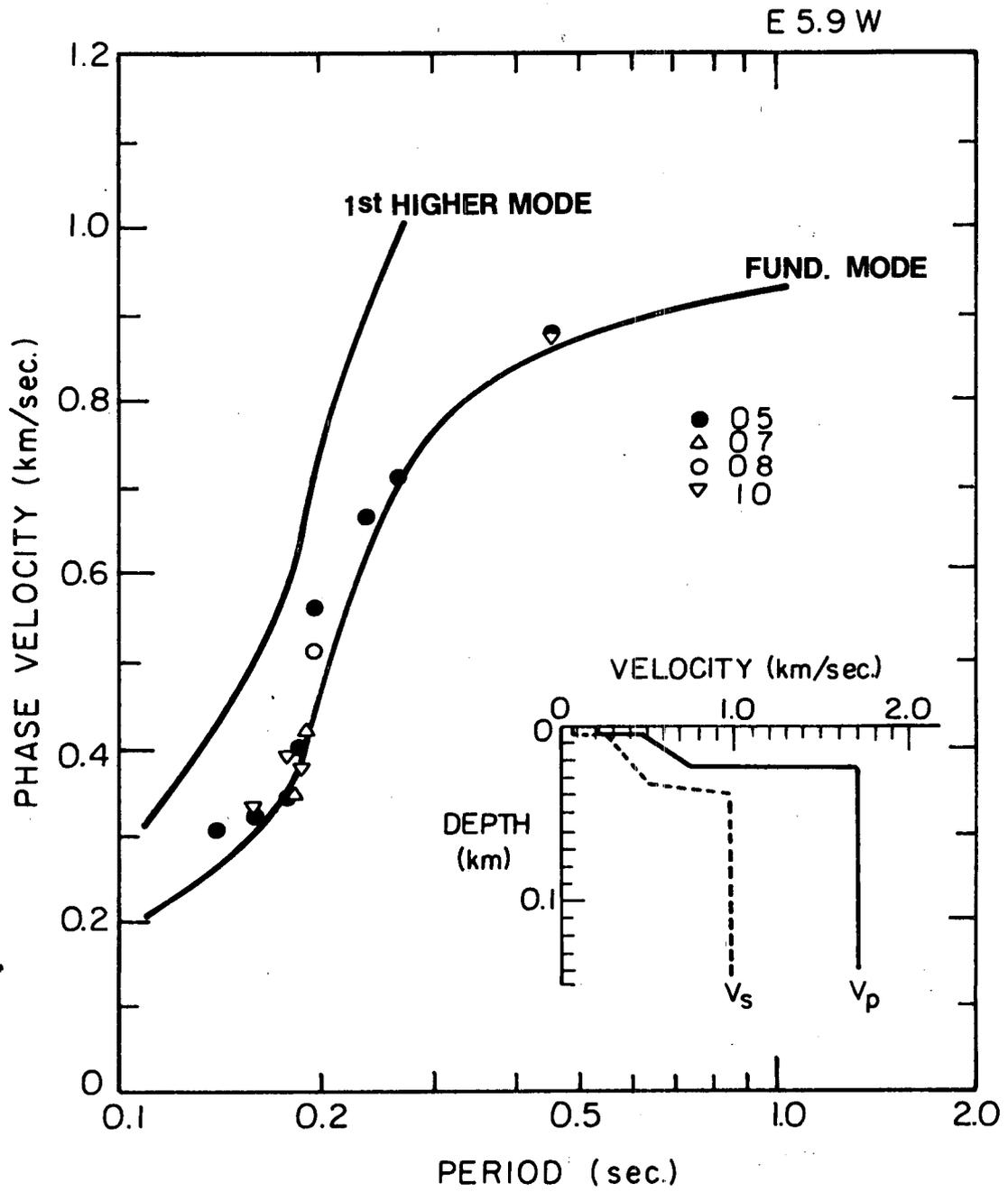
XBL 778 9801

Figure 5.12



XBL 777-9617

Figure 5.13



XBL 771-5005

Figure 5.14

## VI. SUMMARY AND RECOMMENDATIONS

### 6.1 Summary

The main results of a study of microseisms in Grass Valley, Nevada, using a simple field system, are summarized below:

Diurnal variation in the 2-20 Hz noise field is regular. A consistent diurnal variation that repeats from day to day is due apparently to meteorological and cultural sources, with typically 15 dB variation seen from the mid-day high noise level to the low noise level in the early morning hours of 2-4 A.M. Secular variations, due to regional weather patterns, can produce a 5-10 dB range in the early morning minimum noise levels over a duration of a few days.

For spectral stability in investigating spatial variation of noise, at least 28 quiet data blocks, each 12.8 sec long, were taken simultaneously at the network stations and the spectra were averaged for each site. This procedure produced consistent results throughout the area, revealing a characteristically low amplitude smooth noise spectra at hard rock sites, a prominent peak at 4-6Hz at valley sites, and wideband, high amplitude noise, apparently due to very shallow sources, at hot springs sites. Contour maps of noise level, normalized to a reference site, are dominated by the hot springs noise levels outlining the regions of maximum alluvium thickness. Major faults are evident when they produce a shallow lateral contrast in rock properties.

Microseisms in the 2-10 Hz band are predominantly fundamental mode Rayleigh waves, characterized by low velocities and wavelengths as small as 20 m, requiring closely spaced arrays for adequate spatial sampling.

High resolution f-k processing, with proper data sampling, provides a powerful technique for mapping the phase velocity and the direction of propagation of the noise field, revealing local sources and lateral changes in shallow subsurface structure.

No evidence for a significant body wave component in the noise field was found, although it becomes clear that improper spatial sampling can give a false indication through aliasing. Noise emanating from a deep reservoir would be evident as body waves and could be traced to its source given a reasonably accurate velocity model.

## 6.2 Recommendations

Conventional seismic ground noise surveys, conducted as outlined in this study, require a large number of stations for economical implementation. With 100 stations, for example, a week-long survey could provide maps of noise amplitude distribution, P-wave delay time, and microearthquake locations, as well as f-k analyses at many sites, utilizing a 2-3 man crew. It is not clear, however, that such data will be of significant value in delineating a geothermal reservoir.

The amplitude mapping of ground noise in certain frequency bands is a poor exploration technique for delineating buried geothermal systems. The results of the survey described in Chapter 4 indicate that the amplitude variations of microseisms in an area are controlled by the near-surface geology, especially lateral variations in thickness of the alluvial layer. The large amplitude surface wave generated by surface sources and propagating horizontally, will mask weak seismic waves emitted from a buried source. Therefore, amplitude mapping only

reveals information on the very shallow structure.

On the other hand, the technique of f-k analysis can, theoretically, map the wavenumber of the microseisms, discriminating the vertically incident body waves from the surface waves. The yet open question of whether a reservoir acts as a radiator of seismic body waves can be answered through careful f-k analyses in existent geothermal areas. The array to be used for further study must be a non-aliased array of larger diameter than that used in this study. The expansion in array size will improve the resolution around the origin of the f-k diagram. This improvement would provide a more accurate estimate for power at the small wavenumbers, such that the azimuth and the apparent velocity of the long-wavelength body waves are estimated more accurately. The amplitudes of body waves radiating from a source at depth are apparently much smaller than those of the ambient surface waves. In order to extract useful information from the body waves, a sophisticated signal detection and processing scheme is required. However, the f-k analysis technique may fail to detect the geothermal system at depth if our assumption of body wave radiation from the reservoir is not valid, or if the emanating body waves are either attenuated or completely masked by the ambient surface waves. It is fortunate that the ambient surface waves have shorter wavelengths than the body waves because then the detection of weak body waves can be improved by designing a more sophisticated array to cancel short wave-lengths, as is commonly done in conventional seismic reflection surveying.

If the assumption about radiated body waves is indeed valid, and such body waves are detectable, we can trace the recorded wavefronts

to their source, given a reasonably accurate velocity model. There are two schemes which have been used for projecting waves observed at the surface back into the earth and locating the source region, and these methods may be applicable to the geothermal reservoir delineation problem.

The first method is seismic ray tracing described by Julian (1970); Engdahl and Lee (1976). If the array diameter is much smaller than the distance to the buried source, the microseismic field propagates as a plane wave across the array. Estimation of the azimuth and the apparent velocity of the propagating noise field from f-k analysis along with the knowledge of the near-surface velocity distribution, can give us the incident angle of noise. Given a reasonable velocity structure in the area and simultaneously occupied array sites, we can reconstruct ray paths to each site. The intersection of these ray paths outlines the region of the radiating source.

Another approach is much like that used in a conventional reflection survey with two-dimensional surface coverage but without a surface controlled source. The coherent noise fields recorded by a two-dimensional surface array are projected downward into the assumed subsurface model. The reconstruction of the coherent noise field propagating in an upward direction, at a selected frequency, can be carried out by the wave equation migration technique, using a finite-difference approximation such as described by Claerbout (1976). The restriction of this approach to microseismic data is that the noise field must propagate as a spherical wavefront across the geophone array. The spherical wavefront exists in the situation where the array dimension is greater than the

distance to the source. In this case, we can outline the region of radiating sources in terms of the convergent pattern of the extrapolated wave fields.

It is clear that the ray tracing and the wave equation migration are applicable at different source-array distances in the application of delineating geothermal reservoirs. In a practical exploration program, we do not know the depth of geothermal reservoirs, nor do we know the shape of the wavefront across the array. One way of solving the problem is to place a non-aliased array at several sites and search for the evidence of time-invariant, high velocity body waves. As soon as the body waves are detected, one may compare several f-k diagrams, using data of identical recording periods but of different sizes of sub-array. The deterioration of the resolution in the f-k diagrams as we expand the size of the sub-array, indicates that the plane wave assumption is violated and the wavefront migration techniques should be applied. On the other hand, if the noise fields propagate as plane waves across the large array, the resolution in the f-k diagrams will be improved as we expand the size of sub-arrays and the f-k analysis with seismic ray tracing are the proper techniques to locate the noise source.

Based on this study, I conclude that, if the geothermal system is indeed emanating detectable body waves, the analysis of ambient ground motion or seismic noise, can be applied to the delineation of geothermal reservoirs. In fact, if the radiated body waves exist, the method can be one of the most effective geophysical methods in geothermal explorations.

REFERENCES

- Báth, M., 1974, Spectral analysis in geophysics: p. 448-462, Elsevier Scientific Publishing Company.
- Beyer, A., Dey, A., Liaw, A., Majer, E., McEvelly, T.V., Morrison, H.F., and Wollenberg, H., 1976, Geological and geophysical studies in Grass Valley, Nevada: Preliminary open file report, LBL-5262.
- Bendat, J.S., and Piersol, A.G., 1971, Random data: Analysis and measurement procedures, Wiley interscience, p. 193-194.
- Borcherdt, R.D., 1970, Effects of local geology on ground motion near San Francisco Bay: Bull. Seismol. Soc. Am. vol. 60, p. 29-61.
- Brillinger, D.R., 1975, Time series: data analysis and theory: Holt, Rinehard and Winston Inc., p. 125 and p. 151.
- Brune, J.N., and Oliver, J., 1959, The seismic noise of the earth's surface: Bull. Seismol. Soc. Am. vol. 49, p. 349-353.
- Burg, J.P., 1964, Three-dimensional filtering with an array of seismometers: Geophysics. vol. 29, p. 693-713.
- Capon, J., 1969, High-resolution frequency-wavenumber spectrum analysis: Proc. IEEE, vol. 57, p. 1408-1418.
- Capon, J., 1973, Signal processing and frequency wavenumber spectrum analysis for a large aperture seismic array: in Methods of Computational Physics vol. 13, p. 1-59, Academic Press.
- Capon, J., Greenfield, R.J., and Kolker, R.J., 1967, Multidimensional maximum-likelihood processing of a large aperture seismic array, Proc. IEEE, vol. 55, p. 192-211.
- Capon, J., and Goodman, N.R., 1970, Probability distributions for estimators of the frequency-wavenumber spectrum; Proc. IEEE, vol. 58, p. 1785-1786.
- Claerbout, J.F., 1976, Fundamentals of geophysical data processing, McGraw-Hill Book Company, p. 184-226.
- Clacy, G.R.T., 1968, Geothermal ground noise amplitude and frequency spectra in New Zealand volcanic region: J. Geophys. Res. vol. 73, p. 5377-5383.
- Combs, J., and Rotstein, Y., 1975, Microearthquake studies at the Coso geothermal area, China Lake, California: 2nd U.N.Sym. on the Dev. and Use of Geothermal Resources, San Francisco, 1975, p. 909-916.
- Cox, H., 1973, Resolving power and sensitivity to mismatch of optimum array processors: Jour. Acoustical Soc. of A., vol. 54, no. 3, p. 771-785.

- Douze, E.J., 1967, Short-period seismic noise: Bull. Seism. Soc. Am., vol. 57, p. 55-81.
- Douze, E.J., and Sorrells, G.G., 1972, Geothermal ground-noise surveys: Geophysics, vol. 37, p. 813-824.
- Engdahl, E.R., and Lee, W.H.K., 1976, Relocation of local earthquakes by seismic ray tracing: J. Geophys. Res., vol. 81, p. 4400-4406.
- Fix, J.E., 1972, Ambient earth motion in the period range from 0.1 to 2560 sec.: Bull. Seism. Soc. Am., vol. 62, p. 1753-1760.
- Frantti, G.E., 1963, The nature of high-frequency earth noise spectra: Geophysics, vol. 28, p. 547-562.
- Frost, O.L., 1972, An algorithm for linearly constrained adaptive array processing: Proc. IEEE, vol. 60, p. 926-935.
- Goforth, T.T., Douze, E.J., and Sorrells, G.G., 1972, Seismic noise measurements in a geothermal area: Geophysical Prospecting, vol. 20, p. 76-82.
- Gutenberg, B., and Andrews, F., 1956, Bibliography of microseisms: 2nd ed., Seismological Laboratory, California Institute of Technology, Pasadena, 134 p.
- Haubrich, R.A., Munk, W.H., and Snodgrass, F.E., 1963, Comparative spectra of microseisms and swell: Bull. Seismol. Soc. Am., vol. 53, p. 27-37.
- Haubrich, R.A., and Mackenzie, G.S., 1965, Earth noise 5-500 millicycles per second: 2. Reaction of the earth to oceans and the atmosphere: J. Geophys. Res., vol. 70, p. 1429-1440.
- Haubrich, R.A., and McCamy, K., 1969, Microseisms: Coastal and pelagic sources: Rev. of Geophys. Res., vol. 70, p. 539-571.
- Iyer, H.M., 1958, A study of direction of arrival of microseisms at Kew Observatory, Geophys. Jour., vol. 1, p. 32-43.
- Iyer, H.M., 1974, Search for geothermal seismic noise in the East Mesa area, Imperial Valley, California: U.S.G.S. Open-file report no. 74-96. 52 p.
- Iyer, H.M., and Hitchcock, T., 1974, Seismic noise measurements in Yellowstone National Park: Geophysics, vol. 39, p. 389-400.
- Iyer, H.M., and Hitchcock, T., 1976, Seismic noise survey in Long Valley, California: J. Geophys. Res., vol. 81, p. 821-840.
- Jenkins, G.M., and Watts, D.G., 1968, Spectral analysis and its application: Holden-Day Inc., p. 255-257.

- Julian, B.R., 1970, Ray tracing in arbitrary heterogeneous media: Tech. Note 1970-45, Lincoln Lab., Lexington, Mass.
- Kanai, K., and Tanaka, T., 1961, On microtremors. VIII: Bull. of the Earthquake Res. Inst., vol. 39, p. 97-114.
- Kanai, K., Tanaka, T., Morishita, T., and Osada, K., 1966, Observation of microtremors, XI (Matsushiro earthquakes swarm area): Bull. Earthquake Res. Inst., vol. 44, p. 1297-1333.
- Katz, L.J., 1976, Microtremor analysis of local geological conditions: Bull. Seism. Soc. Am., vol. 66, p. 45-60.
- Lacoss, R.T., Kelly, E.M., and Toksöz, M.N., 1969, Estimation of seismic noise structure using arrays: Geophysics, vol. 34, p. 21-38.
- Lysmer, J., and Drake, L.A., 1972, A finite element method for seismology: in Methods in Computational Physics, vol. II, Academic Press.
- Longuet-Higgins, M.S., 1950, A theory of the origin of microseisms: Phil. Trans. Royal Soc. London, Ser. A., vol. 243, p. 1-15.
- Luongo, G., and Rapolla, A., 1973, Seismic noise in Lipari and Vulcano Islands, Southern Thyrrenian Sea, Italy: Geothermics, vol. 2, p. 29-31.
- McEvelly, T.V., and Stauder, W.S.J., 1965, Effect of sedimentary thickness on short-period Rayleigh-wave dispersion: Geophysics, vol. 30, p. 198-203.
- Murphy, A.J., Savino, J., Rynn, J.M.W., Choy, G.L., and McCamy, K., 1972, Observations of long-period (10-100 sec) seismic noise at several worldwide locations: J. Geophys. Res. vol. 77, p. 5042-5049.
- Nicholls, H.R., Rinehart, J.S., 1967, Geophysical study of geyser action in Yellowstone National Park: J. Geophys. Res. vol. 72, p. 4651-4663.
- Oliver, J., and Page, R., 1963, Concurrent storm of long and ultralong period microseisms: Bull. Seism. Soc. Am., vol. 53, p. 15-26.
- Peterson, J., Butler, H.M., Holcomb, L.G., and Hutt, C.R., 1976, The seismic research observatory: Bull. Seism. Soc. Am., Vol. 66, p. 2049-2068.
- Sass, J.H., Lachenbruch, A.H., Monroe, R.J., Greene, G.W., and Moses, T.H., 1971, Heat flow in the western United States: J. Geophys. Res., vol. 76, p. 6376-6413.
- Savino, J., McCamy, K., and Hade, G., 1972, Structures in earth noise beyond twenty seconds--a window for earthquakes: Bull. Seism. Soc. Am., vol. 62, p. 141-176.

- Sorrells, G.G., McDonald, J.A., Der, Z.A., and Herrin, E., 1971, Earth Motion caused by local atmospheric pressure changes: Geophys. J.R. Astr. Soc., vol. 26, p. 83-98.
- Toksoz, M.N., and Lacoss, R.T., 1968, Microseisms: mode structure and sources: Science, vol. 159, p. 872-873.
- Udwadia, F.E., and Trifunac, M.D., 1973, Comparison of earthquake and microtremor ground motions in El Centro, California: Bull. Seismol. Soc. Am., vol. 63, p. 1227-1253.
- Vinnik, L.P., 1971, Origin of longitudinal microseismic waves: Izv., Earth Physics, No. 10, p. 17-30.
- Whiteford, P.C., 1970, Ground movement in Waiotapu geothermal region, New Zealand: Geothermics, (special issue on proceedings of the U.N. Sym. Dev. Util. of Geothermal Resources). 2(part III), p. 478-486.
- Whiteford, P.C., 1975, Studies of the propagation and source locations of geothermal seismic noise, 2nd U.N. Sym. on the Dev. and Use of Geothermal Resource, San Francisco, 1975, p. 1263-1271.
- Whorf, T., 1972, Teleseismic and earth noise monitoring with the Block-Moore Quartz accelerometer. Geophys. J.R. Astr. Soc., vol. 31, p. 205-238.

APPENDIX A  
NOTATION

$A_m(f)$	maximum-likelihood filter coefficients
$b_1$	standardized bandwidth of data window
$ B(k) ^2$	array response function
$c(\tau, \rho)$	cross-correlation function
$c_{mm}(\tau)$	autocorrelation function
$\underline{d}$	directional vector of microseismic field
$E$	expectation
$E(f)$	transfer function of earth medium
$f$	frequency, Hz
$f_H$	cut-off frequency of a low-pass filter
$f_L$	cut-off frequency of a high-pass filter
$f_p$	tape recorder band width
$f_s$	natural frequency of geophone, Hz
$G$	generator constant of geophone
$G(f)$	transfer function of local geology
$h$	thickness of the layer
$H(f)$	transfer function of all system element
$H_A$	amplification of the amplifier
$H_{BW}(f)$	transfer function of Butterworth filter
$H_D(f)$	transfer function of the digitized signal
$H_g(f)$	transfer function of geophone
$H_p(f)$	transfer function of the low-pass filter in the playback system
$H_{RC}(f)$	transfer function of RC low-pass filter

$I$	number of data blocks
$I(f)$	source characteristic of seismic waves
$j$	complex $i$
$\underline{k}$	vector wavenumber
$L$	number of data points in each data block
$m$	index for $m^{\text{th}}$ geophone
$n$	index for $n^{\text{th}}$ geophone
$N$	number of geophones in the array
$P(f, \underline{k})$	frequency-wavenumber spectrum (FKPSD)
$\hat{P}(f, \underline{k})$	the BFM estimate for FKPSD
$\tilde{P}(f, \underline{k})$	the MLM estimate for FKPSD
$\underline{\tilde{q}}$	the inverse of delayed spectral matrix
$\underline{\hat{Q}}$	random noise spectral matrix
$\underline{r}=(x,y)$	vector of geophone location
$R_D$	damping resistance of geophone
$\hat{R}_{mn}(f)$	coherence function estimate between sensors $m$ and $n$
$R_S$	geophone coil resistance
$S(f)$	power spectrum
$\hat{S}_{nn}(f)$	power spectral density estimate for the $n^{\text{th}}$ sensor output
$\bar{S}_{mn}(f)$	averaged cross-power spectrum
$\hat{S}_{mn}(f)$	normalized cross-power spectrum
$\hat{S}$	normalized spectral matrix
$t$	time
$T$	length of each data block in seconds
$\Delta t$	sampling interval of the discrete time series

$U$	energy of the data window
$\underline{V}$	phase velocity vector
$V_n(f)$	velocity spectral density
VCOFS	VCO full scale
$w(\tau)$	data window, lag window, weighting function
$\underline{w}$	weighting vector; $\tilde{\underline{w}}$ = vector of generalized weighting function
$\overset{i}{X}_n(f)$	Fourier transform of the time series $\overset{i}{\phi}_n(t)$
$\beta$	shear-wave velocity
$\zeta$	damping factor of geophone
$\theta$	phase angle
$\nu$	number of degrees of freedom
$\underline{\rho}$	spatial lag
$\sigma(x,y,t,f)$	observed microseisms
$\sigma_i(x,y,t,f)$	intrinsic noise at the site
$\sigma_\ell(x,y,t,f)$	local noise component
$\sigma_m(x,y,t,f)$	microseismic component from the distance source
$\sigma_o$	noise component
$\sigma_1$	component of coherent microseismic signal
$\tau$	temporal lag
$\phi_n(t)$	time series of $n^{\text{th}}$ geophone output
$\overset{i}{\phi}_n(f)$	discrete Fourier transform of time series $\overset{i}{\phi}_n(t)$
$\chi_\nu^2$	chi-square variable with $\nu$ degrees of freedom

APPENDIX B  
DATA PROCESSING NOTES

B.1 Introduction

Three programs have been written in FORTRAN IV for the CDC 7600 machine and RUN 76 compiler at the Lawrence Berkeley Laboratory. Program DUMP reads the digital data from the tape (TAPE 5) generated by the digitizer located in Haviland Hall of the University of California at Berkeley. It then unpacks each 12 bit tape word into one 60 bit CDC word, and writes the data onto another tape (TAPE 3) in binary mode. Program VSD reads the data from TAPE 3, estimates the velocity spectral density of the time series from each geophone output by the method of modified periodogram, and plots the result using both the Graphical Display System (GDS) at the University of California, Berkeley and the line printer (see subroutine YPLOTLG). Program FKPSD reads the array data from TAPE 3 and estimates the frequency wavenumber power spectral density (FKPSD) by either the conventional frequency domain beam forming method (BFM) or the maximum-likelihood method (MLM), and finally presents the result of f-k analysis by the GDS and the line printer.

B.2 Digitization

The field tape is played back at 3 inches per second, i.e., 25 times faster than the recording speed. This shifts the band of 0.25 Hz to 10 Hz up to the band of 6.25 Hz and 250 Hz, which are the limits of our filters. The 12 bit digitizer in Haviland Hall is normally set to digitize 2500 data points per record at the rate of 1000 data points per second (real time), which results in the digitization of 62.5 sec analog data into 2500

digital data in TAPE 5. The header (starting time of each record) and the tailer (the stop time of each record) are written in front of and behind each record in TAPE 5. If the header and/or tailer are not written along with the digital data, the variable NSUP needs to be reset in the subroutine READTAP which is called from the program DUMP.

### B.3 Program DUMP

#### i) Data structure in TAPE 5

As an example, we digitize five records (NSEG = 5) from each geophone output at each channel of the tape recorder, the starting time (header) of each record is denoted by  $T(I)$ ,  $I = 1, 2, 3, 4, 5$ . The sequence of data in TAPE 5 is as follows:

	Ch.2	Ch.3	Ch.4	Ch.5 . . . . Ch.14
T(1)	1	6	11	16 . . . . (NREC-4)
T(2)	2	7	12	17 . . . . (NREC-3)
T(3)	3	8	13	18 . . . . (NREC-2)
T(4)	4	9	14	19 . . . . (NREC-1)
T(5)	5	10	15	20 . . . . <u>(NREC)</u>
				End of file (EOF)

An end-of-file mark is required behind the last record.

#### ii) Input data cards

- (1) The first input data card defines the number of records in the file and the record to be plotted by the GDS.
- (2) An input data card is required for each geophone output. Each card defines the number of the record digitized from the geophone output, the correction factor for geophone polarization, the identification of the geophone output, and the gain factor of the playback system.

iii) Output

The time series of each digital record (2500 data points) with its starting time and stop time are printed, such that the bad record is recognizable and can be deleted. When the number of records in TAPE 5 is large, e.g.,  $NREC > 50$ , it is recommended to dispose the output into microfiche. The unpacked digital data are written in another magnetic tape (TAPE 3) for further processing.

Three types of bad record may occur during this process. The record with a parity error can be skipped by the computer operator under the specification on the control card. In case of parity error, both the number of record in the file (NREC) and the number of the record at that particular geophone output (NSEG) must be reset in input data cards, then, the data in TAPE 5 ought to be reprocessed. Another kind of bad record may happen when the number of words per record does not equal to 502 words (when both the header and tailer are written). If this happens, the translation of this record is skipped by the program and is not written in TAPE 3. The last kind of bad record is the record with clipped amplitude. In this case, the values of clipped data points are listed in front of the printed output data. This kind of bad record shall be deleted by the user under the specifications in the input data cards of programs VSD or FKPSD.

iv) Data structure in TAPE 3

For each record, the unpacked data are written onto TAPE 3 in binary mode by the statements

- (1) WRITE (3) (NDAYS, NHOURS, MIN, NSEC) for the starting time of the record,

- (2) WRITE (3) (NDAYS, NHOURS, MIN, NSEC) for the stop time of the record,
- (3) WRITE (3) (STAT, NB, J) for the record identification, the sequence of record at each geophone output,
- (4) WRITE (3) (NVAL (I), I = 1, 2500) for the values of digital data,
- (5) WRITE (3) (MORE) to indicate the last record of the geophone output if MORE = 0, otherwise MORE = 1.

v) Sample input and output

To unpack and translate one record in TAPE 5 from the site HO.5E sampled from the first record of the 10<sup>th</sup> hour (STAT = H0.5E1001), the sample input cards are:

INPUT DATA CARDS FOR PROGRAM DUMP

```
1 0 0
1 1.0 H0.5E1001 1.0
```

and the printed output is:

RECORD NUMBER 1 CONTAINS 502 WORDS.

J = 1

DIGITISATION STARTING TIME FOR FJ.5E10J1 IS 185 DAYS, 10 HOURS, 0 MINUTES, 4 SECONDS.  
 NOTE THAT DATA WILL START 3 SAMPLES LATER DUE TO THE WRITING OF THE HEADER.  
 STOP TIME 185 DAYS, 10 HOURS, 1 MINUTE, 4 SECONDS.

THE CORRECTIOAL FACTOR FOR POLARIZATION = 1.0 GAIN OF THE PLAY BACK SYSTEM = 1.00

-31	-91	-98	-56	0	79	90	44	0	-25	-22	-4	15	16	-7	-31	-53	-62	-36	0
37	26	2	-7	-31	-83	-108	-48	79	166	-132	16	-85	-90	-36	15	42	32	13	23
59	81	56	-15	-95	-118	-48	55	113	104	50	0	-43	-44	-40	-24	31	55	56	16
-87	-114	-52	31	93	100	66	0	-59	-52	0	63	65	40	9	-23	-71	-95	-64	15
51	104	64	2	-63	-111	-114	-48	59	115	96	36	0	-18	-11	5	10	-5	-18	0
39	71	72	32	-31	-68	-40	27	79	64	0	-79	-68	0	63	76	32	-19	-22	-20
-23	-34	-8	55	90	48	-23	-49	-32	-20	-10	11	31	24	-7	-25	-16	7	27	22
0	-30	-34	-12	11	26	30	25	19	17	17	18	21	26	17	-5	-23	-16	21	52
40	-15	-63	-42	23	74	68	32	-15	-79	-123	-104	-16	95	129	64	-31	-63	-36	-10
-12	-8	23	57	32	-39	-85	-56	0	61	52	8	-26	-26	-6	13	31	53	62	40
0	-31	-4	31	41	25	16	19	19	13	15	31	46	32	-15	-63	-60	-12	23	47
75	80	32	-55	-94	-64	-4	31	45	37	20	6	-3	-23	-29	0	55	72	32	-31
-44	-20	0	-3	-13	-22	-24	-4	23	35	29	12	-15	-45	-48	-28	-2	23	38	24
0	-20	-2	18	20	19	31	47	40	16	-4	0	17	8	-11	0	47	75	60	-12
-22	-29	-8	23	28	0	-19	0	42	33	-11	-29	-2	8	-31	-82	-56	0	67	74
44	-21	-95	-116	-64	23	93	80	0	-103	-112	-40	39	73	80	58	0	-58	-52	0
71	64	0	-58	-60	-12	41	45	36	49	69	40	-31	-81	-64	0	63	72	40	16
19	25	2	-35	-36	0	31	12	-15	-31	-12	42	62	56	0	-46	-8	36	24	-15
-16	23	26	-7	-16	31	72	40	-31	-64	-32	19	51	67	64	32	-15	-54	-53	-28
0	39	62	40	-15	-46	-20	13	10	-2	5	13	-11	-55	-68	-48	-17	-3	7	42
82	64	0	-47	-20	11	4	-5	5	10	-15	-44	-36	-1	21	37	46	28	-7	-29
-16	5	19	35	42	32	8	-1	-2	-7	-17	-6	29	53	36	8	0	15	39	34
10	-15	-47	-69	-52	0	67	106	48	-31	-76	-48	0	38	44	16	-23	-41	-16	15
35	43	48	36	0	-79	-100	-32	55	167	97	64	0	-63	-93	-64	0	63	78	60
34	8	-17	-16	5	18	6	-3	-2	-11	-31	-26	15	51	42	2	-13	15	61	64
24	-21	-35	-25	-12	7	41	55	24	-21	-75	-68	-20	39	62	33	-7	-31	-43	-44
-35	-26	-17	7	43	50	16	-31	-52	-44	-8	39	83	54	64	16	-31	-48	-20	18
31	18	0	-17	-16	0	5	-15	-8	47	103	92	34	-11	-42	-79	-101	-64	0	62
88	88	52	-23	-86	-68	0	71	68	44	45	32	-31	-95	-84	-8	55	64	22	0
-23	-15	-6	0	3	6	14	22	16	-7	-25	-46	-20	19	29	28	8	0	7	38
73	68	16	-29	-44	-41	-24	15	63	80	44	-13	-59	-64	-8	63	101	68	36	-23
-63	-44	15	63	48	0	-44	-26	0	10	-7	-38	-63	-56	-8	51	76	48	-21	-79
-52	-52	0	38	46	29	10	18	26	-15	-77	-80	-16	45	68	58	42	36	16	-27
-55	-36	15	71	50	68	12	-30	-26	-9	-19	-33	-12	29	59	52	24	15	39	38
0	-26	-16	9	26	37	63	59	88	8	-87	-116	-72	0	71	88	48	-27	-66	-36
23	44	16	-19	-22	0	29	34	24	0	-43	-86	-82	-32	31	75	76	24	-39	-77
-76	-32	29	75	77	36	-31	-67	-76	-8	47	36	-11	-36	-28	-2	27	45	28	-23
-54	-24	23	32	4	7	45	55	16	-47	-66	-20	51	91	84	66	44	4	-43	-63
-44	-2	35	52	34	0	-33	-20	23	55	52	24	-31	-66	-62	0	87	100	48	-15
-38	-18	11	13	-11	-21	-18	23	61	52	0	-51	-44	-8	21	47	75	84	32	-35
-76	-53	-24	-8	31	51	112	52	-31	-58	-36	-4	35	65	32	-45	-90	-64	15	67
48	0	-13	-22	-41	-36	0	47	75	80	52	0	-71	-88	-22	47	55	98	48	-23
-74	-56	0	61	52	0	-35	-37	-12	15	47	54	26	0	9	43	45	0	-71	-88
-36	39	57	0	-49	-20	25	0	-63	-72	0	63	64	10	-23	-26	-24	-27	-16	11
41	52	20	-47	-92	-66	0	54	67	71	76	48	-31	-119	-108	-8	83	55	64	64
4	-79	-108	-48	35	52	18	0	11	37	52	44	16	-21	-55	-56	0	57	60	24

0	4	20	29	21	10	17	35	52	44	29	24	12	-15	-39	-28	0	37	36	5
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8	5	9	16	4	-23	-22	31	85	84	32	-43	-55	-56	-48	27	83	92	40	-23
-37	-9	12	15	-22	16	0	-7	4	19	13	3	10	19	16	2	-14	-30	-22	0
4	-27	-44	-20	15	27	36	10	-31	-77	-78	-32	31	75	68	36	4	-17	-63	-84
-48	15	63	65	44	24	0	-27	-35	-25	-14	-9	0	31	63	60	16	-23	-34	-28
-4	31	48	20	7	35	61	16	-55	-82	-50	0	63	123	128	64	0	-61	-77	-64
0	63	72	40	28	45	48	16	-18	-22	0	30	16	-30	-46	-16	27	40	34	20
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0	-30	-38	-16	4	8	23	46	32	-23	-47	-20	11	12	-11	-20	0	47	72	40
-27	-69	-48	0	57	72	52	26	20	27	30	18	-9	-34	-16	35	77	64	16	-29
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-4	23	16	-31	-77	-66	-16	14	10	-7	-4	31	95	139	100	0	-55	-56	-16	55
52	0	-57	-40	35	119	136	64	-39	-95	-80	-24	21	45	56	56	48	32	5	-29
-55	-48	0	71	82	16	-59	-64	-8	35	36	18	8	0	-2	14	41	32	-31	-87
-76	-4	87	145	104	0	-119	-112	-16	63	52	0	-42	-28	0	30	37	26	8	-10
-30	-46	-40	-8	31	62	65	32	-39	-89	-64	0	35	40	32	16	-7	-18	0	27
25	-11	-45	-48	-20	15	47	45	12	-19	-17	7	27	24	14	19	26	10	-15	-28
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88	32	-39	-85	-72	-8	37	26	4	-17	-17	-5	10	31	40	20	6	23	39	16
-31	-77	-76	-24	47	103	126	106	48	-23	-63	-62	-32	31	79	74	16	-37	-57	-52
-44	-16	31	65	64	24	-27	-51	-24	19	32	0	-47	-24	55	88	22	-55	-66	-32
63	104	48	-43	-87	-64	-8	43	55	40	20	22	10	-47	-53	-64	39	119	108	16
-85	-118	-74	0	59	97	90	40	-23	-71	-82	-48	0	42	44	24	0	5	35	43
0	-53	-68	-44	-12	23	75	111	72	0	-41	-24	11	25	2	-27	-16	35	62	34
-15	-47	-44	-16	31	51	24	-11	-28	-24	0	45	78	66	16	-27	-43	-32	0	15
21	14	15	27	34	5	-31	-48	-20	6	14	31	69	72	20	-35	-48	-16	6	3
23	63	48	-21	-65	-24	19	21	6	6	4	-21	-47	-32	11	27	10	3	15	12
-27	-36	27	51	76	0	-87	-84	-16	39	38	12	15	36	38	8	-37	-55	-44	-4
30	40	24	8	7	7	0	0	19	47	62	34	-29	-79	-72	-16	47	65	72	10
-47	-66	-32	27	63	60	32	-11	-38	-22	0	43	48	20	-23	-60	-56	0	79	52
40	-23	-52	-44	-22	11	55	72	34	-23	-36	0	46	41	18	2	-27	-59	-40	23
67	64	22	-15	-63	-56	0	69	48	-19	-63	-57	-12	31	71	83	48	-47	-127	-120
-32	63	53	64	12	-31	-61	-48	0	45	40	4	-12	-1	17	26	15	15	31	44
20	-15	-16	27	63	48	0	-79	-98	-48	35	59	96	40	-23	-46	-16	39	60	32
-45	-101	-76	0	103	108	48	-27	-70	-56	-10	23	22	-7	-25	-14	-1	-5	-11	11
34	39	41	37	16	-15	-35	-24	0	18	12	6	15	25	17	0	-22	-24	-4	45
89	76	16	-22	-20	-21	-53	-79	-52	15	75	74	24	-39	-90	-84	-40	11	46	57
52	37	0	-21	-68	-32	15	35	38	20	1	0	25	55	59	32	1	3	31	55
48	18	0	-8	-5	0	1	15	55	84	40	-23	-22	0	21	17	13	11	0	-42
-58	-8	71	65	24	-39	-59	-51	-32	15	55	32	-31	-34	19	41	4	-27	-20	0
31	42	16	-31	-60	-36	-10	-11	-12	15	54	40	-11	-22	0	0	-47	-75	-34	27
42	2	-45	-56	-38	-26	-12	23	57	52	4	-38	-46	-32	-8	11	12	-11	-24	0
47	75	68	40	0	-47	-48	27	111	104	32	-19	-5	26	36	16	-3	0	25	57
58	32	0	-17	-25	-4	51	64	0	-59	-34	23	53	36	16	21	16	-31	-63	-32
21	43	36	24	8	-15	-33	-22	15	46	32	-7	-26	-12	11	31	32	25	24	18
-6	-24	0	47	48	4	-17	0	21	0	-45	-53	-24	23	39	8	-35	-48	-4	17
-23	-66	-42	0	35	40	42	32	4	-35	-79	-99	-32	63	108	56	-27	-69	-66	-46
-8	55	59	52	-39	-73	-24	25	8	-21	-32	23	79	86	32	-27	-63	-74	-42	39
109	96	24	-25	-33	-13	5	23	45	63	48	4	-12	11	30	20	12	23	35	24
-13	-35	-8	43	58	24	-22	-36	-12	22	47	55	44	0	-53	-68	-24	47	90	56
-23	-57	-16	43	66	48	16	-23	-62	-90	-72	0	70	72	32	-15	-62	-102	-98	-48
31	89	111	56	32	-62	-116	-80	0	52	34	0	-14	-5	14	43	55	40	0	-61

-77	-32	31	62	40	6	-9	-10	-9	-10	-4	18	55	79	48	-31	-75	-56	-10	39
83	52	56	26	27	24	-15	-71	-76	-16	59	78	50	16	-15	-47	-58	-51	-20	31
79	66	18	-15	-37	-47	-36	-4	17	0	-35	-42	-4	41	63	70	68	24	-63	-121
-104	-32	47	82	74	48	20	0	-25	-28	0	35	32	0	-8	23	57	56	16	-31
-60	-36	21	65	73	44	13	-15	-47	-40	0	63	64	32	-13	-55	-68	-32	47	115
106	24	-75	-124	-88	0	75	60	20	-35	-34	0	27	26	24	8	-37	-66	-32	31
67	64	52	32	-31	-107	-114	-24	55	124	64	-31	-62	-38	-28	-34	-8	43	40	-23
-62	-16	43	56	18	-31	-71	-83	-48	31	111	128	64	-47	-97	-64	0	63	73	34
4	1	7	3	-7	-9	0	8	6	10	31	57	58	32	-10	-37	-41	-30	-17	0
55	95	80	16	-23	-25	-19	-25	-34	-43	-41	-20	11	45	61	42	2	-11	13	37
18	-15	-16	35	78	48	-31	-55	-84	-10	51	78	78	40	-31	-72	-32	61	57	40
-31	-40	27	85	76	16	-47	-61	-40	-4	39	73	56	0	-63	-85	-60	-16	18	28
9	-15	-18	15	47	44	20	-5	-31	-59	-68	-42	13	51	48	16	-27	-55	-60	-24
31	61	44	12	0	7	9	-5	-35	-67	-48	23	91	90	32	-43	-72	-36	31	55
97	32	-55	-78	-40	11	55	55	114	64	-25	-72	-40	15	50	59	44	8	-27	-28
0	20	0	-4	31	60	24	-31	-40	15	79	81	8	-79	-104	-40	37	64	36	-13
-45	-28	11	16	-27	-54	-20	35	56	4	-55	-48	0	59	56	20	-6	-2	19	33

#### B.4 Program VSD

The program VSD estimates the velocity spectral density of each geophone output stored in TAPE 3. Each record of 2500 data points is divided into several data blocks of ( $2^{**}NPOW$ ) points per block. The maximum number of data point per block is 512 which is set by the dimension of arrays in the program. Each time series is multiplied by a 10% cosine bell taper before FFT. The power spectral density estimate of each geophone output is obtained by the method of periodogram described in section 4.3.2. This program also calculates 90% confidence limits of the estimated VSD.

##### i) Input data cards

- (1) The first data card specifies the option for punched data deck of VSD, the NPOW to define the data point per block, and the record length in second.
- (2) An input data card is required for each geophone output. This card specifies the termination of data processing, the option for VSD plot by the GDS, the identification of each geophone output, the bad records to be deleted, and the parameters of acquisition and playback systems.

##### ii) Output

The VSD estimate of each geophone output is plotted by subroutine YPLOTLG in a log-linear scale and GDS in a log-log scale along with 90% confidence limits. The estimated VSD are punched on decks for further processing, if the variable IPUNCH is set to be 1.

iii) Sample input and output

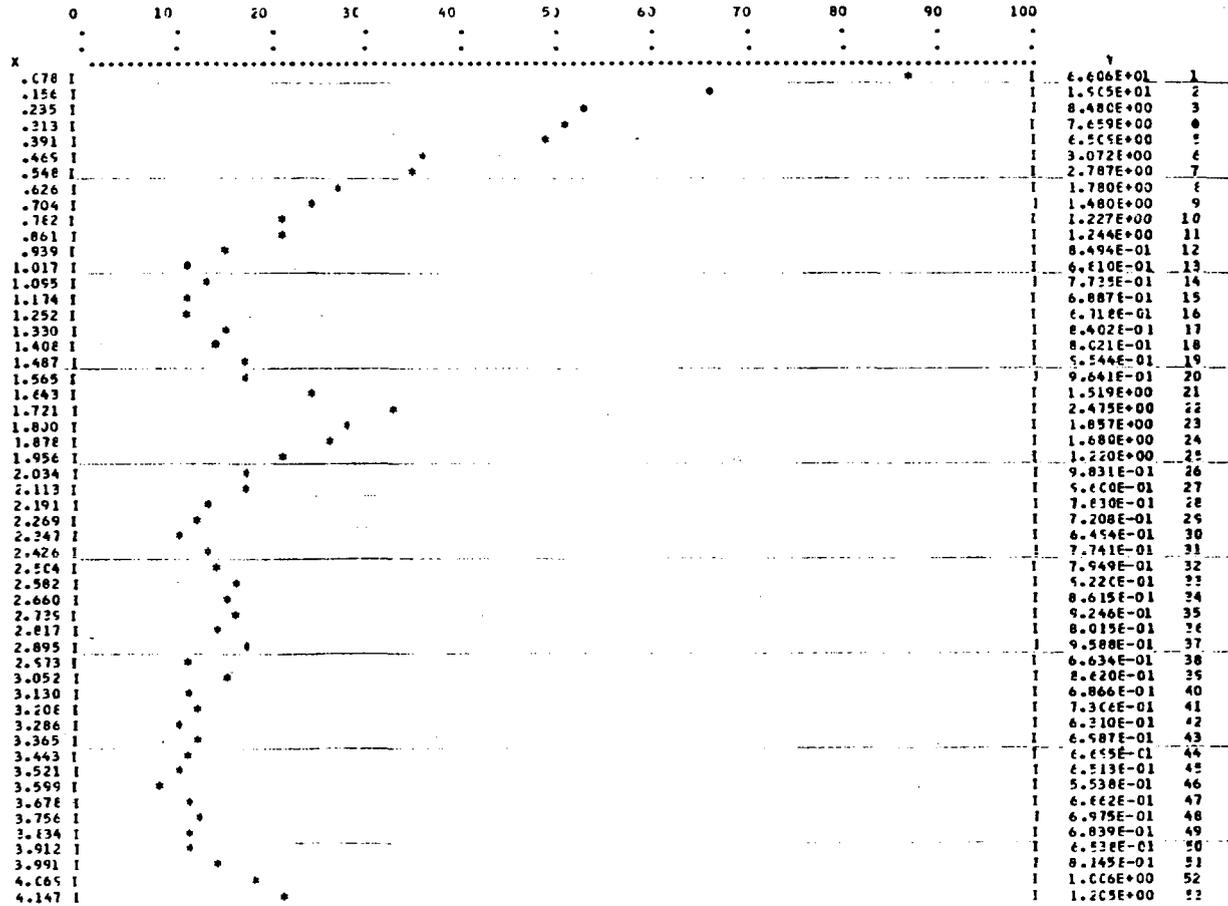
To estimate VSD of the digital data obtained from both site HS and site A3.7N, we need the digital data in magnetic tape (TAPE 3) and the following input data cards

INPUT DATA CARDS FOR PROGRAM VSD

0	9 62.5				
10102	HS	3.40999.00	0.30102.00	30.00	5.00310.00
00102	A3.7N	3.40999.00	0.30114.00	30.00	5.00310.00

Portion of the printed output (VSD of site HS from DC to 8.99 Hz) is presented as follows:

YPLCT..... HS  
 16\* 1 IE\* 256 IK\* 1 FPAX= 1.327E+02 PPIA= 3.168E-01



4.225 I	I	1.488E+00	54
4.304 I	I	1.461E+00	55
4.382 I	I	1.360E+00	56
4.460 I	I	1.031E+00	57
4.538 I	I	8.020E-01	58
4.617 I	I	7.212E-01	59
4.695 I	I	6.031E-01	60
4.773 I	I	6.713E-01	61
4.851 I	I	5.064E-01	62
4.930 I	I	4.736E-01	63
5.008 I	I	4.128E-01	64
5.086 I	I	3.239E-01	65
5.164 I	I	3.995E-01	66
5.243 I	I	3.205E-01	67
5.321 I	I	3.277E-01	68
5.399 I*	I	3.458E-01	69
5.477 I	I	3.349E-01	70
5.556 I	I	3.168E-01	71
5.634 I*	I	3.371E-01	72
5.712 I	I	2.725E-01	73
5.790 I	I	3.872E-01	74
5.868 I	I	4.531E-01	75
5.947 I	I	4.731E-01	76
6.025 I	I	5.591E-01	77
6.103 I	I	4.715E-01	78
6.181 I	I	6.296E-01	79
6.260 I	I	4.525E-01	80
6.338 I	I	5.480E-01	81
6.416 I	I	5.636E-01	82
6.494 I	I	5.031E-01	83
6.573 I	I	6.066E-01	84
6.651 I	I	6.552E-01	85
6.729 I	I	6.335E-01	86
6.807 I	I	6.228E-01	87
6.886 I	I	6.061E-01	88
6.964 I	I	5.554E-01	89
7.042 I	I	6.728E-01	90
7.120 I	I	7.389E-01	91
7.199 I	I	7.055E-01	92
7.277 I	I	7.508E-01	93
7.355 I	I	7.052E-01	94
7.433 I	I	6.555E-01	95
7.512 I	I	6.572E-01	96
7.590 I	I	6.859E-01	97
7.668 I	I	8.258E-01	98
7.746 I	I	7.716E-01	99
7.825 I	I	5.022E-01	100
7.903 I	I	7.265E-01	101
7.981 I	I	5.144E-01	102
8.059 I	I	6.469E-01	103
8.138 I	I	1.039E+00	104
8.216 I	I	1.027E+00	105
8.294 I	I	5.556E-01	106
8.372 I	I	1.146E+00	107
8.451 I	I	1.053E+00	108
8.529 I	I	9.942E-01	109
8.607 I	I	1.025E+00	110
8.685 I	I	1.055E+00	111
8.764 I	I	5.035E-01	112
8.842 I	I	9.840E-01	113
8.920 I	I	5.305E-01	114
8.998 I	I	1.031E+00	115

### B.5 Program FKPSD

The program FKPSD estimates FKPSD of the array geophone outputs stored in TAPE 3. This program is set up to process the data obtained by the maximum of 12 array elements and the maximum of 512 data points per block. Each data block is multiplied by a Hamming window to reduce the leakage effect in the Fourier transform. The FKPSD can be estimated by either BFM described in Section 5.4.2, or MLM described in Section 5.4.3. The variable `ICONVEN = 1` directs the program through the conventional BFM, and the `ICONVEN = 0` directs the program through MLM. At a frequency component the FKPSD are estimated at each of 41 x 41 grid points in a two-dimensional wavenumber space. In addition to estimate FKPSD, this program also provides the options to calculate the coherence between two geophone outputs with its confidence limits and the VSD by specifying `ICOHER = 1` and `NVSDPL ≥ 1`.

We have found that the starting time of each digital record can drift a few msec to 250 msec from the designated starting time on the front panel of the digitizer; this results from inconstant acceleration of the tape drive before recording the digital record on tape. In order to correct the effect of random drifting, the digital data must be plotted by the GDS and compared with the analog data before the estimation of FKPSD.

#### i) Input data cards

- (1) The first card specifies the record length, the variable `NPOW` to define the number of data points per block, the number of data points per record, the maximum number of data blocks to be processed, and the number of records per sensor output

of which starting-time correction is required.

- (2) The second card defines the contour levels in the f-k plot.
- (3) The third card specifies the options for BFM or MLM, for f-k plot by GDS, and for VSD plot.
- (4) A card specifies the termination of the data processing, the number of array elements, the parameters of the playback system, the station identification, the minimum geophone spacing in the array, the scale of output wavenumber, and the options for coherence calculation and FKPSD estimations. This card is required for each array data set.
- (5) Specification of the desired frequency component for estimating FKPSD. This card is required for each array data set.
- (6) Correction time for each record. This card is required for each array data set.
- (7) The coordinate of array elements. This card is required for each array data set.
- (8) This card specifies the identifications of each geophone output, the parameters of geophone, and the option to skip the bad record. A card is required for each geophone output.

ii) Output

The estimated FKPSD in the wavenumber space are normalized with respect to the peak-valued FKPSD at each desired frequency component. The normalized FKPSD are multiplied by -10 to save space of the sign and the decimal point, then, printed in integer format, contoured by symbols, and plotted by GDS. The maximum FKPSD over the frequency band are plotted by the subroutine YPLOT.

iii) Sample input and output

There are two array data sets, E5.9W04 and E5.9W05, stored in TAPE 3. This example skips the array data at E5.9W04 and estimates the FKPSD of the array data at E5.9W05 in four frequency components, i.e. the 7th, 8th, 9th and 10th discrete frequency components. The sequential number of discrete frequency components is defined by the statements FK2.153 and FK2.154 of program FKPSD. The starting-time correction for each record is as follows:

E5.0W0501(1) = 0.0 sec	,	E5.9W0501(2) = 0.04 sec
E5.0W0502(1) = 0.0 sec	,	E5.9W0502(2) = 0.04 sec
E5.9W0503(1) = 0.0 sec	,	E5.9W0503(2) = 0.02 sec
E5.9W0504(1) = 0.0 sec	,	E5.9W0504(2) = 0.0 sec
E5.9W0505(1) = 0.0 sec	,	E5.9W0505(2) = 0.04 sec
E5.9W0506(1) = 0.0 sec	,	E5.9W0506(2) = 0.04 sec
E5.9W0507(1) = 0.0 sec	,	E5.9W0507(2) = 0.04 sec
E5.9W0508(1) = 0.0 sec	,	E5.9W0508(2) = 0.0 sec
E5.9W0509(1) = 0.0 sec	,	E5.9W0509(2) = 0.0 sec
E5.9W0510(1) = 0.0 sec	,	E5.9W0510(2) = 0.04 sec
E5.9W0511(1) = 0.0 sec	,	E5.9W0511(2) = 0.04 sec
E5.9W0512(1) = 0.0 sec	,	E5.9W0512(2) = 0.04 sec.

We take only the first record of each geophone output to estimate FKPSD and set the maximum number of data blocks taken from each channel to be 24, and the number of data points per block to be  $2^6 = 64$ . The sample input data cards are following:

```
*****
*
*           INPUT DATA CARDS FOR PROGRAM FKPSD
*
*****

E2.50      6 2500      24      2
5 -1.0 -3.0 -6.0 -9.0-12.0
0      0      0
1      0      1      1      12      C      0      3 10.0 0.50.007 E5.9W04
20     1      2      3      4      5      6      7      8      9      10     11     12     13     14     15
16     17     18     19     20
C.1200.0000.1200.0000.1200.0000.1200.0000.1200.0000.1200.0000.0C00.0000.1200.0000.1200.000
C.1200.0000.0800.0000.1200.0000.1200.0000
-0.0190 0.0110 C.0000 0.0250 C.0150 0.0110 -0.0108 0.0063 0.0000 0.0177
0.0108 0.0063 -0.0153 -0.0089 0.0000 -0.0125 0.0153 -0.0089 -0.0217 -0.0125
C.0000 -0.0219 C.0217 -0.0125
E5.9W0401 2 1 3      1.0 3.4999.0 0.30
E5.9W0402 2 1 3      1.0 1.35 3.00 0.60
E5.9W0403 2 1 3      1.0 1.35 3.00 0.60
E5.9W0404 2 1 3      1.0 3.40999.0 0.30
E5.9W0405 2 1 3      1.0 1.35 3.00 0.60
E5.9W0406 2 1 3      1.0 3.40999.0 0.30
E5.9W0407 2 1 3      1.0 3.40999.0 0.30
E5.9W0408 3 1 2 4      1.0 4.30999.0 0.60
E5.9W0409 3 1 2 4      1.0 1.35999.0 0.60
E5.9W0410 2 1 3      1.0 3.40999.0 0.30
E5.9W0411 2 1 3      1.0 3.40999.0 0.30
E5.9W0412 2 1 3      1.0 1.35 3.50 0.60
0      0      1      0      12      0      0      3 10.0 0.50.007 E5.9W05
4      7      8      9      10
C.0000.0400.0000.0400.0000.0200.0000.0000.0000.0400.0000.0400.0000.0400.0000.0000.000
C.0000.0000.0000.0400.0000.0400.0000.040
-0.0190 0.0110 C.0000 0.0250 C.0150 0.0110 -0.0108 0.0063 0.0000 0.0177
0.0108 0.0063 -0.0153 -0.0089 0.0000 -0.0125 0.0153 -0.0089 -0.0217 -0.0125
C.0000 -0.0219 C.0217 -0.0125
E5.9W0501 2 2 3      1.0 3.4999.0 0.30
E5.9W0502 3 2 3 4      1.0 1.35 3.00 0.60
E5.9W0503 2 2 3      1.0 1.35 3.00 0.60
E5.9W0504 2 2 3      1.0 3.40999.0 0.30
E5.9W0505 2 2 3      1.0 1.35 3.00 0.60
E5.9W0506 2 2 3      1.0 3.40999.0 0.30
E5.9W0507 2 2 3      1.0 3.40999.0 0.30
E5.9W0508 2 2 3      1.0 4.30999.0 0.60
E5.9W0509 2 2 3      1.0 1.35999.0 0.60
E5.9W0510 2 2 3      1.0 3.40999.0 0.30
E5.9W0511 2 2 3      1.0 3.40999.0 0.30
E5.9W0512 2 2 3      1.0 1.35 3.50 0.60
114.0 5.0 30.0
*****
```

The printed output for the 9th Frequency component at E5.9W05 is



B.6 Program listing

```

*****
*
*                               PROGRAM DUMP
*
*****

DUMP,7,55.471209,LIAH
DISKHGG,14000.
RUN76,SC,NL77777.
FETCHMT,TAPE5,05,P4,32264,1F.
FETCHPS,BKYJLIB,TEMP,GDS.
FETCHPS,BKYLGOB,GDSCC,GDSCC.
COPYI,TEMP,LIB.
RETURN,TEMP.
LINK,X,F=LGO,P=LIB,P=BKYIC.
STAGE,TAPE3,08,W,08665.
ENTERMT,TAPE3,08665.
RETURN,TAPE3.
EXIT.
DUMP,0.
FIN.
GDSCC.
DISPOSE,OUTPUT,MF.
RETURN,LGO,LIB.
RETURN,TAPE99,GDSCC.
PROGRAM DUMP(INPUT,OUTPLT,PUNCH,PUNCHB,TAPE1=PUNCHB,TAPE3,TAPE5,TADUMP.2
1PE99,TAPE92,TAPE93,TAPE94) DUMP.3
C*****DUMP.4
C DUMP.5
C DUMP.6
C PROGRAM DUMP DUMP.7
C SUBROUTINE REACTAP DUMP.8
C TRANS DUMP.9
C UNPK DUMP.10
C TIME DUMP.11
C APLCT DUMP.12
C DUMP.13
C DUMP.14
C DUMP.15
C DUMP.16
C DUMP.17
C.....THIS PROGRAM READS THE TAPE (TAPE 5) WRITTEN BY THE DIGITIZER
C OF HAVILAND HALL, TRANSLATES THE DATA, AND UN-PACKS EACH WORD
C INTO 60 BITS WORD, AND PLCTS THE TIME SERIES BY CALCOMP UNDER
C REQUESTED. DUMP.18
C DUMP.19
C DUMP.20
C DUMP.21
C.....IPRINT, WHEN IPRINT = J THAT RECORD WILL BE PRINTED IN OCTAL FORM. DUMP.22
C.....J = THE SEQUENCE NUMBER OF RECORD IN TAPE 5. DUMP.23
C.....NB = THE SEQUENCE NUMBER OF EACH GEOPHONE OUTPUT. DUMP.24
C.....NNPL = THE TOTAL NUMBER OF NPL. DUMP.25
C.....NPL = J, THE JTH TIME SERIES WILL BE PLOTTED. DUMP.26
C.....NREC = THE TOTAL NUMBER OF RECORD DIGITIZED ON TAPE 5. DUMP.27
C.....NSEG = THE NUMBER OF RECORD FROM EACH GEOPHONE OUTPUT. DUMP.28
C.....NSKIP = THE SEQUENCE NUMBER IN THE TAPE TO BE SKIPPED. DUMP.29
C.....PB = THE GAIN FACTOR OF THE PLY-BACK SYSTEM. DUMP.30
C.....PULA = THE CORRECTION FACTOR OF POLARIZATION OF THE SEISMOMETER AND DUMP.31
C..... D/OK AMPLIFIER. DUMP.32
C.....STAT = STATION IDENTIFICATION. DUMP.33
C DUMP.34
C.....THE INPUT DATA CARDS ARE AS FOLLOWING DUMP.35
C (1) NREC,NNPL,NSKIP(313) DUMP.36
C (2) NSEG,PULA,STAT,PB,IPRINT(12,F8.2,A10,F5.2,I3) FOR EACH GEO- DUMP.37

```

```

C          PHONE OUTPUT.                                DUMP.38
C                                                    DUMP.39
C.....AUTHOR    ALFRED LIANG-CHI LIAN
C                ENGINEERING GEOSCIENCE
C                UNIVERSITY OF CALIFORNIA, BERKELEY
C.....DATE      SEPTEMBER 1977
C*****
DIMENSION SPECS(12)                                DUMP.40
DIMENSION NPL(55)                                  DUMP.41
COMMON /BLK1/N1(512),LEN,MCATA                     DUMP.42
COMMON /BLK2/NB,NSKIP                               DUMP.43
COMMON /BLK3/N(2500),NH(12),NT(12)                 DUMP.44
                                                    DUMP.45
                                                    DUMP.46
                                                    DUMP.47
C.....GVERRIDES INSUFFICIENT DATA READ ERROR ON BKVID. DUMP.48
CALL FET(5,04CC 00C0 00C0 0000 0000B,8*64)         DUMP.49
                                                    DUMP.50
                                                    DUMP.51
SPECS(12)=99.                                       DUMP.52
MDATA=2500                                           DUMP.53
IDEC=1                                               DUMP.54
J=0                                                  DUMP.55
NB=0                                                 DUMP.56
READ 9,NREC,NNPL,NSKIP                              DUMP.57
9  FORMAT (3I3)                                       DUMP.58
99 READ 10,NSEG,POLA,STAT,PB,IPRINT                 DUMP.59
10  FORMAT (12,F8.2,A10,F5.2,I3)                     DUMP.60
KPB=1.0/PB                                          DUMP.61
98  J=J+1                                             DUMP.62
     NB=NB+1                                          DUMP.63
     CALL REACTAP(NREC,J,IPRINT,NNPL,SPECS,AVAL)     DUMP.64
     IF (INVAL.EQ.0) GO TO 51                         DUMP.65
20  CALL TRANS (MC,SAMP,IDEC,STAT,J)                 DUMP.66
                                                    DUMP.67
DO 5 I=1,MDATA                                       DUMP.68
N(I)=N(I)*POLA                                       DUMP.69
IF (IABS(N(I)).LE.2040) GO TO 5                       DUMP.70
PRINT 6,I,N(I)                                       DUMP.71
6  FORMAT (1H0,1X,*N( *,14,*) = *,16)               DUMP.72
5  CONTINUE                                          DUMP.73
     IF (J.LE.NSKIP) GO TO 4                          DUMP.74
DO 55 I=1,MDATA                                       DUMP.75
N(I)=N(I)*RPB                                       DUMP.76
55  CONTINUE                                          DUMP.77
     PRINT 800,POLA,PB                                DUMP.78
800 FORMAT (1H0,1X,*THE CORRECTIONAL FACTOR FOR POLARIZATION = *,F5.1, DUMP.79
15X,*GAIN OF THE PLAY BACK SYSTEM = *,F5.2,/)       DUMP.80
PRINT 2,(N(I),I=1,2500)                              DUMP.81
2  FORMAT (1X,516,3X,516,3X,516,3X,516)             DUMP.82
4  WRITE (3) (STAT,NB,J)                             DUMP.83
     WRITE (3) (N(I),I=1,2500)                       DUMP.84
C  PUNCH 3,(N(I),I=1,2500)                           DUMP.85
C 3  FORMAT (2004)                                    DUMP.86
DO 40 NL=1,NNPL                                       DUMP.87
IF (J.EQ.NPL(NL)) GO TO 41                            DUMP.88
40  CONTINUE                                          DUMP.89
     GO TO 7                                           DUMP.90
41  CALL APL0T(STAT,NB)                               DUMP.91
     IF (J.GE.NREC) GO TO 7                            DUMP.92
     CALL NXTFRM(SPECS)                               DUMP.93
7  CONTINUE                                          DUMP.94
49  MORE=1                                           DUMP.95

```

```
50 CONTINUE                                DUMP.96
    IF (NB.GE.NSEG) MCRE=C                  DUMP.97
    IF (NB.GE.NSEG) NB=0                    DUMP.98
    WRITE (3) (MCRE)                        DUMP.99
    GO TO 52                                DUMP.100
51 PRINT 53,STAT,NB,J                      DUMP.101
53 FORMAT (1H0,5X,*THE PROGRAM SKIPS THE RECORD OF *,A10,12,* AS REQU)DUMP.102
    TESTED.*,* J = *,13,////////////////////DUMP.103
    IF (J.GE.NREC) GC TC 6C                 DUMP.104
    IF (NB.LT.NSEG) GO TO 56                DUMP.105
    PRINT 54                                 DUMP.106
54 FORMAT (1H0,*AAAAA*)                   DUMP.107
    MORE=0                                   DUMP.108
    NB=0                                     DUMP.109
52 IF (J.GE.NREC) GO TO 6C                 DUMP.110
    IF (MORE.EQ.1) GO TO 98                 DUMP.111
    GO TO 99                                 DUMP.112
60 CONTINUE                                DUMP.113
    CALL GDSEND(SPECS)                      DUMP.114
70 STOP                                    DUMP.115
    END                                      DUMP.116
```

```
SUBROUTINE APLCT(LABEL2,NB)                APL0T.2
COMMON /BLK3/N(2500)                       APL0T.3
DIMENSION SPECS(30),X(2500),Y(2500),BUFY(500),BUFY(500) APL0T.4
DIMENSION GXSPEC(4),XTICK(3,2),YTICK(3,2)  APL0T.5
DIMENSION LINE(2)                          APL0T.6
SPECS(1)=0.0                               APL0T.7
SPECS(2)=1.0                               APL0T.8
SPECS(3)=2500.0                             APL0T.9
SPECS(4)=1.0                               APL0T.10
SPECS(5)=2048.0                             APL0T.11
SPECS(6)=-2048.0                           APL0T.12
SPECS(7)=46.88                             APL0T.13
SPECS(8)=10.0                              APL0T.14
SPECS(11)=1.0                              APL0T.15
SPECS(12)=99.0                             APL0T.16
SPECS(13)=2500.0                           APL0T.17
SPECS(14)=1.0                              APL0T.18
SPECS(15)=1.0                              APL0T.19
DO 1 I=1,2500                               APL0T.20
  X(I)=I                                    APL0T.21
  Y(I)=FLOAT(N(I))                         APL0T.22
1 CONTINUE                                  APL0T.23
CALL PFLILI(X,Y,BUFY,BUFY,SPECS)          APL0T.24
LINE(1)=LABEL2                             APL0T.25
LINE(2)=0                                   APL0T.26
RULE =3.0                                   APL0T.27
SPECS(17)=0.2                              APL0T.28
SPECS(18)=0.2                              APL0T.29
SPECS(19)=0.0                              APL0T.30
SPECS(20)=0.0                              APL0T.31
SPECS(21)=1.0                              APL0T.32
SPECS(22)=45.0                             APL0T.33
SPECS(23)=10.5                             APL0T.34
CALL TITLEG(RULE,LINE,SPECS)              APL0T.35
SPECS(22)=46.0                             APL0T.36
SPECS(28)=0.0                              APL0T.37
VALUE=FLOAT(NB)                            APL0T.38
CALL DECVAL(RULE,VALUE,SPECS)            APL0T.39
SPECS(7)=45.0                              APL0T.40
```



```

C      IF (NNPL.EQ.0) GO TO 72
C      CALL GDSND(SPELS)
C 72  STOP
      END

```

```

READTAP.
READTAP.
READTAP.
READTAP.

```

```

SUBROUTINE TIME (L,NDAYS,NHCURS,MIN,NSEC)
DIMENSION L(12)
NDAYS=(L(1).AND.1)*200+(L(2).AND.4)*25+(L(2).AND.3)*40+(L(3).AND.6
1)*5+(L(3).AND.1)*8+(L(4).AND.7)
NHCURS=(L(5).AND.1)*2C+(L(6).AND.4)/4*10+(L(6).AND.3)*4+(L(7).AND.6
1D.6)/2
MIN=(L(8).AND.7)*10+(L(9).AND.7)*2+(L(10).AND.4)/4
NSEC=(L(10).AND.1)*40+(L(11).AND.6)*5+(L(11).AND.1)*8+(L(12).AND.7
1)
RETURN
END

```

```

TIME.2
TIME.3
TIME.4
TIME.5
TIME.6
TIME.7
TIME.8
TIME.9
TIME.10
TIME.11
TIME.12

```

```

SUBROUTINE TRANS (MC,SAMP,IDEQ,STAT,J)

```

```

TRANS.2
TRANS.3

```

```

COMMENT... THE FIRST DATA CARD (12,FB.2,A10,F10.0) INDICATES THE HEADER-TRANS.4
C OPTION,THE SAMPLE INTERVAL IN SEC.,THE STATION ID,AND THE DEFAULT OPTTRANS.5
C DEFAULT OPTION IS ALWAYS LEFT BLANK THE PROGRAM TERMINATES IF ANYTHINGTRANS.6
CHED IN THIS FIELD. THE HEADER-TAILER OPTION CODE IS AS FOLLOWS 1=NO HEATTRANS.7
C TAILER 2=HEADER AND TAILER 3=HEADER ONLY 4=TAILER ONLY. THE ITRANS.8
C TA ARE THE WRITTEN OUTPUT OF THE PROGRAM DUMPTAPE FOR THE HAV.HALL DIGTRANS.9

```

```

COMMON /BLK1/N1(512),LEN,MDATA
COMMON /BLK2/NB,NSKIP
COMMON /BLK3/N(2500),NH(12),NT(12)
IF (MDATA .LE. 1) MDATA =2500
MD3 = MDATA-3
MD2 = MDATA-2
C      READ 2, NCHEK,SAMP,STAT,DEFAULT
C 2  FORMAT(12,FB.4,A10,F10.0)
C
      NCHEK=2
      SAMP=0.0125
      DEFAULT=0.0
C
      IF (DEFAULT.L1..0001) GO TO 110
      PRINT 12
      12 FORMAT (* PROBLEM WITH NUMBER OF CARDS IN THE INPUT DECK.  PROGRTRANS.26
      1AM TERMINATED. *)
      STOP
      110 CONTINUE
      IF (NCHEK-2) 10,20,30
      10 CONTINUE
      II=0
      DO 3 I=1,500
      DO 2 JW=1,5
      II=II+1
      N(II)=SHIFT(N1(I),12*JW).AND.7777B
      2 CONTINUE
      3 CONTINUE
C      READ (4,3) (N(I),I=1,MDATA)
C 3  FORMAT(2004)
      PRINT 11,STAT,NB
      11 FORMAT(1HU,*NO HEADER-TAILER INDICATED BY CONTROL CARD FOR *,A10,I)TRANS.44

```

```

TRANS.10
TRANS.11
TRANS.12
TRANS.13
TRANS.14
TRANS.15
TRANS.16
TRANS.17
TRANS.18
TRANS.19
TRANS.20
TRANS.21
TRANS.22
TRANS.23
TRANS.24
TRANS.25
TRANS.26
TRANS.27
TRANS.28
TRANS.29
TRANS.30
TRANS.31
TRANS.32
TRANS.33
TRANS.34
TRANS.35
TRANS.36
TRANS.37
TRANS.38
TRANS.39
TRANS.40
TRANS.41
TRANS.42
TRANS.43
TRANS.44

```

```
12) TRANS.45
  GO TO 40 TRANS.46
20 CONTINUE TRANS.47
C READ (4,4) ((NH(I),I=1,12),(N(IF),IF=1,17)) TRANS.48
C 4 FJRMAT(1201,1704) TRANS.49
C READ (4,3) (N(IF),IH=18,MD3) TRANS.50
C READ (4,5) (N(IU),IU=MD2,MDATA),(NT(JE),JE=1,12)) TRANS.51
C 5 FORMAT(304,12C1) TRANS.52
  CALL UNPK(IWORD) TRANS.53
C PRINT 22,IWORD TRANS.54
C 22 FORMAT (1H0,35X,*IWORD = *,I5) TRANS.55
21 DO 23 J2=1,8 TRANS.56
  NT(J2)=SHIFT(N1(IWORD),3*J2+36).AND.7B TRANS.57
23 CONTINUE TRANS.58
  DO 24 J2=9,12 TRANS.59
  NT(J2)=SHIFT(N1(LEN),3*(J2-8)).AND.7B TRANS.60
24 CONTINUE TRANS.61
  CALL TIME(NH,NDAYS,NHOURS,MIN,NSEC) TRANS.62
  WRITE (3) (NDAYS,NHCURS,MIN,NSEC) TRANS.63
  PRINT 25,J TRANS.64
25 FORMAT (1H0,* J = *,I5) TRANS.65
50 CONTINUE TRANS.66
  PRINT 6,STAT,NB,NDAYS,NHCURS,MIN,NSEC TRANS.67
  6 FORMAT(1H0,* DIGITISATION STARTING TIME FOR *,A10,I2,* IS *,I3,* DTTRANS.72
  1AYS, *,I2,* HOURS, *,I2,* MINUTES, *,I2,* SECONDS.*/,* NOTE THAT TRANS.73
  1DATA WILL START 3 SAMPLES LATER DUE TO THE WRITING OF THE HEADER.*TRANS.74
2) TRANS.75
31 CALL TIME(NT,NDAYS,NHOURS,MIN,NSEC) TRANS.76
  WRITE (5) (NDAYS,NHOURS,MIN,NSEC) TRANS.77
51 CONTINUE TRANS.78
  PRINT 7,NDAYS,NHCURS,MIN,NSEC TRANS.79
  7 FORMAT(* STOP TIME*,I4,* DAYS, *,I2,* HOURS, *,I2,* MINUTES, *,I2,TRANS.81
  1* SECONDS.*////) TRANS.82
  GO TO 40 TRANS.83
30 CONTINUE TRANS.84
  IF (NCHEK.EQ.4) GO TO 35 TRANS.85
C READ (4,4) ((NH(IY),IY=1,12),(N(IR),IR=1,17)) TRANS.86
C READ (4,3) (N(KJ),KJ=18,MDATA) TRANS.87
  CALL UNPK(IWORD) TRANS.88
  CALL TIME(NH,NDAYS,NHCURS,MIN,NSEC) TRANS.89
  PRINT 6,STAT,NDAYS,NHCURS,MIN,NSEC TRANS.90
  GO TO 40 TRANS.91
35 CONTINUE TRANS.92
C READ (4,3) (N(IW),IW=1,NCDATA) TRANS.93
C READ (4,9) (NT(KN),KN=1,12) TRANS.94
C 9 FORMAT(1201) TRANS.95
  II=0 TRANS.96
  DO 103 I=1,500 TRANS.97
  DO 102 Jw=1,5 TRANS.98
  II=II+1 TRANS.99
  N(II)=SHIFT(N1(I),12*Jw).AND.7777B TRANS.100
102 CONTINUE TRANS.101
103 CONTINUE TRANS.102
```



```
*****  
*  
*                               PROGRAM VSD                               *  
*  
*****
```

```
VSD,7,55.471209,LIAh  
DISKHOG,14000.  
RUN76,SC,NL7777.  
FETCHMT,TAPE3,D8,32277,1F.  
FETCHPS,BKYL1B,TEMP,GDS.  
FETCHPS,BKYLGOB,GDSCL,GDSCL.  
COPYI,TEMP,LIB.  
LINK,X,F=LGO,P=LIB,PP={LC=99999}.  
EXIT.  
RETURN,LGO,LIB.  
DUMP,O.  
FIN.  
DISPOSE,OUTPUT,MF.  
RETURN,TAPE99,GDSCL.
```

```
PROGRAM VSD(INPUT,OUTPUT,PUNCHB,TAPE6=PUNCHb,TAPE97,TAPE99,TAPE3,TVSD.2  
LAPE92,TAPE93,TAPE94) VSD.3  
VSD.4
```

```
CCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCVSD.5  
C. VSD.6  
C. VSD.7  
C. PROGRAM VSD READS THE DIGITAL DATA FROM TAPE 3 AND ESTIMATES VSD.8  
C. THE VELOCITY SPECTRAL DENSITY (VSD) OF THE TIME SERIES BY THE VSD.9  
C. METHOD OF MODIFIED PERICDUGRAM. VSD.10  
C. A 10 PERCENT COSINE BELL TAPER IS APPLIED TO EACH DATA BLOCK. VSD.11  
C. VSD IS DEFINED AS THE SQUARE ROOT OF POWER SPECTRAL DENSITY. VSD.12  
C. THE TAPE 3 IS GENERATED BY THE PROGRAM DUMP WHICH UNPACKES AND VSD.13  
C. TRANSLATES THE TAPE 5. THE TAPE 5 IS GENERATED BY THE DIGITIZER VSD.14  
C. IN HAVILAND HALL. EACH DATA BLOCK CONTAINS 2**NPOW DATA POINTS. VSD.15  
C. THE ARRAY DIMENSIONS ARE SET UP TO PROCESS THE DATA BLOCK WITH VSD.16  
C. MAXIMUM OF 512 DATA PLINTS PER DATA BLOCK. THEREFORE, EACH VSD.17  
C. RECORD OF 2500 PCINTS IN TAPE 3 IS DIVIDED INTO 4 DATA BLOCK VSD.18  
C. WITH 512 PCINTS PER BLOCK. VSD.19  
C. VSD.20  
C. VSD.21  
C. BWCUT = CUT OFF FREQUENCY (HZ) OF THE BUTTERWORTH FILTER(LOW-PASS). VSD.22  
C. DMPFT = DAMPING FACTOR OF THE SEISMOMETER. VSD.23  
C. ENW = ENERGY OF THE LAG WINDOW. VSD.24  
C. ERRBD = LOWER ERROR BAR AT 90 PERCENT CONFIDENCE. VSD.25  
C. ERRBU = UPPER ERROR BAR AT 90 PERCENT CONFIDENCE. VSD.26  
C. FS = SQRT(POWER SPECTAL DENSITY). VSD.27  
C. GAIN = THE GAIN OF THE AMPLIFIER, IN DECIBAL. VSD.28  
C. IIVSPL=1 WILL PLGT THE INDIVIDUAL FOURIER SPECTRUM. VSD.29  
C. IMVSPL=1 WILL PLOT THE AVERAGED FOURIER SPECTRUM. VSD.30  
C. INDEX3 CAN BE ANY INTEGER. VSD.31  
C. INDEX4 CAN BE ANY INTEGER. VSD.32  
C. IPUNCH=1 WILL PUNCH VSD ON CARD. VSD.33  
C. LABEL1 = STATION IDENTIFICATION. VSD.34  
C. MOREZ =0 INDICATES NO MORE INPUT DATA CARD. VSD.35  
C. MTAPER = THE PERCENTAGE OF THE FIRST AND THE LAST TIME DATA BEING TAPVSD.36  
C. BY A COSINE BELL. VSD.37  
C. NBW = NUMBER OF BUTTERWORTH FILTER USED AS ANTI-ALIASING FILTER. VSD.38  
C. NPOW NPTT = 2**NPOW VSD.39  
C. NPTF = THE NUMBER OF FREQUENCY COMPONENTS. VSD.40  
C. NPTT = THE NUMBER OF DATA POINT IN TIME DCMAIN. VSD.41  
C. NREC = NUMBER OF VSD TO BE AVERAGED. VSD.42
```

```

C. NSKIP(NNSKIP) = SEQUENTIAL NUMBER OF DATA BLOCK TO BE SKIPPED. VSD.43
C. NVAL(NSMP) = DIGITAL TIME SERIES OF ONE RECORD BLOCK. VSD.44
C. NSMP = NUMBER OF SAMPLING POINT IN EACH RECORD VSD.45
C. POWDST = POWER SPECTRAL DENSITY VSD.46
C. RCCUT = CUT OFF FREQUENCY(FZ) OF 2-POLE RC LOW-PASS FILTER. VSD.47
C. RD = DAMPING RESISTANCE OF THE SEISMOMETER, IN KILO-OHMS. VSD.48
C. RECLen = LENGTH OF EACH RECORD BLOCK OF 2500 POINTS (IN SECOND). VSD.49
C. RS = COIL RESISTANCE OF THE SEISMOMETER. IN KILO-OHMS. VSD.50
C. SAMIN = SAMPLING INTERVAL IN REAL TIME DOMAIN. VSD.51
C. VCOFS = FULL SCALE VOLTAGE OF VOLTAGE CONTROLLED OSCILLATOR. VSD.52
C. VS = VELOCITY SPECTRAL DENSITY, IN MILLIMICRON/SEC/SQRT(HZ) VSD.53
C. VSD.54
C. VSD.55
C. THE INPUT DATA CARDS ARE VSD.56
C. (1) IPUNCH,NPUN,RECLen (215,F5.0) VSD.57
C. (2) MORE2,IIVSPL,IMVSPL,INDEX3,INDEX4,LABEL1,NNSKIP,(NSKIP(I),I=1,VSD.58
C. 6),RS,RD,DMPFT,GAIN,RCCUT,VCOFS,NBW,BWCUT(511,A10,712,6F6.2, VSD.59
C. 11,F4.1) FOR EACH SEISMOMETER OUTPUT. VSD.60
C. VSD.61
C.....AUTHOR ALFRED LIANG-CHI LIAW
C ENGINEERING GEOSCIENCE
C UNIVERSITY OF CALIFORNIA, BERKELEY
C.....DATE SEPTEMBER 1977
C. VSD.62
CCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCC VSD.63
COMPLEX DATA,FUATA,CFDATA,PCWDST VSD.65
COMPLEX X VSD.66
COMMON /BLK1/FREQ(257) VSD.67
COMMON /BLK2/VS(257) VSD.68
COMMON /BLK3/ ERPLU(257),ERRBD(257) VSD.69
COMMON /BLK4/NREC,LABEL1 VSD.70
COMMON /BLK5/INDEX1,INDEX2 VSD.71
COMMON /BLK6/RLFS(257) VSD.72
COMMON /BLK7/STAT VSD.73
COMMON /BLK8/II VSD.74
DIMENSION NVAL(2500),DATA(512),FDATA(257),CFDATA(257),POWDST(257) VSD.75
DIMENSION SUMFS(257),FS(257),SPECS(12),RFFS(257) VSD.76
DIMENSION PHASE(257) VSD.77
DIMENSION NSKIP(6),CHI5(10),CH195(20) VSD.78
DIMENSION DATAID(512) VSD.79
DIMENSION A(257,12),X(12) VSD.80
DATA CHI95/0.0039,0.052,0.117,0.178,0.229,0.272,0.310,0.342,0.369,VSD.81
10.394,0.416,0.436,0.453,0.469,0.484,0.498,0.510,0.522,0.532,0.543/VSD.82
DATA CHI5/3.84,3.00,2.60,2.37,2.21,2.10,2.01,1.94,1.88,1.83/ VSD.83
DATA NSMP/2500,MTAPER/10,2/1.645/ VSD.84
READ 1,IPUNCH,NPUN,RECLen VSD.85
1 FORMAT (215,F5.0) VSD.86
SPECS(12)=99. VSD.87
NPTT=2**NPUN VSD.88
NPTF1=(NPTT/2)+1 VSD.89
NPTF=NPTF1-1 VSD.90
SAMIN=RECLen/(NSMP-1) VSD.91
PERIOD =(SAMIN)*(NPTT-1) VSD.92
PI=4.*ATAN(1.) VSD.93
PI2=PI*2.0 VSD.94
M1=NPTT*MTAPER/100 VSD.95
M2=NPTT-M1 VSD.96
NTAPER=0 VSD.97
NTRFUN=0 VSD.98
VSD.99
VSD.100

```

```
C.....GENERATING THE ARRAY OF FREQUENCY
DU 21 I=1,NPTF
21 FREQ(I)=1/PERIOD
97 IA=0
NS=12

98 READ 2,MGRE2,IIVSPL,IMVSPL,INDEX3,INDEX4,LABEL1,NNSKIP,(NSKIP(I),IVSD.101
I=1,6),RS,RD,DMPFT,GAIN,RCCUT,VCOFS,NBW,BWCUT VSD.102
2 FORMAT (5I1,A10,7I2,6F6.2,11,F4.1) VSD.103
PRINT 100,MGRE2,IIVSPL,IMVSPL,INDEX3,INDEX4,LABEL1,NNSKIP,(NSKIP(I) VSD.104
I),I=1,6),RS,RD,DMPFT,GAIN,RCCUT,VCOFS,NBW,BWCUT VSD.105
108 FURMAT (2X,5I1,A10,7I2,6F6.2,11,F4.1) VSD.106
VSD.107
IA=IA+1 VSD.108
IF (EOF(5LINPUT).NE.0) GO TO 999 VSD.109
DU 20 KI=1,NPTF VSD.110
SUMFS(KI)=0.0 VSD.111
20 CONTINUE VSD.112
NREC=0 VSD.113
VSD.114
99 READ (3) (NDAYS,NHOURS,MIN,NSEC) VSD.115
IF (EOF(3).NE.0) GO TO 599 VSD.116
READ (3) (NDAYST,NHOURST,MINST,NSECT) VSD.117
READ (3) (STATN,NB,J) VSD.118
ENCODE (10,5,STAT) STATN,NB VSD.119
5 FORMAT (R0,A2) VSD.120
READ (3) (NVAL(I),I=1,2500) VSD.121
READ (3) (MORE) VSD.122
VSD.123
PRINT 3,STATN,NB,J,MORE VSD.124
3 FORMAT (1H0,* STATN IN TAPE 3 = *,A10,* NB = *,I1,* J = *,I3,* MO VSD.125
IRE = *,I1) VSD.126
VSD.127
IF (IMVSPL.NE.1) GO TO 33 VSD.128
VSD.129
VSD.130
VSD.131
C PRINT 4,STAT,NPTI,IIVSPL,IMVSPL,INDEX3 VSD.132
4 FURMAT (1H1,//////////,50X,*THE FOLLOWING DATA CAME FROM *,A10 VSD.133
1,/,50X,*NUMBER OF POINT IN TIME DCMAIN = *,I4,/,50X,*INDEX1 = *,I VSD.134
1,/,50X,*INDEX2 = *,I1,/,50X,*INDEX3 = *,I1,//////////) VSD.135
DU 19 I=1,NNSKIP VSD.136
IF (NB.EG.NSKIP(I)) GO TO 331 VSD.137
19 CONTINUE VSD.138
C.....ADJUSTING THE DC SHIFT. VSD.139
VSD.140
SHIFT=0.0 VSD.141
DU 666 IDC=1,2500 VSD.142
SHIFT=SHIFT+FLOAT(NVAL(IDC)) VSD.143
666 CONTINUE VSD.144
SHIFT=SHIFT/2500.0 VSD.145
DU 667 IDC=1,2500 VSD.146
NVAL(IDC)=NVAL(IDC)-SHIFT VSD.147
667 CONTINUE VSD.148
29 LJ=1 VSD.149
VSD.150
VSD.151
VSD.152
VSD.153
VSD.154
VSD.155
C.....APPLYING COSINE BELL TAPER. VSD.156
90 ENW=0.0 VSD.157
DU 10 J=1,NPTT VSD.158
NVAL(J)=NVAL(J+LJ) VSD.159
IF (J.LT.M1) GO TO 7 VSD.160
IF (J.GT.M2) GO TO 9 VSD.161
TAPER=1.0 VSD.162
```

```
GO TO 8 VSD.163
7 ARG=PI*(M1-J+1)/M1 VSD.164
  TAPER=(1.0+COS(ARG))/2.0 VSD.165
  GO TO 8 VSD.166
9 ARG=PI*(J-M2)/M1 VSD.167
  TAPER=(1.0+COS(ARG))/2.0 VSD.168
8 CONTINUE VSD.169
  ENW=ENW+TAPER*TAPER VSD.170
  DATA(J)=CMPLX(TAPER*NVAL(J),0.0) VSD.171
C DATAWID(J)=TAPER VSD.172
10 CONTINUE VSD.173
C PRINT 109,ENW VSD.174
109 FORMAT (1H0,50X,*ENW = *,E12.3) VSD.175
VSD.176
C IF (NTAPER.NE.0) GO TO 25 VSD.177
C PMAX1=1.5 VSD.178
C PMINI=DATAWID(1) VSD.179
C CALL YPLOTLG(DATAWID,1,NPTT,1,PMAX1,PMINI,STAT) VSD.180
C NTAPER=1 VSD.181
25 CONTINUE VSD.182
C IF (LJ.LT.512) GO TO 15 VSD.183
C PRINT 14 VSD.184
14 FORMAT (//////) VSD.185
15 CONTINUE VSD.186
VSD.187
C PRINT 11, (NVAL(J),J=1,NPTT) VSD.188
11 FORMAT (2X,5I6,3X,5I6,3X,5I6,3X,5I6) VSD.189
GO TO 333 VSD.190
331 PRINT 332,STAT,NE VSD.191
332 FORMAT (1H0,10X,*THE PRGKAM SKIPS THE RECORD *,A10,I3) VSD.192
GO TO 33 VSD.193
VSD.194
333 CALL FFT(DATA,NPLW,NPTT,PI) VSD.195
DO 30 KJ=1,NPTF VSD.196
  FDATA(KJ)=DATA(KJ+1)*2.0*SAMIN VSD.197
  CFDATA(KJ)=CCNJG(FDATA(KJ)) VSD.198
  FREQ(KJ)=(KJ-1)/PERIOD VSD.199
  PWDST(KJ)=FDATA(KJ)*CFDATA(KJ) VSD.200
  PCWDST(KJ)=PCWDST(KJ)/PERIOD VSD.201
  PDWDST(KJ)=PCWDST(KJ)/ENW VSD.202
  FS(KJ)=SQRT(REAL(PLWDST(KJ))) VSD.203
  SJMM=FS(KJ) VSD.204
  SUMFS(KJ)=SUMFS(KJ)+SUMM VSD.205
30 CONTINUE VSD.206
VSD.207
C.....PLOT THE INDIVIDUAL FOURIER SPECTRUM BY THE YPLOT ROUTINE. VSD.208
VSD.209
CP PMAX1=0.00 VSD.210
CP PMINI=FS(1) VSD.211
CP DO 220 I=1,NPTF VSD.212
CP IF(FS(I).GT.PMAX1) PMAX1=FS(I) VSD.213
CP IF(FS(I).LT.PMINI) PMINI=FS(I) VSD.214
CP220 CONTINUE VSD.215
CP CALL YPLOTLG(FS,1,NPTF,1,PMAX1,PMINI,STAT) VSD.216
VSD.217
LJ=LJ+NPTT VSD.218
NREC=NREC+1 VSD.219
IF (IIVSPL.NE.1) GO TO 36 VSD.220
II=0 VSD.221
IF (NREC.EQ.1) GO TO 34 VSD.222
CALL NXTFRM(SPECS) VSD.223
VSD.224
```

```
34 CALL PLOTFS(NPTF) VSD.225
36 IF (LJ.LT.2049) GO TO 90 VSD.226
33 IF (MORE.NE.0) GO TO 99 VSD.227
C 33 IF (NREC.LT.NNREC) GO TO 99 VSD.228
IF (IMVSPL.NE.1) GO TO 45 VSD.229
DO 40 KK=1,NPTF VSD.230
FS(KK)=SUMFS(KK)/NREC VSD.231
40 CONTINUE VSD.232
C.....CORRECTING THE SYSTEM RESPONSE VSD.233
CALL GEOPHCN(VS,FS,DMPFT,PI2,NPTF,RS,RD) VSD.234
IF (NBW.NE.0) CALL ANTIALI(VS,PI2,NBW,BWCUT,NPTF) VSD.235
CALL AMPVCO(VS,NPTF,GAIN,VCCFS,RCCUT) VSD.236
C.....CALCULATING THE CONFIDENCE LIMITS VSD.237
NFRDM=2*NREC VSD.238
IF (NFRDM.GT.10) GO TO 51 VSD.239
CHI2=CHI5(NFRDM) VSD.240
GO TO 52 VSD.241
51 CALL CHISC(NFRDM,Z,CHI2) VSD.242
52 IF (NFRDM.GT.20) GO TO 53 VSD.243
CHI1=CHI95(NFRDM) VSD.244
GO TO 54 VSD.245
53 Z1=-Z VSD.246
CALL CHISC(NFRDM,Z1,CHI1) VSD.247
54 CONTINUE VSD.248
DO 55 KK=1,NPTF VSD.249
ERRBU(KK)=VS(KK)/SQRT(CHI1) VSD.250
ERRBD(KK)=VS(KK)/SQRT(CHI2) VSD.251
55 CONTINUE VSD.252
PRINT 35, NREC,LABEL1,CHI1,CHI2 VSD.253
35 FORMAT (I1,//////////,35X,*THE FOLLOWING DATA ARE THE ESTIMATED VSD.254
ITES OF VELOCITY SPECTRAL DENSITY*,/,35X, *OVER*,I4,* RECORDS OF *,VSD.255
IA10,/,35X,*CHISC OVER DF =*,2F10.3) VSD.256
PRINT 350,NPTT,SAMIN VSD.257
350 FORMAT (35X,*THERE ARE*,I5,* DATA PLINTS IN EACH RECORD, WITH SAMP VSD.258
LING INTERVAL *,F12.3,* SECCNDS.*) VSD.259
VSD.260
VSD.261
VSD.262
VSD.263
VSD.264
VSD.265
VSD.266
VSD.267
VSD.268
VSD.269
VSD.270
VSD.271
VSD.272
VSD.273
VSD.274
VSD.275
VSD.276
VSD.277
VSD.278
VSD.279
VSD.280
VSD.281
VSD.282
VSD.283
VSD.284
VSD.285
VSD.286
PMAI=0.00
PMINI=VS(1)
DO 230 I=1,NPTF
IF(VS(I).GT.PMAI) PMAI=VS(I)
IF(VS(I).LT.PMINI) PMINI=VS(I)
230 CONTINUE
CALL YPLOTLG(VS,I,NPTF,I,PMAI,PMINI,LABELI)
IF (IPUNCH.NE.1) GO TO 443
WRITE (6) (VS(I),I=1,NPTF)
WRITE (6) (CHI1,CHI2)
443 CALL SECONDD(TIME)
PRINT 444, TIME
444 FORMAT (15X,*TIME = *,F15.4)
IF (IMVSPL.NE.1) GO TO 44
CALL NXTRFM(SPECS)
II=1
50 CALL PLOTFS(NPTF)
44 CONTINUE
```

```
45 IF (MOPE2.NE.0) GO TO 98
999 CALL GDSSEND(SPECS)
STOP
END
```

VSD.287  
VSD.288  
VSD.289  
VSD.290

```
      SUBROUTINE AMPVCO(VS,NPTF,GAIN,VCCFS,RCCUT)
C.....SUBROUTINE AMPVCO IS USED TO CORRECT THE GAIN AND THE LOW-PASS
C FILTER OF THE AMPLIFIER AND THE VCO FULL SCALE.
C THE LOW-PASS FILTER AT AS110 IS A 2-POLE RC FILTER.
COMMON /BLK1/FREQ(257)
DIMENSION VS(257)
GAIN=GAIN/20.0
GAIN=10.0**GAIN
A=GAIN*2048.0/VCOFS
DO 1 I=1,NPTF
B=FREQ(I)/RCCUT
B2=B*B+1
VS(I)=VS(I)*B2/A
VS(I)=VS(I)*1.0E+09
1 CONTINUE
RETURN
END
```

AMPVC.2  
AMPVC.3  
AMPVC.4  
AMPVC.5  
AMPVC.6  
AMPVC.7  
AMPVC.8  
AMPVC.9  
AMPVC.10  
AMPVC.11  
AMPVC.12  
AMPVC.13  
AMPVC.14  
AMPVC.15  
AMPVC.16  
AMPVC.17  
AMPVC.18  
AMPVC.19  
AMPVC.20

```
      SUBROUTINE ANTIALI(VS,PI2,NBW,BWCUT,NPTF)
C.....SUBROUTINE ANTIALI IS USED TO CORRECT MODULUS OF THE TRANSFER
C FUNCTION OF ARBITRARY BUTTERWORTH LOW PASS FILTER USED AS ANTI-ALIASING
C FILTER.
C NBW = NUMBER OF STAGE OF FILTERS.
C NPOLE = NUMBER OF POLE OF THE FILTER.
COMMON /BLK1/FREQ(257)
DIMENSION VS(257)
NPOLE=4
DO 1 I=1,NPTF
N=2*NPOLE
A=FREQ(I)/BWCUT
FOR HIGH PASS FILTER, A=1.0/A
A=A**N
B=SQRT(1+A)
B=B**NBW
VS(I)=VS(I)*B
1 CONTINUE
RETURN
END
```

ANTAL.2  
ANTAL.3  
ANTAL.4  
ANTAL.5  
ANTAL.6  
ANTAL.7  
ANTAL.8  
ANTAL.9  
ANTAL.10  
ANTAL.11  
ANTAL.12  
ANTAL.13  
ANTAL.14  
ANTAL.15  
ANTAL.16  
ANTAL.17  
ANTAL.18  
ANTAL.19  
ANTAL.20  
ANTAL.21

```
      SUBROUTINE CHISC(NFRDM,2,CHI)
C.....THIS SUBROUTINE CALCULATES PERCENTILES OF THE CHI-SQUARE/DEGREE OF
C.....FREEDOM DISTRIBUTIONS FOR LARGE VALUES OF DEGREE OF FREEDOM.
DF=FLUAT(NFRDM)
A=2.0/(9.0*DF)
B=SQRT(A)
C=1.0-A+2*B
CHI=C*C*C
RETURN
```

CHIS2.2  
CHIS2.3  
CHIS2.4  
CHIS2.5  
CHIS2.6  
CHIS2.7  
CHIS2.8  
CHIS2.9  
CHIS2.10  
CHIS2.11

END

CHIS2.12

SUBROUTINE FFT(A,M,N,PI)

FFT\*R.2

C.....THIS FFT SUBROUTINE PERFORMS THE FORWARD FAST FOURIER TRANSFORM  
 C ONLY. AFTER COOLEY, LEWIS, WELCH, OBTAINED FROM L.R. RABINER AND B.  
 C GOLD--- THEORY AND APPLICATION OF DIGITAL SIGNAL PROCESSING.  
 C.....N=2\*\*M  
 C.....A(N) = THE ARRAY OF DATA TO BE TRANSFORMED.  
 C.....FOURIER TRANSFORM IS THE FORM OF (REAL)-I(IMAG).

FFT\*R.3  
 FFT\*R.4  
 FFT\*R.5  
 FFT\*R.6  
 FFT\*R.7  
 FFT\*R.8  
 FFT\*R.9  
 FFT\*R.10  
 FFT\*R.11  
 FFT\*R.12  
 FFT\*R.13  
 FFT\*R.14  
 FFT\*R.15  
 FFT\*R.16  
 FFT\*R.17  
 FFT\*R.18  
 FFT\*R.19  
 FFT\*R.20  
 FFT\*R.21  
 FFT\*R.22  
 FFT\*R.23  
 FFT\*R.24  
 FFT\*R.25  
 FFT\*R.26  
 FFT\*R.27  
 FFT\*R.28  
 FFT\*R.29  
 FFT\*R.30  
 FFT\*R.31  
 FFT\*R.32  
 FFT\*R.33  
 FFT\*R.34  
 FFT\*R.35  
 FFT\*R.36  
 FFT\*R.37  
 FFT\*R.38

COMPLEX A(N),U,W,T  
 NV2=N/2  
 NM1=N-1  
 J=1  
 DO 7 I=1,NM1  
 IF (I.GE.J) GO TO 5  
 T=A(J)  
 A(J)=A(I)  
 A(I)=T  
 5 K=NV2  
 6 IF(K.GE.J) GO TO 7  
 J=J-K  
 K=K/2  
 GO TO 6  
 7 J=J+K  
 DO 20 L=1,M  
 LE=2\*\*L  
 LE1=LE/2  
 U=(1.0,0.0)  
 W=CMPLX(COS(PI/LE1),SIN(PI/LE1))  
 DO 20 J=1,LE1  
 DO 10 I=J,N,LE  
 IP=I+LE1  
 T=A(IP)\*U  
 A(IP)=A(I)-T  
 10 A(I)=A(I)+T  
 20 U=U\*W  
 RETURN  
 END

SUBROUTINE GEOPHCN(VS,FS,DMPFT,PI2,NPTF,RS,RD)

GEOPH.2  
 GEOPH.3  
 GEOPH.4  
 GEOPH.5  
 GEOPH.6  
 GEOPH.7  
 GEOPH.8  
 GEOPH.9  
 GEOPH.10  
 GEOPH.11  
 GEOPH.12  
 GEOPH.13  
 GEOPH.14  
 GEOPH.15  
 GEOPH.16  
 GEOPH.17  
 GEOPH.18  
 GEOPH.19  
 GEOPH.20

C.....MODULUS CORRECTION ONLY.  
 C.....SUBROUTINE GEOPHCN IS USED TO CORRECT THE VELOCITY SENSITIVITY,  
 C DAMPING RESISTANCE, AND GENERATOR CONSTANCE.  
 C.....WG = NATURAL FREQUENCY OF THE SEISMOMETER (IN HZ).  
 C GS = GENERATOR CONSTANCE OF THE SEISMOMETER (IN VOLT/METER/SEC).  
 C RS = COIL RESISTANCE.  
 C RD = DAMPING RESISTANCE.  
 C DMPFT = DAMPING FACTOR.

COMMON /BLK1/FREQ(257)  
 DIMENSION VS(257),FS(257)  
 DATA WG/4.5/,GS/77.17/  
 WS=WG\*PI2  
 WS2=WS\*WS  
 RD2=2.0\*DMPFT  
 RD=RD\*1000.0  
 RS=RS\*1000.0

DO 1 I=1,NPTF	GEOPH.21
WE=PI2*FREQ(I)	GEOPH.22
WE2=WE*WE	GEOPH.23
A=WE2-WS2	GEOPH.24
A2=A*A	GEOPH.25
B=RO2*WS*WE	GEOPH.26
B2=B*B	GEOPH.27
C=SQRT(A2+B2)	GEOPH.28
AMP=WE*WE/C	GEOPH.29
VS(I)=FS(I)/AMP	GEOPH.30
VS(I)=VS(I)/GS	GEOPH.31
IF (RD.GE.999000.0) GO TO 1	GEOPH.32
R=RS+RD	GEOPH.33
VS(I)=VS(I)*R/RD	GEOPH.34
1 CONTINUE	GEOPH.35
RETURN	GEOPH.36
END	GEOPH.37

SUBROUTINE PLOTFS(NPTF)	PLFS.2
C.....THIS SUBROUTINE PLOTS THE VELOCITY SPECTRAL DENSITY IN LOG-LOG	PLFS.3
C.....SCALE.	PLFS.4
	PLFS.5
	PLFS.6
DIMENSION SPECS(30),LINE(4)	PLFS.7
COMMON /BLK1/X(257)	PLFS.8
COMMON /BLK2/Y(257)	PLFS.9
COMMON /BLK3/ERRBU(257),ERRBC(257)	PLFS.10
COMMON /BLK4/NREC,LABEL1	PLFS.11
COMMON /BLK5/INDEX1,INDEX2	PLFS.12
COMMON /BLK7/STAT	PLFS.13
COMMON /BLK8/I1	PLFS.14
SPECS(1)=1.0	PLFS.15
SPECS(2)=1.5	PLFS.16
SPECS(3)=25.0	PLFS.17
SPECS(4)=0.3	PLFS.18
DO 8 I=1,15,NPTF	PLFS.19
IF (Y(I).GT.1.0E+03) GO TO 6	PLFS.20
IF (Y(I).LT.1.0E-01) GO TO 9	PLFS.21
8 CONTINUE	PLFS.22
SPECS(5)=1.0E+03	PLFS.23
SPECS(6)=1.0E-01	PLFS.24
GO TO 5	PLFS.25
6 SPECS(5)=1.0E+04	PLFS.26
SPECS(6)=1.0	PLFS.27
GO TO 5	PLFS.28
9 SPECS(5)=1.0E+01	PLFS.29
SPECS(6)=1.0E-03	PLFS.30
5 SPECS(7)=7.875	PLFS.31
SPECS(8)=5.2	PLFS.32
SPECS(11)=1.0	PLFS.33
SPECS(12)=99.	PLFS.34
CALL AXLGLG(SPECS)	PLFS.35
YMAX=SPECS(5)	PLFS.36
YMIN=SPECS(6)	PLFS.37
DO 10 I=1,NPTF	PLFS.38
IF (Y(I).GT.YMAX) Y(I)=YMAX	PLFS.39
IF (Y(I).LT.YMIN) Y(I)=YMIN	PLFS.40
10 CONTINUE	PLFS.41
C PRINT 11,(X(I),Y(I),I=1,409)	PLFS.42
C 11 FORMAT (10X,*FREQ = *,F10.7,10X,*AMPLITUDE SPECTRUM *,E13.3)	PLFS.43
SPECS(25)=-0.15	PLFS.44

CALL SAXLGT(SPECS)	PLFS.45
SPECS(27)=-0.15	PLFS.46
CALL SAXLGR(SPECS)	PLFS.47
SPECS(13)=FLCAT(NPTF)	PLFS.48
SPECS(14)=1.0	PLFS.49
SPECS(15)=1.0	PLFS.50
SPECS(30)=97.	PLFS.51
CALL SLLGLG(X,Y,SPECS)	PLFS.52
IF (11.EQ.0) GO TO 15	PLFS.53
	PLFS.54
SPECS(13)=FLCAT(NPTF)	PLFS.55
DU 12 I=1,164	PLFS.56
Y(I)=ERRBU(I)	PLFS.57
IF (Y(I).GT.YMAX) Y(I)=YMAX	PLFS.58
IF (Y(I).LT.YMIN) Y(I)=YMIN	PLFS.59
12 CONTINUE	PLFS.60
SPECS(16)=17.0	PLFS.61
SPECS(17)=0.05	PLFS.62
SPECS(18)=0.05	PLFS.63
CALL PSLGLG(X,Y,SPECS)	PLFS.64
DU 13 I=1,NPTF	PLFS.65
Y(I)=ERRBD(I)	PLFS.66
IF (Y(I).GT.YMAX) Y(I)=YMAX	PLFS.67
IF (Y(I).LT.YMIN) Y(I)=YMIN	PLFS.68
13 CONTINUE	PLFS.69
CALL PSLGLG(X,Y,SPECS)	PLFS.70
15 SPECS(17)=-1	PLFS.71
SPECS(18)=-1	PLFS.72
SPECS(9)=10.0	PLFS.73
SPECS(19)=0.0	PLFS.74
SPECS(21)=1.0	PLFS.75
SPECS(24)=0.0	PLFS.76
SPECS(29)=2.0	PLFS.77
CALL NCLGB(SPECS)	PLFS.78
SPECS(20)=0.0	PLFS.79
SPECS(26)=0.0	PLFS.80
CALL NOLGL(SPECS)	PLFS.81
SPECS(17)=0.15	PLFS.82
SPECS(18)=0.15	PLFS.83
SPECS(24)=0.2	PLFS.84
CALL TITLEB(14HFREQUENCY (HZ),SPECS)	PLFS.85
SPECS(20)=90.0	PLFS.86
SPECS(26)=0.4	PLFS.87
CALL TITLEL(30HVSD (MILLIMICRON/SEC/SQRT(HZ)),SPECS)	PLFS.88
SPECS(17)=0.15	PLFS.89
SPECS(18)=0.15	PLFS.90
SPECS(19)=0.0	PLFS.91
SPECS(20)=0.0	PLFS.92
SPECS(21)=1.0	PLFS.93
SPECS(23)=6.6	PLFS.94
RULE=1.0	PLFS.95
IF (11.EQ.0) GO TO 20	PLFS.96
SPECS(22)=2.15	PLFS.97
SPECS(28)=0.0	PLFS.98
VALUE=NREC	PLFS.99
CALL DECVAL(RULE,VALUE,SPECS)	PLFS.100
LINE(1)=LABEL	PLFS.101
LINE (2)=10H RECORDS	PLFS.102
LINE(3)=0	PLFS.103
GO TO 21	PLFS.104
20 LINE(1)=STAT	PLFS.105
LINE(2)=0	PLFS.106

```

21 SPECS(22)=0.5
CALL TITLEG(RULE,LINE,SPECS)
RETURN
END

```

PLFS.107  
PLFS.108  
PLFS.109  
PLFS.110

```

SUBROUTINE YPLOTG(Y,IB,IE,IK,PMAX,PMIN,TITLE)
COMMON /BLK1/X(257)
DIMENSION Y(1),KMAP(100),ISCALE(11)
DATA KBLANK,KY/1H ,1H*/
DATA ISCALE/ 0,10,20,30,40,50,60,70,80,90,100/
DO 2 I=1,100
2 KMAP(I)=KBLANK
C....PRINT BEGINNING OF GRAPH
QMAX=ALOG10(PMAX)
QMIN=ALOG10(PMIN)
C=100.0/(QMAX-QMIN)
B=C*QMIN
SCALE=1.0/C
PRINT 5,TITLE,IB,IE,IK,PMAX,PMIN
5 FORMAT (1H1,10X,*YPLG.....*,A1C,/,10X,*IB=*,15,2X,*IE=*,15,2X,*I
IK=*,15,2X,*PMAX=*,E12.3,2X,*PMIN=*,E12.3,///)
PRINT 6, (ISCALE(I),I=1,11)
6 FORMAT(11X,5(12,8X),1X,12,5(7X,13))
DO 7 I=1,2
7 PRINT 8
8 FORMAT(13X,(11(1H. ,9X)))
PRINT 12
12 FORMAT ( 6X,1HX,7X,10CF.....,7X,1HY
1Y)
C....PRINT BODY OF GRAPH
DO 10 I=IB,IE,IK
YLG=ALOG10(Y(I))
IY=C*YLG-B
C CHECK THAT Y IS ON SCALE
IF(IY .LT. 1 .OR. IY .GT. 100) GO TO 13
KMAP(IY)=KY
PRINT 9, X(I),(KMAP(J),J=1,100),Y(I),I
9 FORMAT ( 2X,F9.3,1X,1H1,100A1,1H1,2X,E10.3,1X,14)
GO TO 14
13 PRINT 9, X(I),(KMAP(J),J=1,100),Y(I)
GO TO 10
14 KMAP(IY)=KBLANK
10 CONTINUE
RETURN
END

```

YPLLG.2  
YPLLG.3  
YPLLG.4  
YPLLG.5  
YPLLG.6  
YPLLG.7  
YPLLG.8  
YPLLG.9  
YPLLG.10  
YPLLG.11  
YPLLG.12  
YPLLG.13  
YPLLG.14  
YPLLG.15  
YPLLG.16  
YPLLG.17  
YPLLG.18  
YPLLG.19  
YPLLG.20  
YPLLG.21  
YPLLG.22  
YPLLG.23  
YPLLG.24  
YPLLG.25  
YPLLG.26  
YPLLG.27  
YPLLG.28  
YPLLG.29  
YPLLG.30  
YPLLG.31  
YPLLG.32  
YPLLG.33  
YPLLG.34  
YPLLG.35  
YPLLG.36  
YPLLG.37  
YPLLG.38  
YPLLG.39  
YPLLG.40  
YPLLG.41  
YPLLG.42  
YPLLG.43

```

SUBROUTINE YPLOT(Y,IB,IE,IK,PMAX,PMIN,TITLE)
COMMON /BLK1/X(257)
DIMENSION Y(1),KMAP(100),ISCALE(11)
DATA KBLANK,KY/1H ,1H*/
DATA ISCALE/ 0,10,20,30,40,50,60,70,80,90,100/
DO 2 I=1,100
2 KMAP(I)=KBLANK
C....PRINT BEGINNING OF GRAPH
C=100.0/(PMAX-PMIN)
B=C*PMIN
SCALE=1.0/C

```

YPLLI.2  
YPLLI.3  
YPLLI.4  
YPLLI.5  
YPLLI.6  
YPLLI.7  
YPLLI.8  
YPLLI.9  
YPLLI.10  
YPLLI.11  
YPLLI.12

```
PRINT 5,TITLE,IB,IE,IK,PMAX,PMIN YPLLI.13
5 FORMAT (1H1,10X,*YPLCT.....*,A10,/,10X,*IB=*,I5,2X,*IE=*,I5,2X,*IYPLLI.14
IK=*,I5,2X,*PMAX=*,E12.3,2X,*PMIN=*,E12.3,///) YPLLI.15
PRINT 6, (ISCALE(I),I=1,11) YPLLI.16
6 FORMAT(11X,5(I2,8X),1X,I2,5(7X,13)) YPLLI.17
DO 7 I=1,2 YPLLI.18
7 PRINT 8 YPLLI.19
8 FORMAT(13X,(11(1H. ,9X))) YPLLI.20
PRINT 12 YPLLI.21
12 FORMAT ( 6X,1FX,7X,10CF..... YPLLI.22
1.....,7X,1HYPLLI.23
1Y) YPLLI.24
C....PRINT BODY OF GRAPH YPLLI.25
DO 10 I=IB,IE,IK YPLLI.26
IY=C*Y(I)-B YPLLI.27
C CHECK THAT Y IS ON SCALE YPLLI.28
IF(IY .LT. 1 .OR. IY .GT. 100) GO TO 13 YPLLI.29
KMAP(IY)=KY YPLLI.30
PRINT 9, X(I),(KMAP(J),J=1,100),Y(I),I YPLLI.31
9 FORMAT ( 2X,F9.3,1X,1H1,100A1,1H1,2X,E10.3,1X,14) YPLLI.32
GO TO 14 YPLLI.33
13 PRINT 9, X(I),(KMAP(J),J=1,100),Y(I) YPLLI.34
GO TO 10 YPLLI.35
14 KMAP(IY)=KBLANK YPLLI.36
10 CONTINUE YPLLI.37
RETURN YPLLI.38
END YPLLI.39
```

0	9 02.5					
10102	HS	3.40999.00	0.30102.00	30.00	5.00310.00	
00102	A3.7A	3.40999.00	0.30114.00	30.00	5.00310.00	





```
LARGE C.XF(12,24,64) FK2.100
DIMENSION PMAXDB(20),FYP(30) FK2.101
DIMENSION SMNPF(12,12) FK2.102
DIMENSION V(12,24) FK2.103
DIMENSION NVAL(2500) FK2.104
DIMENSION X(12),MLF(12),EKX(12) FK2.105
DIMENSION NFY(50),FKPMAX(50) FK2.106
DIMENSION DATA(512),NP(41,41) FK2.107
DIMENSION SPECS(12),NSKIP(10) FK2.108
DIMENSION PHASE(226),w(512),T(226),NIB(12),WSPA(12),DTIME(72),NVSDFK2.109
I(12) FK2.110
COMMON /BLK1/FREQ(226) FK2.111
COMMON /BLK2/COH(226) FK2.112
COMMON /BLK5/STATN(12) FK2.113
COMMON /BLK7/XK(41),YK(41) FK2.114
COMMON /BLK8/P(41,41) FK2.115
COMMON /BLK9/NKX,NKY FK2.116
COMMON /BLK10/PLEVEL(10),NPLEV FK2.117
COMMON /BLK11/STA(2) FK2.118
COMMON /BLK12/VEL(5),NVEL,AZIMU(5) FK2.119
COMMON /BLK13/ERRBU(164),ERRBU(164) FK2.120
COMMON /BLK14/V(226),N(512),NPTF FK2.121
DATA MTAPER/10/ FK2.122
DATA Z95/1.645/ FK2.123
DATA NKX/41/,NKY/41/ FK2.124
DATA PMAXREF/10000.0/ FK2.125
CALL SECLND(TIME) FK2.126
FK2.127
FK2.128
C.....-----FK2.129
C.....READ INPUT DATA FK2.130
C.....-----FK2.131
PRINT 14,TIME FK2.132
READ 1,RECLEN,NPCH,NSMP,NBMAX,N2500 FK2.133
ICUN=0 FK2.134
STA(2)=0 FK2.135
CZERO=CMLPX(0.0,0.0) FK2.136
CI=CMLPX(0.0,1.0) FK2.137
SPECS(12)=99. FK2.138
NPTT=2**NPOW FK2.139
C. DUE TO THE LIMIT OF LCM, SET MAX. NPTF=226. FK2.140
NPTF=NPTT/2 FK2.141
IF (NPTF.GE.226) NPTF=226 FK2.142
SAMIN=RECLEN/FLOAT(NSMP-1) FK2.143
PERIOD=(SAMIN)*FLOAT(NPTT-1) FK2.144
PI=4.*ATAN(1.0) FK2.145
PI2=2.0*PI FK2.146
CALL HAMMIN(NPTT,PI2,w,ENW) FK2.147
M1=NPTT*MTAPER/100 FK2.148
M2=NPTT-M1 FK2.149
NTAPEK=0 FK2.150
READ 2,NPLEV,(PLEVEL(I),I=1,10) FK2.151
FK2.152
DO 110 I=1,NPTF FK2.153
110 FREQ(I)=FLOAT(I)/PERIOD FK2.154
FK2.155
READ 3,ICUNVEN,IFKPLLT,NVSDPL,(NVSD(I),I=1,12) FK2.156
IF (ICUNVEN.NE.1) GO TO 198 FK2.157
DO 140 I=1,12 FK2.158
140 WSPA(I)=1.0 FK2.159
198 READ 4,MORE2,ICGHER,IFKPSD,ISKAPP,NS,NAV1,ITRANS,NBW,BWCUT,SCALK,SFK2.160
1PACIN,STA(I) FK2.161
```

```
READ 5,NVNFY,(NFY(I),I=1,NNFY)
NF1=NFY(2)-NFY(1)
DO 220 I=2,NNFY
  I1=I+1
  NF2=NFY(I1)-NFY(I)
  IF (NF1.NE.NF2) GO TO 227
  IFKYPL=1
  IF (I1.GE.NNFY) GO TO 229
220 CONTINUE
GO TO 229
227 IFKYPL=0
229 CONTINUE
C. STATEMENTS FK2.176 AND FK2.220 ARE USED ONLY WHEN THAT
C. 3 RECORDS/ARRAY ELEMENT, N2500=2, AND THE FIRST RECORD BEING TO
C. ESTIMATE FKPSD.
  IDEL=-1
  NDTIME=N2500*NS
  READ 6,(DTIME(I),I=1,NDTIME)
  NAV=NAV1*2+1
  NAV2=NPTF-NAV1
  XKIN=(-1.0)/(2.0*SPACIN)
  XKIN=XKIN*SCALK
  DELK=(-2.0*XKIN)/FLOAT(NKX-1)
  READ 7,(X(I),I=1,NS)
  IF (EOF(5LINPUT).NE.0) GO TO 2672
  CALL SECOND(TIME)
  PRINT 14,TIME
  NSKCON=0
  DO 730 J=1,NS
    NSCON=NSCON+1.
  READ 8,STAT,(NNSKIP,(NSKIP(I),I=1,9),GAIN,RS,RD,DMPFT
  PRINT 15,STAT,NPTT,ICCHER,IFKPSD,STA(1),SPACIN,XKIN,DELK,X(J)
  PRINT 16,ICONVEN,IFKPLCT,NVSDPL,GAIN
  NB512=0
  M=0
  GO TO 377
373 PRINT 17

377 READ (3) (NDAYS,NHOURS,MIN,NSEC)
  IF (EOF(3).NE.0) GO TO 373
  READ (3) (NDAYST,NPCUST,MINT,NSELT)
  READ (3) (STATN(J),NG,JJ)
  IF (STATN(J).NE.STAT) GO TO 2675
  READ(3) (NVAL(I),I=1,NSMP)
  READ (3) (MORE)

  IF (ISKARR.NE.0) GO TO 602
  DO 460 I=1,NNSKIP
    IF (NB.EQ.NSKIP(I)) GO TO 602
460 CONTINUE

C.....-----
C.....ADJUSTING THE DC SHIFT.
C.....-----
  SHIFT=0.0
  DO 480 IDC=1,NSMP
    SHIFT=SHIFT+FLOAT(NVAL(IDC))
480 CONTINUE
  SHIFT=SHIFT/FLOAT(NSMP)
  DO 510 IDC=1,NSMP
    NVAL(IDC)=NVAL(IDC)-SHIFT
510 CONTINUE
```

FK2.162  
FK2.163  
FK2.164  
FK2.165  
FK2.166  
FK2.167  
FK2.168  
FK2.169  
FK2.170  
FK2.171  
FK2.172  
FK2.173  
FK2.174  
FK2.175  
FK2.1751  
FK2.176  
FK2.177  
FK2.176  
FK2.179  
FK2.180  
FK2.181  
FK2.182  
FK2.183  
FK2.184  
FK2.185  
FK2.186  
FK2.187  
FK2.188  
FK2.189  
FK2.190  
FK2.191  
FK2.192  
FK2.193  
FK2.194  
FK2.195  
FK2.196  
FK2.197  
FK2.198  
FK2.199  
FK2.200  
FK2.201  
FK2.202  
FK2.203  
FK2.204  
FK2.205  
FK2.206  
FK2.207  
FK2.208  
FK2.209  
FK2.210  
FK2.211  
FK2.212  
FK2.213  
FK2.214  
FK2.215  
FK2.216  
FK2.217  
FK2.218  
FK2.219  
FK2.220  
FK2.221  
FK2.222

LJ=1	FK2.223
C. SEE NOTE AT FK2.174 ABOUT THE STATEMENT FK2.226	FK2.224
IDEL=IDEL+2	FK2.225
	FK2.226
	FK2.227
NADV=CTIME(IDEL)/SAMIN	FK2.228
	FK2.229
C.....-----	FK2.230
C.....APPLYING HANNING WINDOW	FK2.231
C.....-----	FK2.232
521 DO 540 K=1,NPTT	FK2.233
RAN=(RANF(0)-0.5)*3.65	FK2.234
NVAL(K+LJ)=NVAL(K+LJ)+RAN	FK2.235
C    DATA(K)=CMPLX(NVAL(K+LJ+NADV)*GAIN*(K),0.0)	FK2.236
540 CONTINUE	FK2.237
	FK2.238
PRINT 18,STATN(J),NB,CTIME(IDEL),NADV	FK2.239
	FK2.240
IF (NB512.NE.0) GO TO 633	FK2.241
PRINT 19, (NVAL(JN),JA=1,NPTT)	FK2.242
IF (ICON.NE.0) GO TO 633	FK2.243
CALL SECOND(TIME)	FK2.244
PRINT 14,TIME	FK2.245
GO TO 633	FK2.246
602 PRINT 20,STATN(J),NB	FK2.247
IF (MORE.EQ.0.AND.NB512.EQ.0) GO TO 621	FK2.248
GO TO 725	FK2.249
621 PRINT 21,STAT	FK2.250
IF (NSKLN.GE.NS) GO TO 198	FK2.251
GO TO 730	FK2.252
	FK2.253
C.....-----	FK2.254
C.....FFT AND THE CORRECTION FOR SYSTEM RESPONSE	FK2.255
C.....-----	FK2.256
633 CALL FFT(DATA,NPOM,NPTT,P1)	FK2.257
DO 640 NF=1,NPTF	FK2.258
DATA(NF)=CONJG(DATA(NF))	FK2.259
640 CONTINUE	FK2.260
IF (KS.GT.1.0) CALL GICFON(DATA,P12,CMPFT,RO,KS,NPTF)	FK2.261
IF (NBW.NE.0) CALL ANTIALLI(DATA,P12,NBW,DELUT,NPTF)	FK2.262
IF (ICON.NE.0) GO TO 672	FK2.263
ICON=1	FK2.264
CALL SECOND(TIME)	FK2.265
PRINT 14,TIME	FK2.266
672 NB512=NB512+1	FK2.267
DO 700 NF=1,NPTF	FK2.268
700 XF(J,NB512,NF)=DATA(NF+1)*2.0*SAMIN	FK2.269
	FK2.271
	FK2.270
IF (NB512.GE.NBMAX) GO TO 726	FK2.272
	FK2.273
720 CONTINUE	FK2.274
LJ=LJ+NPTT	FK2.275
LJCK=LJ+NPTT	FK2.276
IF (LJCK.LT.NSMP) GO TO 521	FK2.277
725 IF (MORE.NE.0) GO TO 377	FK2.278
726 IF (MORE.NE.0) GO TO 377	FK2.279
NREC=NB512	FK2.280
730 CONTINUE	FK2.281
CALL SECOND(TIME)	FK2.282
PRINT 14,TIME	FK2.283
PRINT 22,NB512,NAV	FK2.284

```
IF (NB512.EQ.0) GO TO 158
TERM2=1.0/(2.0*FLCAT(NB512*NAV)-2.C)
TERM3=SQRT(FLCAT(NB512*NAV)*TERM2)*Z95
CALL SECOND(TIME)
PRINT 14,TIME
FK2.265
FK2.280
FK2.287
FK2.288
FK2.289
FK2.290
FK2.291
FK2.292
C.....-----
C.....CONSTRUCTING THE SPETRAL MATRIX SM(JS,LS,NF), IN THE UPPER TRIANGLEFK2.293
C.....-----FK2.294
DO 1070 NF=1,NPTF
DO 1060 JS=1,NS
DO 1050 LS=JS,NS
SF =CZER0
DO 1040 M=1,NREC
F=XF(JS,M,NF)
FF=CONJG(XF(LS,M,NF))
F=F*FF
SF=SF+F
FK2.295
FK2.296
FK2.297
FK2.298
FK2.299
FK2.300
FK2.301
FK2.302
FK2.303
1040 CONTINUE
SM(JS,LS,NF)=SF/FLCAT(NREC)
FK2.304
FK2.305
1050 CONTINUE
FK2.306
1060 CONTINUE
FK2.307
1070 CONTINUE
FK2.308
CALL SECOND(TIME)
FK2.309
PRINT 14,TIME
FK2.310
PRINT 879,(SM(1,1,NF),NF=1,NPTF)
FK2.311
PRINT 879,(SM(1,6,NF),NF=1,NPTF)
FK2.312
879 FORMAT (1H0,2X,*SM1*,//,24(2X,12E10.2,/)
FK2.313
FK2.314
C.....-----
C.....PLOTING VELOCITY SPECTRAL DENSITY OF INDIVIDUAL GEOPHONE OUTPUT. FK2.315
C.....-----FK2.317
IF (NVSDPL.LT.1) GO TO 1227
DO 1220 IVSD=1,NVSDPL
NV=NVS(IVSD)
FK2.318
FK2.319
READ 24, AMGAIN,VCCFS,RCCUT
FK2.320
PRINT 25, AMGAIN,VCCFS,RCCUT,ENW
FK2.321
FK2.322
DO 1200 NF=1,NPTF
FK2.323
TVS=REAL(SM(NV,NV,NF))/(PERIOD*ENW)
FK2.324
VS(NF)=SQRT(TVS)
FK2.325
1200 CONTINUE
FK2.326
IF (VCCFS.LT.1.0) GO TO 1212
FK2.327
CALL AMPVCO(VS,NPTF,AMGAIN,VCCFS,RCCUT)
FK2.328
1212 CONTINUE
FK2.329
PRINT 879,(VS(NF),NF=1,NPTF)
FK2.330
CALL PLOTVS
FK2.331
CALL NXTFRM(SPECS)
FK2.332
1220 CONTINUE
FK2.333
1227 CONTINUE
FK2.334
FK2.335
FK2.336
C.....-----
C.....RUNNING AVERAGE OVER (NAV) FREQUENCY COMPONENTS.
FK2.337
C.....-----FK2.338
IF (NAV.LE.1) GO TO 1377
FK2.339
DO 1360 JS=1,NS
FK2.340
DO 1350 LS=JS,NS
FK2.341
DO 1320 NF=1,NPTF
FK2.342
IF (NF.LE.NAV1) GO TO 1316
FK2.343
IF (NF.GE.NAV2) GO TO 1316
FK2.344
SF=(0.0,0.0)
FK2.345
FK2.346
```

```
DO 1300 NA=1,NAVI
II=NF-NA
I2=NF+NA
SF=SF+SM(JS,LS,I1)+SM(JS,LS,I2)
1300 CONTINUE
SF=SF+SM(JS,LS,NF)
T(NF)=SF/LOAT(NAV)
GO TO 1320
1316 T(NF)=SM(JS,LS,NF)
1320 CONTINUE
DO 1340 NF=1,NPTF
SM(JS,LS,NF)=T(NF)
1340 CONTINUE
1350 CONTINUE
1360 CONTINUE
JS=1
LS=1
PRINT 26,(SM(JS,LS,NF),NF=1,NPTF)
1377 CONTINUE
-----
C.....
C.....CALCULATING THE COHERENCE
C.....
IF (ICUHER.NE.1) GO TO 1569
DO 1540 JS=1,NS
DO 1530 LS=JS,NS
IF (JS.EQ.LS) GO TO 1530
DO 1460 NF=1,NPTF
SMTEMP=SM(JS,JS,NF)*SM(LS,LS,NF)
DEN=CSQRT(SMTEMP)
SM(JS,LS,NF)=SM(JS,LS,NF)/DEN
CUH(NF)=CABS(SM(JS,LS,NF))
1460 CONTINUE
CALL CONFLIM(TERM2,TERM3,NPTF)
CALL YPLOT(CUH,1,NPTF,1,1.0,0.0,STATN(JS),LS,FPEW)
PRINT 27,NB512
PRINT 28
PRINT 29,(ERRBU(IE),ERRFD(IE),IE=1,NPTF)
ISTA=STATN(JS)
IF (JS.GT.2) GO TO 1530
CALL PLOTCUH(ISTA,LS,NPTF)
CALL NXTRFM(SPECS)
1530 CONTINUE
1540 CONTINUE
DO 1570 JS=1,NS
DO 1560 NF=1,NPTF
SM(JS,JS,NF)=1.0
1560 CONTINUE
1570 CONTINUE
1569 IF (IFKPSD.EQ.0) GO TO 2671
PRINT 30
-----
C.....
C.....DO LOOP 2600 EVALUATES THE FKPSD AT DIFFERENT FREQUENCIES.
C.....
CALL SECOND(TIME)
PRINT 14,TIME
PDB=0.0
DO 2600 NF=1,NNFY
NVEL=0
```

```
NFP=NFY(NF) FK2.409
FRECY=FREQ(NFP) FK2.410
IF (FRECY.LT.8.0) GO TO 1627 FK2.411
XKIN=(-1.0)/(2.0*SPACIN) FK2.412
XKIN=XKIN*2.0*SCALN FK2.413
DELK=(-2.0*XKIN)/FLOAT(NKX-1) FK2.414
1627 CONTINUE FK2.415
FK2.416
C.....FK2.417
C.....NORMALIZED THE SPECTRAL MATRIX AT SFELECTED FREQUENCIES. FK2.418
C.....FK2.419
IF (ICORER.EQ.1) GO TO 1716 FK2.420
DO 1700 JS=1,NS FK2.421
DO 1690 LS=JS,NS FK2.422
IF (JS.EQ.LS) GO TO 1690 FK2.423
SMTEMP=SM(JS,JS,NFP)*SM(LS,LS,NFP) FK2.424
DEN=CSQRT(SMTEMP) FK2.425
SM(JS,LS,NFP)=SM(JS,LS,NFP)/DEN FK2.426
1690 CONTINUE FK2.427
1700 CONTINUE FK2.428
DO 1710 JS=1,NS FK2.429
SM(JS,JS,NFP)=1.0 FK2.430
1710 CONTINUE FK2.431
1716 CONTINUE FK2.432
PRINT 41 FK2.433
FK2.434
FK2.435
1722 PMAX1=0.0 FK2.436
DO 2350 KX=1,NKX FK2.437
XK(KX)=XKIN+DELK*FLOAT(KX-1) FK2.438
DO 2340 KY=1,NKY FK2.439
YK(KY)=XKIN+DELK*FLOAT(KY-1) FK2.440
FK2.441
IF (ICONVEN.EQ.1) GO TO 2213 FK2.442
FK2.443
C.....FK2.444
C.....CONSTRUCTING THE WAVE NUMBER MATRIX. FK2.445
C.....FK2.446
DO 2050 JS=1,NS FK2.447
DO 2040 LS=JS,NS FK2.448
SMNPF(JS,LS)=SM(JS,LS,NFP) FK2.449
IF (JS.EQ.LS) GO TO 2021 FK2.450
T1=X(JS)-X(LS) FK2.451
DX=REAL(T1) FK2.452
DY=AIMAG(T1) FK2.453
AEXP=XK(KX)*DX+YK(KY)*DY FK2.454
AEXP=AEXP*PI2 FK2.455
E=CMPLX(0.0,AEXP) FK2.456
E=CEXP(E) FK2.457
SMNPF(JS,LS)=SMNPF(JS,LS)*E FK2.458
SMNPF(LS,JS)=CONJG(SMNPF(JS,LS)) FK2.459
GO TO 2036 FK2.460
C.....FK2.461
C.....CONSTRUCTING V MATRIX (V) = (SM,I) FK2.462
C.....FK2.463
2021 V(JS,LS)=1.0 FK2.464
GO TO 2040 FK2.465
2036 V(JS,LS)=SMNPF(JS,LS) FK2.466
V(LS,JS)=SMNPF(LS,JS) FK2.467
2040 CONTINUE FK2.468
2050 CONTINUE FK2.469
NSS=NS+1 FK2.470
```

```
NSE=NS*2
DO 2100 JS=1,NS
DO 2070 LS=NSS,NSE
LS1=LS-NS
IF (JS.EQ.LS1) GO TO 2067
V(JS,LS)=CZERO
GO TO 2070
2067 V(JS,LS)=1.0
2070 CONTINUE
2100 CONTINUE
-----
C.....
C.....FIND THE INVERSE OF SPECTRAL MATRIX
C.....
CALL CMXDIV(NS,NS,DET,12,V)
DO 2130 JS=1,NS
DO 2120 LS=1,NS
NSS=LS+NS
SMNPF(JS,LS)=V(JS,NSS)
2120 CONTINUE
2130 CONTINUE
-----
C.....
C.....CONSTRUCTING THE MAXIMUM LIKELIHOOD FILTER.
C.....
SUMQJL=(0.0,0.0)
DO 2160 JS=1,NS
DO 2150 LS=1,NS
SUMQJL=SUMQJL+SMNPF(JS,LS)
2150 CONTINUE
2160 CONTINUE
DO 2210 JS=1,NS
SUMQL=(0.0,0.0)
DO 2170 LS=1,NS
SUMQL=SUMQL+SMNPF(JS,LS)
2170 CONTINUE
MLF(JS)=SUMQL/SUMQJL
MLF(JS)=CONJG(MLF(JS))
2210 CONTINUE
GO TO 2224
2213 DO 2220 LS=1,NS
MLF(LS)=CMPLX(WSPA(LS),C.0)
2220 CONTINUE
2224 DO 2250 JS=1,NS
TX=REAL(X(JS))
TY=AIMAG(X(JS))
BEXP=XK(KX)*TX+YK(KY)*TY
BEXP=BEXP*PI2
E=CMPLX(0.0,BEXP)
EKX(JS)=CEXP(E)
2250 CONTINUE
-----
C.....
C.....ESTIMATE FKPSD
C.....
SUM=0.
NRC=0
DO 2320 M=1,NREC
SUMFUN=CZERO
DO 2310 JS=1,NS
```

```
FK2.471
FK2.472
FK2.473
FK2.474
FK2.475
FK2.476
FK2.477
FK2.478
FK2.479
FK2.480
FK2.481
FK2.482
FK2.483
FK2.484
FK2.485
FK2.486
FK2.487
FK2.488
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FK2.490
FK2.491
FK2.492
FK2.493
FK2.494
FK2.495
FK2.496
FK2.497
FK2.498
FK2.499
FK2.500
FK2.501
FK2.502
FK2.503
FK2.504
FK2.505
FK2.506
FK2.507
FK2.508
FK2.509
FK2.510
FK2.511
FK2.512
FK2.513
FK2.514
FK2.515
FK2.516
FK2.517
FK2.518
FK2.519
FK2.520
FK2.521
FK2.522
FK2.523
FK2.524
FK2.525
FK2.526
FK2.527
FK2.528
FK2.529
FK2.530
FK2.531
FK2.532
```

```
CFUN=MLF(JS)*XF(JS,M,NFP)*EKX(JS) FK2.533
SUMFUN=SUMFUN+CFUN FK2.534
2310 CONTINUE FK2.535
FUN=CABS(SUMFUN) FK2.536
SUMM=FUN*FUN FK2.537
SUM=SUM+SUMM FK2.538
NBC=NBC+1 FK2.539
2320 CONTINUE FK2.540
P(KX,KY)=SUM/FLCAT(NBC) FK2.541
IF (P(KX,KY).LT.PMAX1) GO TO 2340 FK2.542
PMAX1=P(KX,KY) FK2.543
2340 CONTINUE FK2.544
2350 CONTINUE FK2.545
FKPMAX(NF)=PMAX1 FK2.546
IF (FKPMAX(NF).LT.PDB) GO TO 2353 FK2.547
PDB=FKPMAX(NF) FK2.548
2353 CONTINUE FK2.549
FK2.550
C..... FK2.551
C.....TRANSFORMING THE FKPSD INTO DECIBEL SCALE WITH RESPECT TO THE PEAK FK2.552
C.....FKPSD IN EACH FREQUENCY COMPONENT, THEN PRINT THE RESULTS. FK2.553
C..... FK2.554
-----
PRINT 31,FREQCY,STA(1),PMAX1,XKIN,DELK,NFY(NF) FK2.555
PRINT 32 FK2.556
PRINT 33 FK2.557
DO 2460 KY=1,NKY FK2.558
DO 2470 KX=1,NKX FK2.559
P(KX,KY)=10.0*ALOG10(P(KX,KY)/PMAX1) FK2.560
NP(KX,KY)=-10*P(KX,KY) FK2.561
IF (P(KX,KY).LT.0.0) GO TO 2470 FK2.562
XKPMAX=XK(KX)*XK(KX) FK2.563
YKPMAX=YK(KY)*YK(KY) FK2.564
NVEL=NVEL+1 FK2.565
DELKH=0.5*DELK FK2.566
TEMPK=SQRT(XKPMAX+YKPMAX) FK2.567
IF (TEMPK.LT.DELKH) GO TO 2454 FK2.568
VEL(NVEL)=FREQCY/TEMPK FK2.569
GO TO 2457 FK2.570
2454 VEL(NVEL)=999.999 FK2.571
THETA=0.000 FK2.572
AZIMU(NVEL)=0.000 FK2.573
GO TO 2470 FK2.574
2457 CONTINUE FK2.575
THETA=57.2958*ATAN2(YK(KY),XK(KX)) FK2.576
AZIMU(NVEL)=90.0-THETA FK2.577
IF (AZIMU(NVEL).GE.0.0) GO TO 2470 FK2.578
AZIMU(NVEL)=360.0+AZIMU(NVEL) FK2.579
2470 CONTINUE FK2.580
2480 CONTINUE FK2.581
DO 2530 JKY=1,NKY FK2.582
JY=NKY-JKY+1 FK2.583
IF (JY.EQ.21) GO TO 2510 FK2.584
PRINT 34,(NP(KX,JY),KX=1,NKX) FK2.585
GO TO 2530 FK2.586
2510 PRINT 35,(NP(KX,JY),KX=1,NKX) FK2.587
2530 CONTINUE FK2.588
PRINT 33 FK2.589
PRINT 32 FK2.590
PRINT 36,(VEL(IV),IV=1,NVEL) FK2.591
PRINT 37,(AZIMU(IV),IV=1,NVEL) FK2.592
IF (IFKPL0T.EQ.1) CALL FKPL0T(FREQCY,ICONVEN,NAV,PMAX1) FK2.593
```

```
CALL SYMCNT(P,NKX,NKY) FK2.595
CALL SECOND(TIME) FK2.596
PRINT 14,TIME FK2.597
IF (NF.EQ.NNFY.AND.MORE2.EQ.0) GO TO 2600 FK2.598
CALL NXTFRM(SPECS) FK2.599
2600 CONTINUE FK2.600
PRINT 38,STA(1) FK2.601
DO 2650 I=1,NNFY FK2.602
NFP=NFY(I) FK2.603
FYP(I)=FREQ(NFP) FK2.604
PMAxDB(I)=10.0*ALOG10(FKPPMAX(I)/PMAxREF) FK2.605
PRINT 39,I,NFP,FYP(I),FKPPMAX(I),PMAxDB(I) FK2.606
2650 CONTINUE FK2.607
IF (IFKYPL.EQ.1) CALL YPLOT(PMAxDB,1,NNFY,1,30.,-30.,STA(1),1,FYP) FK2.608
2671 IF (MORE2.NE.0) GO TO 198 FK2.609
2672 CALL GOSEND(SPECS) FK2.610
GO TO 2710 FK2.611
2675 PRINT 40,J,STATN(J),STAT FK2.612
2710 CONTINUE FK2.613
CALL SECOND(TIME) FK2.614
PRINT 14,TIME FK2.615
FK2.616
FK2.617
1 FORMAT (F5.0,4I5) FK2.618
2 FORMAT (I5,10F5.0) FK2.619
3 FORMAT (I5I5) FK2.620
4 FORMAT (8I5,3F5.0,A10) FK2.621
5 FORMAT (I6I5) FK2.622
6 FORMAT (I6F5.3) FK2.623
7 FORMAT (I0F8.4) FK2.624
8 FORMAT (A10,I0I2,4F5.0) FK2.625
14 FORMAT (15X,*TIME = *,F15.4) FK2.626
15 FORMAT (1H0,///,17H ,33X,*THE DATA CAME FROM *,A10FK2.627
1,/,50X,*NUMBER OF POINT IN TIME DOMAIN = *,I4,/,50X,*ICOHET = *,I1FK2.628
1,/,50X,*IFKPSD = *,I1,/,50X,*STA = *,A10,/,50X,*SEISMOMETER SPACINFK2.629
1G IN KM = *,F6.3,/,50X,*INITIAL KX AND KY = *,F7.2,/,50X,*DELTA K FK2.630
1 = *,F7.3,/,50X,*COORDINATES OF THE SFISMOMETER IN KM = *,2F7.3,/) FK2.631
16 FORMAT (1H0,49X,*ICONVEN = *,I2,/,50X,*IFKPLOT = *,I2,/,50X,*NVSDPFK2.632
1L = *,I2,/,50X,*GAIN = *,E12.3,/) FK2.633
17 FORMAT (1H0,10X,*END OF FILE MARK ENCUNTED*,/) FK2.634
18 FORMAT (1H0,10X,*STATICA *,A10,I3,* TIME CORRECTION IS *,F8.3,* SEFK2.635
1CUNDS, WHICH CORR. TO *,I5,* POINTS ADV. IN DIGITAL DATA.*) FK2.636
19 FORMAT (2X,5I6,3X,5I6,3X,5I6,3X,5I6) FK2.637
20 FORMAT (1H0,10X,*THE PRCGKAM SKIPS THE RECORD *,A10,I3,* AS REQUESTFK2.638
1TED.*) FK2.639
21 FORMAT (1H0,/,10X,*THE DIGITAL DATA FROM THE STATION *,A10,* GIVE FK2.640
1NB512=0 WHEN MURE=0.*) FK2.641
22 FORMAT (1H0,5X,*NC. OF 512PTS DATA BLOCK = *,I5,*RUNNING AVE. OVEFK2.642
1R *,I4,* POINTS.*,/) FK2.643
23 FORMAT (1H0,2X,*SM1*,//,24(2X,12E10.2,/) FK2.644
24 FORMAT (3F5.0) FK2.645
25 FORMAT (1H0,* AMPLIFIER GAIN = *,F6.2,* VCG FULL SCALE = *,F5.1,*FK2.646
1 RC CUT = *,F5.1,* W ENERGY = *,E10.3) FK2.647
26 FORMAT (1H0,2X,*S*2*,//,24(2X,12E10.2,/) FK2.648
27 FORMAT (//,2X,*NB512 = *,I5) FK2.649
28 FORMAT (3X,*95 PERCENT CONFIDENT LIMIT*,/) FK2.650
29 FORMAT (10(3X,2F5.2)) FK2.651
30 FORMAT (1H6) FK2.652
31 FORMAT (1H1,1X,*FREQUENCY WAVENUMBER POWER SPECTRAL DENSITY AT FREFK2.653
1QUENCY = *,F7.2,* AT*,A10,/,2X,*MAXIMAL PSD = *,E12.3,/,2X,*NGFK2.654
1TE THAT THE VALUES PRINTED ARE (-10)(PSD/PMAx1) IN Db.*,/,2X,*X295FK2.655
1 = *,F9.3,2X,*DELTA K = *,F9.3,//,2X,*NFY = *,I3,//////////) FK2.656
```

```
32 FORMAT (64X,1H0) FK2.657
33 FORMAT (2X,123H . . . . .) .FK2.658
1 . . . . . .FK2.659
1 . . . . .) FK2.660
34 FORMAT (1H0,1X,41(I3)) FK2.661
35 FORMAT (1H0,*0*,41(I3),* 0*) FK2.662
36 FORMAT (1H0,2X,*VELOCITY (KM/SEC) = *,5F10.3) FK2.663
37 FORMAT (1H0,2X,*AZIMUTH (DEGREE) = *,5F10.3) FK2.664
38 FORMAT (1H1,4X,A10,/) FK2.665
39 FORMAT (2I5,* FREQUENCY = *,F7.2,* MAXIMUM FKPSD = *,E12.3,F10.3FK2.666
1,* DB*) FK2.667
40 FORMAT (1H0,///,* THE PROGRAM STOPS DUE TO STATN(*,I2,*) = *,A10FK2.668
1,* WHICH CONFLICTS WITH STAT = *,A10) FK2.669
41 FORMAT (1HQ) FK2.670
STOP FK2.671
END FK2.672
```

```
SUBROUTINE AMPVCO(VS,NPTF,GAIN,VCCFS,RCCUT) AMPVC.2
AMPVC.3
C.....SUBROUTINE AMPVCO CORRECTS THE GAIN AND THE LOW-PASS FILTER OF THE AMPVC.4
C AMPLIFIER AND THE VCO FULL SCALE. AMPVC.5
C THE LOW-PASS FILTER AT AS110 IS 2-POLE RC FILTER. AMPVC.6
AMPVC.7
COMMON /BLK1/FREQ(226) AMPVC.8
DIMENSION VS(226) AMPVC.9
GAIN=GAIN/20.0 AMPVC.10
GAIN=10.0**GAIN AMPVC.11
A=GAIN*2048.0/VCCFS AMPVC.12
DO 1 I=1,NPTF AMPVC.13
B=FREQ(I)/RCCUT AMPVC.14
B2=B*B+1 AMPVC.15
VS(I)=VS(I)*B2/A AMPVC.16
VS(I)=VS(I)*1.0F+09 AMPVC.17
1 CONTINUE AMPVC.18
RETURN AMPVC.19
END AMPVC.20
```

```
SUBROUTINE ANTAL1(VS,PI2,NBW,BWCUT,NPTF) ANTAL.2
ANTAL.3
C.....SUBROUTINE ANTAL1 IS USED TO CORRECT MODULUS OF THE TRANSFER ANTAL.4
C FUNCTION OF ARBITRARY LOW PASS BUTTERWORTH FILTER SUCH AS THE ANTAL.5
C ANTI-ALIASING FILTER USED IN THE HAVILAND HALL. ANTAL.6
C NBW = NUMBER OF FILTER STAGE ANTAL.7
C NPOLE = NUMBER OF POLE OF THE FILTER. ANTAL.8
ANTAL.9
COMPLEX VS ANTAL.10
COMMON /BLK1/FREQ(226) ANTAL.11
DIMENSION VS(226) ANTAL.12
NPOLE=4 ANTAL.13
DO 1 I=1,NPTF ANTAL.14
N=2*NPOLE ANTAL.15
A=FREQ(I)/BWCUT ANTAL.16
A=A**N ANTAL.17
B=SQRT(1+A) ANTAL.18
B=B**NBW ANTAL.19
VS(I)=VS(I)*B ANTAL.20
1 CONTINUE ANTAL.21
RETURN ANTAL.22
END ANTAL.23
```

```

SUBROUTINE CMXDIV(IN,IM,DET,IDIM,V)
C REPLACES B BY (A INVERSE)*B. DESTROYS A.
C A IS N BY N, B IS N BY M. A AND B ARE STORED IN V(N,N+M).

COMPLEX SAVE,PIVOT,DETERM,DET,V,CAUS
DIMENSION V(IDIM,1)
INTEGER P
N=IN
M=IM

C
NM1=N-1
NP1=N+1
NPM=N+M
DETERM=(1.0E0,0.0E0)
P=1
202 CONTINUE
IMIN=P+1

C FIND PIVOT
RPIVOT=0.0E0
DO 204 I=P,N
  RSAVE=ABS(V(I,P))
  IF (RSAVE .LE. RPIVOT) GO TO 209
  RPIVOT=RSAVE
  IBIG=I
209 CONTINUE
IF (RPIVOT .GT. 0.0E0) GO TO 210
DETERM=(0.0E0,0.0E0)
GO TO 249

C FOR EXCHANGE AND POW. REDUCTION
210 PIVOT=V(IBIG,P)
DETERM=DETERM*PIVOT
DO 219 J=P,NPM
  SAVE=V(IBIG,J)/PIVOT
  V(IBIG,J)=V(P,J)
  V(P,J)=SAVE
219 CONTINUE
IF (IBIG .NE. P) DETERM=-DETERM
IF (P .GE. N) GO TO 250
V(P,P)=(1.0E0,0.0E0)

C BLOCK REDUCTION
DO 239 I=IMIN,N
  DO 238 J=IMIN,NPM
    V(I,J)=V(I,J)-V(I,P)*V(P,J)
238 CONTINUE
V(I,P)=(0.0E0,0.0E0)
239 CONTINUE
P=P+1
249 GO TO 202

C BACK SUBSTITUTION
250 CONTINUE
IF (M .LE. 0) GO TO 259
DO 259 K=NP1,NPM
  DO 253 P=1,NM1
    I=N-P
    DO 252 J=I,NM1
      V(I,K)=V(I,K)-V(J+1,K)*V(I,J+1)
252 CONTINUE
253 CONTINUE
259 CONTINUE
C
```

```

CMXDV.2
CMXDV.3
CMXDV.4
CMXDV.5
CMXDV.6
CMXDV.7
CMXDV.8
CMXDV.9
CMXDV.10
CMXDV.11
CMXDV.12
CMXDV.13
CMXDV.14
CMXDV.15
CMXDV.16
CMXDV.17
CMXDV.18
CMXDV.19
CMXDV.20
CMXDV.21
CMXDV.22
CMXDV.23
CMXDV.24
CMXDV.25
CMXDV.26
CMXDV.27
CMXDV.28
CMXDV.29
CMXDV.30
CMXDV.31
CMXDV.32
CMXDV.33
CMXDV.34
CMXDV.35
CMXDV.36
CMXDV.37
CMXDV.38
CMXDV.39
CMXDV.40
CMXDV.41
CMXDV.42
CMXDV.43
CMXDV.44
CMXDV.45
CMXDV.46
CMXDV.47
CMXDV.48
CMXDV.49
CMXDV.50
CMXDV.51
CMXDV.52
CMXDV.53
CMXDV.54
CMXDV.55
CMXDV.56
CMXDV.57
CMXDV.58
CMXDV.59
CMXDV.60
CMXDV.61
CMXDV.62
```

```

290 CONTINUE
DET=DETERM
RETURN
END

```

CMXDV.63  
CMXDV.64  
CMXDV.65  
CMXDV.66  
CMXDV.67

```

SUBROUTINE CONFLIM(TERM2,TERM3,NPTF)
C.....THIS ROUTINE CALCULATES THE CONFIDENT LIMITS OF COHERENCE ESTIMATE
COMMON /BLK2/COF(226)
COMMON /BLK13/ERRBU(164),ERRBC(164)
DIMENSION W(164)
DO 2 NF=1,NPTF
A=(1.0+COF(NF))/(1.0-COF(NF))
W(NF)=0.5*A*LOG(A)
B=W(NF)-TERM2
CU=B+TERM3
CD=B-TERM3
ERRBU(NF)=TANH(CU)
ERRBC(NF)=TANH(CD)
2 CONTINUE
RETURN
END

```

CONFL.2  
CONFL.3  
CONFL.4  
CONFL.5  
CONFL.6  
CONFL.7  
CONFL.8  
CONFL.9  
CONFL.10  
CONFL.11  
CONFL.12  
CONFL.13  
CONFL.14  
CONFL.15  
CONFL.16  
CONFL.17  
CONFL.18

```

SUBROUTINE FFT(A,M,N,PI)
C.....THIS FFT SUBROUTINE PERFORMS THE FORWARD FAST FOURIER TRANSFORM
C ONLY. AFTER COOLEY, LEWIS, WELCH, OBTAINED FROM L.R. RABINER AND B.
C GOLD--- THEOREY AND APPLICATION OF DIGITAL SIGNAL PROCESSING.
C.....N=2**M
C.....A(N) = THE ARRAY OF DATA TO BE TRANSFORMED.
COMPLEX A(N),U,W,T
NV2=N/2
NM1=N-1
J=1
DO 7 I=1,NM1
IF (1.0.E.J) GO TO 5
T=A(J)
A(J)=A(I)
A(I)=T
5 K=NV2
6 IF(K.GE.J) GO TO 7
J=J-K
K=K/2
GO TO 6
7 J=J+K
DO 20 L=1,M
LE=2**L
LE1=LE/2
U=(1.0,0.)
W=CMPLX(COS(PI/LE1),SIN(PI/LE1))
DO 10 J=1,LE1
DO 10 I=J,N,LE
IP=I+LE1
T=A(IP)*U
A(IP)=A(I)-T
10 A(I)=A(I)+T
20 U=U*W
RETURN

```

FFT.2  
FFT.3  
FFT.4  
FFT.5  
FFT.6  
FFT.7  
FFT.8  
FFT.9  
FFT.10  
FFT.11  
FFT.12  
FFT.13  
FFT.14  
FFT.15  
FFT.16  
FFT.17  
FFT.18  
FFT.19  
FFT.20  
FFT.21  
FFT.22  
FFT.23  
FFT.24  
FFT.25  
FFT.26  
FFT.27  
FFT.28  
FFT.29  
FFT.30  
FFT.31  
FFT.32  
FFT.33  
FFT.34  
FFT.35  
FFT.36  
FFT.37

END

FFT.38

SUBROUTINE FKPLCT(FRECCY,ICONV,N,NAV,PMAX)  
 ... THIS SUBROUTINE CONTOURS THE FREQUENCY WAVEP  
 AT A SELECTED FREQUENCY ON A K-KY PLANT.

```

COMMON /BLK7/X(41),Y(41)
COMMON /BLK8/Z(41,41)
COMMON /BLK9/NKA,NKY
COMMON /BLK10/ZLEVEL(10),NL
COMMON /BLK11/LINE(2)
COMMON /BLK12/VEL(5),NVEL,AZIML(5)
DIMENSION BUFGYZ(100),SPECS(35)
SPECS(1)=1.0
SPECS(2)=1.5
SPECS(3)=X(NKX)
SPECS(4)=X(1)
SPECS(5)=Y(NKY)
SPECS(6)=Y(1)
SPECS(7)=9.0
SPECS(8)=5.0
SPECS(11)=1.0
SPECS(12)=99.
SPECS(14)=2.0
SPECS(10)=2.0
CALL GDLI1(SPECS)
SPECS(30)=97.
SPECS(31)=NKA
SPECS(32)=NKY
SPECS(33)=FLOAT(NL)
LENGTH=67
CALL CONLI(X,Y,Z,ZLEVEL,BUFGYZ,LENGTH,SPECS)
SPECS(17)=0.15
SPECS(18)=0.15
SPECS(19)=0.0
SPECS(20)=0.0
SPECS(21)=1.0
SPECS(22)=6.1
SPECS(23)=0.6
SPECS(29)=3.0
RULE=1.0
CALL SIGVAL (RULE,FRECCY,0,SPECS)
SPECS(22)=6.9
CALL TITLEG(RULE,2HZ,SPECS)
SPECS(22)=7.4
SPECS(28)=0.0
VALUE=NAV
CALL DECVAL(RULE,VALUE,SPECS)
SPECS(22)=4.5
IF (ICONV.EQ.0) CALL TITLEG(RULE,4H.R.,SPECS)
IF (ICONV.EQ.1) CALL TITLEG(RULE,5HCONV.,SPECS)
SPECS(22)=2.5
CALL TITLEG(RULE,LINE,SPECS)
SPECS(17)=0.1
SPECS(18)=0.1
SPECS(19)=0.0
SPECS(24)=0.0
SPECS(28)=1.0
CALL NDDL18(SPECS)
SPECS(26)=0.0

```

FKPL.2  
 FKPL.3  
 DENSITYFKPL.4  
 FKPL.5  
 FKPL.6  
 FKPL.7  
 FKPL.8  
 FKPL.9  
 FKPL.10  
 FKPL.11  
 FKPL.12  
 FKPL.13  
 FKPL.14  
 FKPL.15  
 FKPL.16  
 FKPL.17  
 FKPL.18  
 FKPL.19  
 FKPL.20  
 FKPL.21  
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 FKPL.52  
 FKPL.53  
 FKPL.54  
 FKPL.55  
 FKPL.56  
 FKPL.57  
 FKPL.58  
 FKPL.59  
 FKPL.60

```

CALL MODLIL(SPECS)
SPECS(17)=0.15
SPECS(18)=0.15
SPECS(24)=0.2
CALL TITLE6(14HKX (CYCLES/KM),SPECS)
SPECS(26)=0.4
CALL TITLE6(14HKY (CYCLES/KM),SPECS)
SPECS(22)=6.1
SPECS(23)=6.0
SPECS(17)=0.1
SPECS(18)=0.1
CALL TITLE6(RULE,11LEVELS (DL),SPECS)
SPECS(23)=4.0
CALL TITLE6(RULE,17VELOCITY (KM/SEC),SPECS)
SPECS(17)=0.1
SPECS(18)=0.1
SPECS(22)=6.3
SPECS(28)=1.0
DO 5 L=1,NL
SPECS(23)=5.7-C.2*FLOAT(L-1)
VALUE=ZLEVEL(L)
CALL DECVAL(RULE,VALUE,SPECS)
5 CONTINUE
SPECS(28)=3.0
DO 6 L=1,NVEL
SPECS(23)=3.8-0.2*FLOAT(L-1)
VALUE=VF(L)
CALL DECVAL (RULE,VALUE,SPECS)
6 CONTINUE
SPECS(22)=7.0
SPECS(23)=1.0
DO 7 L=1,NVEL
SPECS(23)=3.8-0.2*FLOAT(L-1)
VALUE=AZIMU(L)
CALL DECVAL(RULE,VALUE,SPECS)
7 CONTINUE
SPECS(22)=6.3
SPECS(23)=1.5
SPECS(24)=4.0
CALL SIGVAL(RULE,PMAXI,IPOWER,SPECS)
SPECS(22)=6.1
CALL TITLE6(RULE,13PMAX = E,SPECS)
SPECS(22)=7.4
SPECS(28)=0.
VALUE=FLOAT(IPOWER)
CALL DECVAL(RULE,VALUE,SPECS)
RETURN
END

```

```

FKPL.61
FKPL.62
FKPL.63
FKPL.64
FKPL.65
FKPL.66
FKPL.67
FKPL.68
FKPL.69
FKPL.70
FKPL.71
FKPL.72
FKPL.73
FKPL.74
FKPL.75
FKPL.76
FKPL.77
FKPL.78
FKPL.79
FKPL.80
FKPL.81
FKPL.82
FKPL.83
FKPL.84
FKPL.85
FKPL.86
FKPL.87
FKPL.88
FKPL.89
FKPL.90
FKPL.91
FKPL.92
FKPL.93
FKPL.94
FKPL.95
FKPL.96
FKPL.97
FKPL.98
FKPL.99
FKPL.100
FKPL.101
FKPL.102
FKPL.103
FKPL.104
FKPL.105
FKPL.106
FKPL.107
FKPL.108

```

SUBROUTINE GEOPHCN(DATA,PI2,DMPFT,RD,RS,NPTF)

```

GEOPH.2
GEOPH.3
GEOPH.4
GEOPH.5
GEOPH.6
GEOPH.7
GEOPH.8
GEOPH.9
GEOPH.10
GEOPH.11
GEOPH.12
GEOPH.13

```

```

C.....SUBROUTINE GEOPHCN IS USED TO CORRECT THE VELOCITY SENSITIVITY,
C DAMPING RESISTANCE, AND GENERATOR CONSTANCE.
C.....WG = NATURAL FREQUENCY OF THE SEISMOMETER (IN HZ).
C GS = GENERATOR CONSTANCE OF THE SEISMOMETER (IN VOLT/METER/SEC).
C RS = COIL RESISTANCE.
C RD = DAMPING RESISTANCE.
C DMPFT = DAMPING FACTOR.

```

```

COMPLEX DATA,E,PH
COMMON /BLK1/FREQ(226)

```

```

DIMENSION PHASE(257)
DIMENSION DATA(512)
DATA WG/4.57,GS/77.17/
DATA LAB/10MPHASE TEST/
WS=WS*PI2
WS2=WS*WS
RJ2=2.0*DMPFT
RD=RD*1000.0
RS=RS*1000.0
DO 1 I=2,NPTF
WE=PI2*FREQU(I-1)
WE2=XF*WE
A=WE2-WS2
A2=A*A
B=RJ2*WS*WE
B2=B*B
C=SQR(A2+B2)
AMP=WE2/C
PHASE(I)=ATAN2(B,A)
PA=PHASE(I)
E=CMPLX(0.0,PA)
PH=CEXP(E)
PRINT 2,PHASE,PA,PH,AMP,DATA(I)
2 FORMAT (5X,8E12.3)
DATA(I)=DATA(I)/(AMP*PH)
IF (RD.GE.999000.0) GO TO 1
K=RS*RD
DATA(I)=DATA(I)*F/RD
1 CONTINUE
RETURN
END

```

GEOPH.14  
GEOPH.15  
GEOPH.16  
GEOPH.17  
GEOPH.18  
GEOPH.19  
GEOPH.20  
GEOPH.21  
GEOPH.22  
GEOPH.23  
GEOPH.24  
GEOPH.25  
GEOPH.26  
GEOPH.27  
GEOPH.28  
GEOPH.29  
GEOPH.30  
GEOPH.31  
GEOPH.32  
GEOPH.33  
GEOPH.34  
GEOPH.35  
GEOPH.36  
GEOPH.37  
GEOPH.38  
GEOPH.39  
GEOPH.40  
GEOPH.41  
GEOPH.42  
GEOPH.43  
GEOPH.44

```

SUBROUTINE HAMMN(NPTT,PI2,W,ENW)
DIMENSION W(NPTT)
ALFA=0.54
ALFA1=1.0-ALFA
A=FLOAT((NPTT-1)/2)
ENW=0.0
DO 5 M=1,NPTT
B=FLOAT(M-1)
ANG=PI2*(B-A)/FLOAT(NPTT)
W(M)=ALFA+ALFA1*CCS(ANG)
ENW=ENW+W(M)*W(M)
5 CONTINUE
RETURN
END

```

HAMMN.2  
HAMMN.3  
HAMMN.4  
HAMMN.5  
HAMMN.6  
HAMMN.7  
HAMMN.8  
HAMMN.9  
HAMMN.10  
HAMMN.11  
HAMMN.12  
HAMMN.13  
HAMMN.14  
HAMMN.15

```

SUBROUTINE PLOTCH(LABEL1,L,NPTF)

```

C.....THIS SUBROUTINE PLOTS THE COHERENCE BETWEEN TWO SEISMOMETER  
C OUTPUTS.

```

COMMON /BLK1/X(226)
COMMON /BLK2/Y(226)
COMMON /BLK3/ERRBU(164),ERRBC(164)
DIMENSION SPECS(30),LINE(4)
SPECS(1)=1.0
SPECS(2)=1.5
SPECS(3)=25.0
SPECS(4)=0.3

```

PLCOH.2  
PLCOH.3  
PLCOH.4  
PLCOH.5  
PLCOH.6  
PLCOH.7  
PLCOH.8  
PLCOH.9  
PLCOH.10  
PLCOH.11  
PLCOH.12  
PLCOH.13  
PLCOH.14

```
SPECS(5)=1.0
SPECS(6)=0.0
5 SPECS(7)=7.675
SPECS(8)=5.2
SPECS(10)=10.0
SPECS(11)=1.0
SPECS(12)=99.
CALL AXLGLI(SPECS)
SPECS(25)=-0.15
CALL SAXLGT(SPECS)
SPECS(27)=-0.15
CALL SAXLIR(SPECS)
SPECS(13)=FLOAT(NPTF)
SPECS(14)=1.0
SPECS(15)=1.0
SPECS(30)=97.
CALL SLLGLI(X,Y,SPECS)
DO 10 I=1,NPTF
Y(I)=ERRORD(I)
10 CONTINUE
SPECS(16)=17.0
SPECS(17)=0.05
SPECS(18)=0.05
CALL PSLGLI(X,Y,SPECS)
DO 11 I=1,NPTF
Y(I)=ERRORD(I)
11 CONTINUE
CALL PSLGLI(X,Y,SPECS)
15 SPECS(17)=.1
SPECS(18)=.1
SPECS(19)=0.0
SPECS(21)=1.0
SPECS(24)=0.0
SPECS(29)=2.0
CALL NGLEGB(SPECS)
SPECS(20)=0.0
SPECS(26)=0.0
SPECS(23)=1.0
CALL NGDLTL(SPECS)
SPECS(17)=0.15
SPECS(18)=0.15
SPECS(24)=0.2
CALL TITLEG(14+FREQUENCY (F2),SPECS)
SPECS(20)=90.0
SPECS(26)=0.4
CALL TITLEL(9HCOHERENCE,SPECS)
SPECS(17)=0.15
SPECS(18)=0.15
SPECS(19)=0.0
SPECS(20)=0.0
SPECS(21)=1.0
SPECS(22)=2.0
SPECS(23)=6.0
RULE=1.0
LINE(1)=LABEL1
LINE(2)=0
CALL TITLEG(RULE,LINE,SPECS)
SPECS(22)=4.5
SPECS(28)=0.
VALUE=FLOAT(L)
CALL DECVAL(RULE,VALUE,SPECS)
RETURN
```

PLCOH.15  
PLCOH.16  
PLCOH.17  
PLCOH.18  
PLCOH.19  
PLCOH.20  
PLCOH.21  
PLCOH.22  
PLCOH.23  
PLCOH.24  
PLCOH.25  
PLCOH.26  
PLCOH.27  
PLCOH.28  
PLCOH.29  
PLCOH.30  
PLCOH.31  
PLCOH.32  
PLCOH.33  
PLCOH.34  
PLCOH.35  
PLCOH.36  
PLCOH.37  
PLCOH.38  
PLCOH.39  
PLCOH.40  
PLCOH.41  
PLCOH.42  
PLCOH.43  
PLCOH.44  
PLCOH.45  
PLCOH.46  
PLCOH.47  
PLCOH.48  
PLCOH.49  
PLCOH.50  
PLCOH.51  
PLCOH.52  
PLCOH.53  
PLCOH.54  
PLCOH.55  
PLCOH.56  
PLCOH.57  
PLCOH.58  
PLCOH.59  
PLCOH.60  
PLCOH.61  
PLCOH.62  
PLCOH.63  
PLCOH.64  
PLCOH.65  
PLCOH.66  
PLCOH.67  
PLCOH.68  
PLCOH.69  
PLCOH.70  
PLCOH.71  
PLCOH.72  
PLCOH.73  
PLCOH.74  
PLCOH.75  
PLCOH.76

END

PLCOH.77  
PLCOH.78

SUBROUTINE PLUTVS

C.....THIS SUBROUTINE PLOTS THE VELOCITY SPECTRAL DENSITY IN LOG\*LOG  
C.....SCALE.

DIMENSION SPECS(30),LINE(4)  
COMMON /BLK1/X(226)  
COMMON /BLK11/STAT(2)  
COMMON /BLK14/Y(226),NREC,NPTF  
SPECS(1)=1.0  
SPECS(2)=1.5  
SPECS(3)=25.0  
SPECS(4)=0.3  
DO 8 I=1,NPTF  
IF (Y(I).GT.1.0E+03) GO TO 6  
IF (Y(I).LT.1.0E-01) GO TO 5  
8 CONTINUE  
SPECS(5)=1.0E+03  
SPECS(6)=1.0E-01  
GO TO 5  
6 SPECS(5)=1.0E+04  
SPECS(6)=1.0  
GO TO 5  
9 SPECS(5)=1.0E+01  
SPECS(6)=1.0E-03  
5 SPECS(7)=7.075  
SPECS(8)=5.2  
SPECS(11)=1.0  
SPECS(12)=9.9  
CALL AXLGLG(SPECS)  
YMAX=SPECS(5)  
YMIN=SPECS(6)  
DO 10 I=1,NPTF  
IF (Y(I).GT.YMAX) Y(I)=YMAX  
IF (Y(I).LT.YMIN) Y(I)=YMIN  
10 CONTINUE  
SPECS(25)=-0.15  
CALL SAXLGT(SPECS)  
SPECS(27)=-0.15  
CALL SAXLGR(SPECS)  
SPECS(13)=FLOAT(NPTF)  
SPECS(14)=1.0  
SPECS(15)=1.0  
SPECS(30)=97.  
CALL SLLGLG(X,Y,SPECS)  
15 SPECS(17)=.1  
SPECS(18)=.1  
SPECS(19)=10.0  
SPECS(19)=0.0  
SPECS(21)=1.0  
SPECS(24)=0.0  
SPECS(29)=2.0  
CALL NOLGB(SPECS)  
SPECS(20)=0.0  
SPECS(26)=0.0  
CALL NOLGL(SPECS)  
SPECS(17)=0.15  
SPECS(18)=0.15

PLVS.2  
PLVS.3  
PLVS.4  
PLVS.5  
PLVS.6  
PLVS.7  
PLVS.8  
PLVS.9  
PLVS.10  
PLVS.11  
PLVS.12  
PLVS.13  
PLVS.14  
PLVS.15  
PLVS.16  
PLVS.17  
PLVS.18  
PLVS.19  
PLVS.20  
PLVS.21  
PLVS.22  
PLVS.23  
PLVS.24  
PLVS.25  
PLVS.26  
PLVS.27  
PLVS.28  
PLVS.29  
PLVS.30  
PLVS.31  
PLVS.32  
PLVS.33  
PLVS.34  
PLVS.35  
PLVS.36  
PLVS.37  
PLVS.38  
PLVS.39  
PLVS.40  
PLVS.41  
PLVS.42  
PLVS.43  
PLVS.44  
PLVS.45  
PLVS.46  
PLVS.47  
PLVS.48  
PLVS.49  
PLVS.50  
PLVS.51  
PLVS.52  
PLVS.53  
PLVS.54  
PLVS.55  
PLVS.56  
PLVS.57  
PLVS.58  
PLVS.59

```

SPECS(24)=0.2
CALL TITLEB(14HFREQUENCY (HZ),SPECS)
SPECS(20)=90.0
SPECS(26)=0.4
CALL TITLEC(30HVSD (MILLIMICRON/SEC/SQRT(HZ)),SPECS)
SPECS(17)=0.15
SPECS(15)=0.15
SPECS(19)=0.0
SPECS(20)=0.0
SPECS(21)=1.0
SPECS(23)=0.8
RULE=1.0
SPECS(22)=2.15
SPECS(28)=0.0
VALUE=NREC
CALL DFCVAL(RULE,VALUE,SPECS)
LINE(1)=STAT(1)
LINE(2)=10H RECRDS
LINE(3)=0
SPECS(22)=0.5
CALL TITLEG(RULE,LINE,SPECS)
RETURN
END

```

PLVS.60  
PLVS.61  
PLVS.62  
PLVS.63  
PLVS.64  
PLVS.65  
PLVS.66  
PLVS.67  
PLVS.68  
PLVS.69  
PLVS.70  
PLVS.71  
PLVS.72  
PLVS.73  
PLVS.74  
PLVS.75  
PLVS.76  
PLVS.77  
PLVS.78  
PLVS.79  
PLVS.80  
PLVS.81  
PLVS.82

```

SUBROUTINE SYMCOB(P,KX,KY)

```

SYMCO.2

```

C.....THIS SUBROUTINE PLOTS THE FKPSD BY SYMBOLS AT A FREQUENCY
C COMPONENT.

```

SYMCO.3  
SYMCO.4  
SYMCO.5  
SYMCO.6  
SYMCO.7  
SYMCO.8  
SYMCO.9  
SYMCO.10  
SYMCO.11  
SYMCO.12  
SYMCO.13  
SYMCO.14  
SYMCO.15  
SYMCO.16  
SYMCO.17  
SYMCO.18  
SYMCO.19  
SYMCO.20  
SYMCO.21  
SYMCO.22  
SYMCO.23  
SYMCO.24  
SYMCO.25  
SYMCO.26  
SYMCO.27  
SYMCO.28  
SYMCO.29  
SYMCO.30  
SYMCO.31  
SYMCO.32  
SYMCO.33  
SYMCO.34  
SYMCO.35  
SYMCO.36  
SYMCO.37  
SYMCO.38

```

COMMON /BLK10/PLEVEL(10),NPLEV
DIMENSION P(41,41),KEY(41)
DATA K1,K2,K3,K4,K5,K6/1H0,1H*,1H/,1H*,1H*,1H*/
DATA K7/1H /
PRINT 750
750 FORMAT (1H0)
PRINT 760, (PLEVEL(I),I=1,NPLEV)
760 FORMAT (1H1,* CONTINUF LEVELS = *,7E12.3)
PRINT 762
762 FORMAT (1HC,//////////)
PRINT 700
700 FORMAT (64X,1FD)
PRINT 761
761 FORMAT (2X,123# . . . . .)
1 . . . . .)
DO 100 JKY=1,NKY
JY=NKY-JKY+1
DO 200 KX=1,NKX
IF (P(KX,JY).GE.PLEVEL(1)) GO TO 10
IF (P(KX,JY).GE.PLEVEL(2)) GO TO 20
IF (P(KX,JY).GE.PLEVEL(3)) GO TO 30
IF (P(KX,JY).GE.PLEVEL(4)) GO TO 40
IF (P(KX,JY).GE.PLEVEL(5)) GO TO 50
IF (P(KX,JY).GE.PLEVEL(6)) GO TO 60
KEY(KX)=K7
GO TO 200
60 KEY(KX)=K6
GO TO 200
50 KEY(KX)=K5
GO TO 200
40 KEY(KX)=K4

```

```

GO TO 200
30 KEY(KX)=K3
GO TO 200
20 KEY(KX)=K2
GO TO 200
10 KEY(KX)=K1
200 CONTINUE
IF (JY.EQ.21) GO TO 151
PRINT 150,(KEY(J),J=1,NKX)
150 FORMAT (1H0,1X,4I(2X,A1))
GO TO 100
151 PRINT 152,(KEY(J),J=1,NKX)
152 FORMAT (1H0,*0*,4I(2X,A1),* 0*)
100 CONTINUE
PRINT 761
PRINT 700
RETURN
END

```

SYMCO.39  
SYMCO.40  
SYMCO.41  
SYMCO.42  
SYMCO.43  
SYMCO.44  
SYMCO.45  
SYMCO.46  
SYMCO.47  
SYMCO.48  
SYMCO.49  
SYMCO.50  
SYMCO.51  
SYMCO.52  
SYMCO.53  
SYMCO.54  
SYMCO.55  
SYMCO.56

```

SUBROUTINE YPLOT(Y,IB,IE,IK,PMAX,PMIN,TITLE,L,X)
DIMENSION X(1)
DIMENSION Y(1),KMAP(100),ISCALE(11)
DATA KBLANK,KY/1F,1H*/
DATA ISCALE/ 0,10,20,30,40,50,60,70,80,90,100/
DO 2 I=1,100
2 KMAP(I)=KBLANK
C....PRINT BEGINNING OF GRAPH
C=100.0/(PMAX-PMIN)
B=C*PMIN
SCALE=1.0/C
PRINT 5,TITLE,L,IB,IE,IK,PMAX,PMIN
5 FORMAT (1H1,10X,*YPLOT.....*,A10,15,/,10X,*IB=*,15,2X,*IE=*,15,2XYPLOT.15
1,*IK=*,15,2X,*PMAX=*,E12.3,2X,*PMIN=*,E12.3,///)
YPLOT.16
PRINT 6, (ISCALE(I),I=1,11)
YPLOT.17
6 FORMAT(11X,5(I2,8X),1X,12,5(7X,13))
YPLOT.18
DO 7 I=1,2
YPLOT.19
7 PRINT 8
YPLOT.20
8 FORMAT(13X,(11(1H, 9X)))
YPLOT.21
PRINT 12
YPLOT.22
12 FORMAT ( 6X,1HX,7X,10CH.....,7X,1HYPLOT.24
1.....,7X,1HY
YPLOT.25
C....PRINT BODY OF GRAPH
YPLOT.26
DO 10 I=IB,IE,IK
YPLOT.27
IY=C*Y(I)-B
YPLOT.28
C CHECK THAT Y IS ON SCALE
YPLOT.29
IF(IY.LT. 1 .OR. IY.GT. 100) GO TO 13
YPLOT.30
KMAP(IY)=KY
YPLOT.31
PRINT 9, X(I),(KMAP(J),J=1,100),Y(I),I
YPLOT.32
9 FORMAT ( 2X,F9.3,1X,1H1,100A1,1F1,2X,E10.3,1X,14)
YPLOT.33
GO TO 14
YPLOT.34
13 PRINT 9, X(I),(KMAP(J),J=1,100),Y(I)
YPLOT.35
GO TO 10
YPLOT.36
14 KMAP(IY)=KBLANK
YPLOT.37
10 CONTINUE
YPLOT.38
RETURN
YPLOT.39
END
YPLOT.40

```

62.50 6 2500 24 2  
5 -1.0 -3.0 -6.0 -9.0-12.0

